

Astroparticle physics - New messengers of the Universe

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Université Paris Cité
Laboratoire APC

European biased perspective



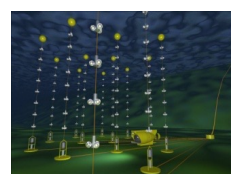
©APPEC

Outline

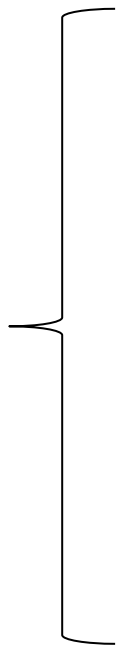


Motivations for a Multi-messenger approach

Success of multi-wavelength
Onset of Astroparticle Physics
Key scientific questions



Historical aspects
Detection principles
Achievements
Future challenges



Cosmic-rays

VHE gamma-rays

Neutrinos

Gravitational waves

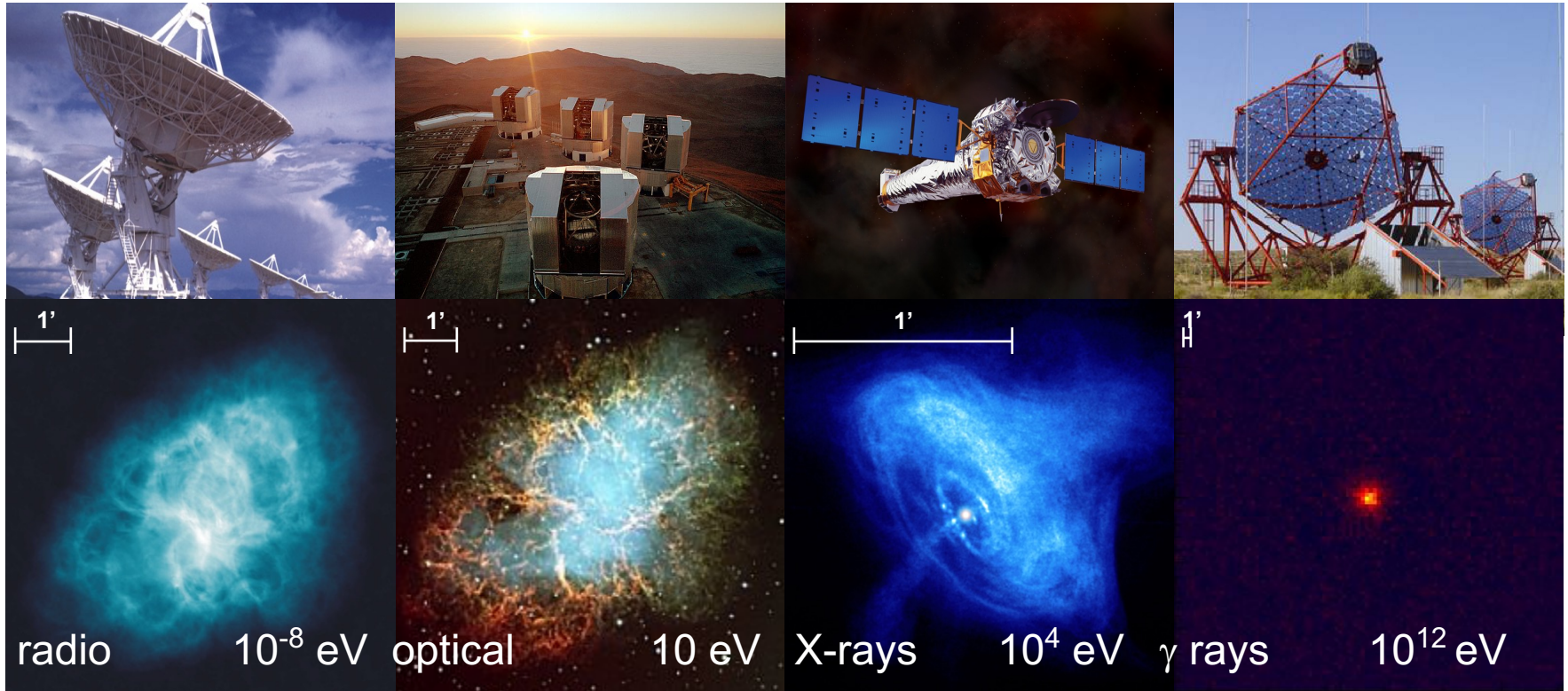
→ Livia Conti

Concluding remark

Multi-wavelength Astronomy

The Crab Nebula

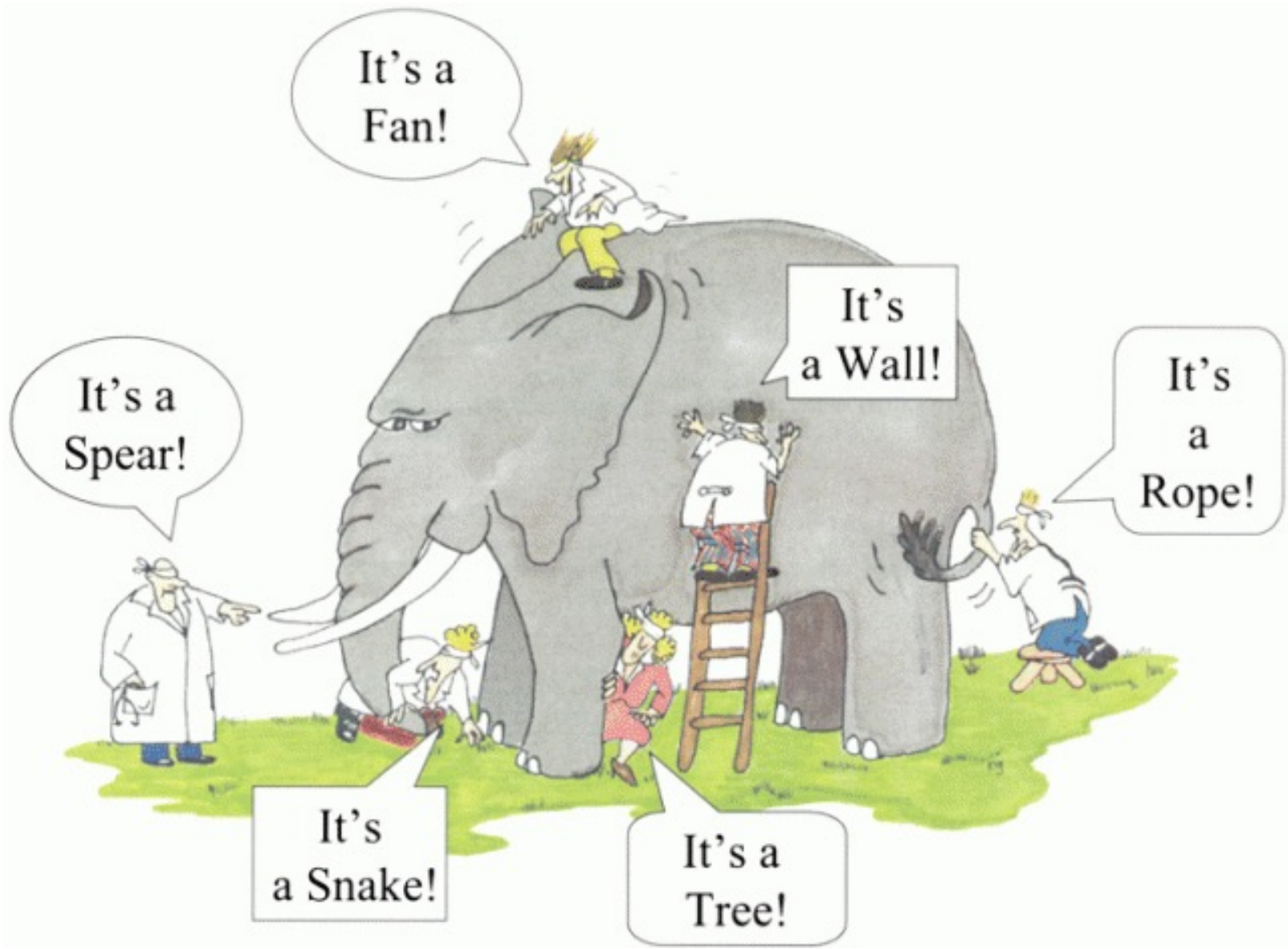
(first source seen in the TeV gamma domain)



Thermal emission, dusts, molecular clouds, non-thermal processes...

Multiwavelength studies enable to get a more complete modelling of the source

Why several messengers ?



What is Astroparticle Physics?

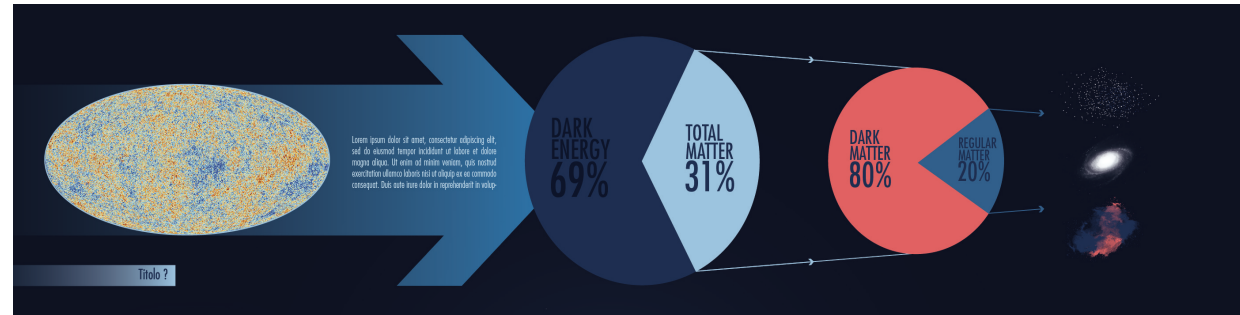
Astroparticle physics, also called **particle astrophysics**, is a branch of **particle physics** that studies elementary particles of astronomical origin and their relation to **astrophysics** and **cosmology**. It is a relatively new field of research emerging at the intersection of particle physics, **astronomy**, astrophysics, **detector physics**, **relativity**, **solid state physics**, and cosmology. Partly motivated by the discovery of **neutrino oscillation**, the field has undergone rapid development, both theoretically and experimentally, since the early 2000s.^[1]

- 1) Associate physics at different scales to explain the phenomena
 - E.g. Nuclear Physics and Gravity to understand the equilibrium of the stars
- 2) Use multi-messenger probes adapted to the corresponding space-time scales and study cosmological events in different times and depths of interaction
 - Neutrinos and Gravitational waves to probe “deep and early processes” vs electromagnetic interactions coming at later stages
 - Establish a global, low latency network to share fast enough the incoming signals
- 3) Discover new physics by comparing multi-messenger representations /cartographies of the Universe
 - E.g. how different can be an “early” vs a “late” cartography of the Universe (Hubble values tension) or black hole populations vs this of “living” stars

Key questions

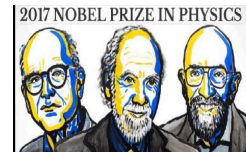
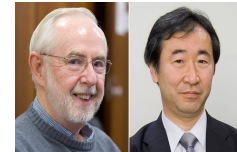
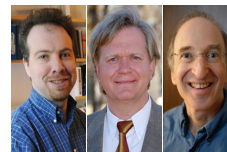
- The Primordial Universe
 - Inflation

- The Dark Universe
 - Dark energy
 - Dark matter
 - Matter/antimatter

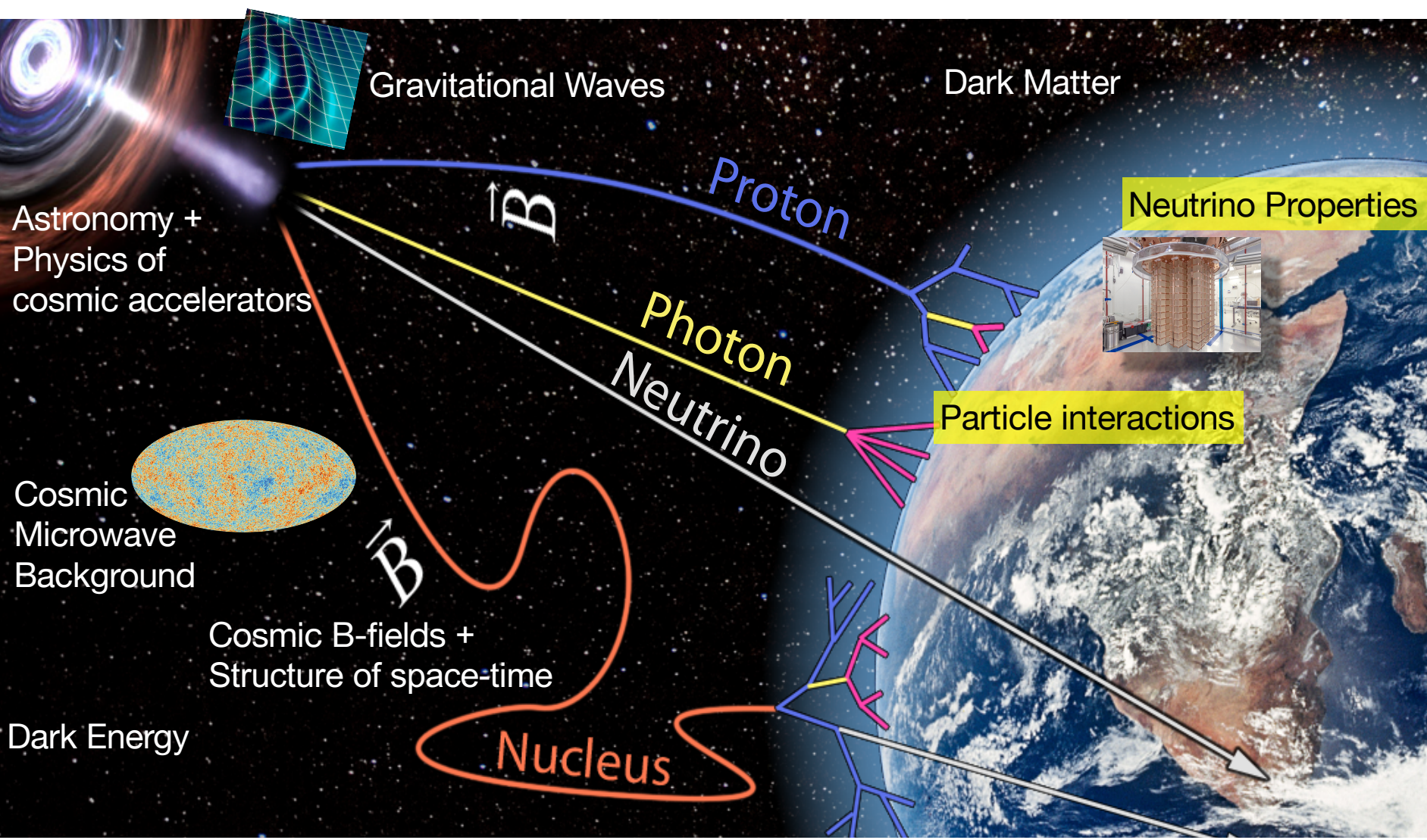


- The Violent Universe
 - Nature of black holes, neutron stars and white dwarfs
 - Formation and evolution of galaxies
 - Violent phenomena
 - Physics of dense matter and strong EM fields

A Nobel Prize field !



Multi-Messenger Astroparticle Physics



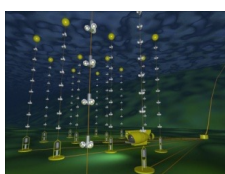
A very large Scientific Scope

Outline

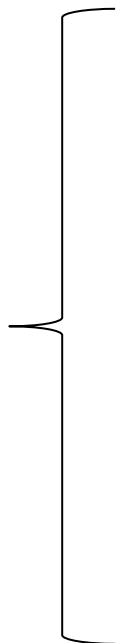


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VHE gamma-rays

Neutrinos

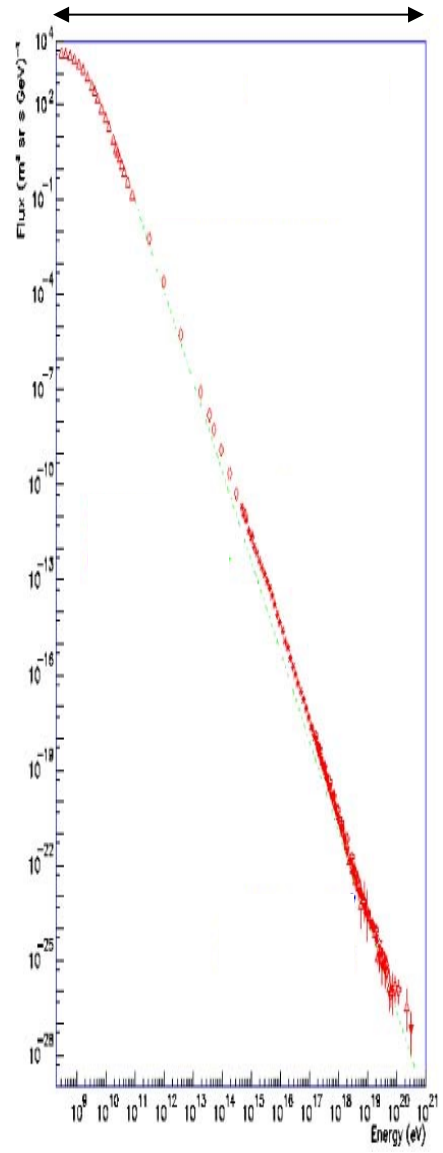
Gravitational waves

→ Livia Conti

Concluding remark

Cosmic rays, challenging since 1900

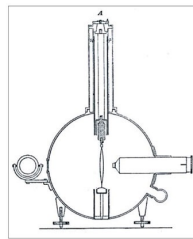
> 12 orders of magnitude!



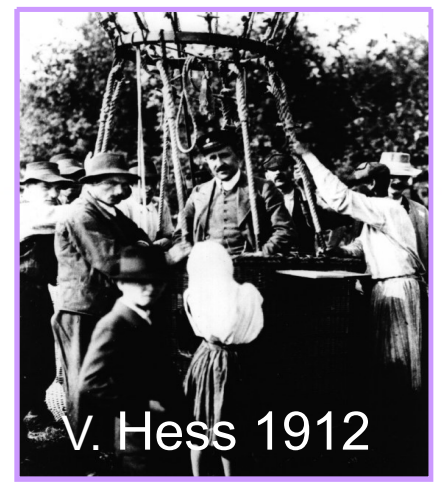
4 per cm² per second

32 orders of magnitude (/m²/sr/GeV)

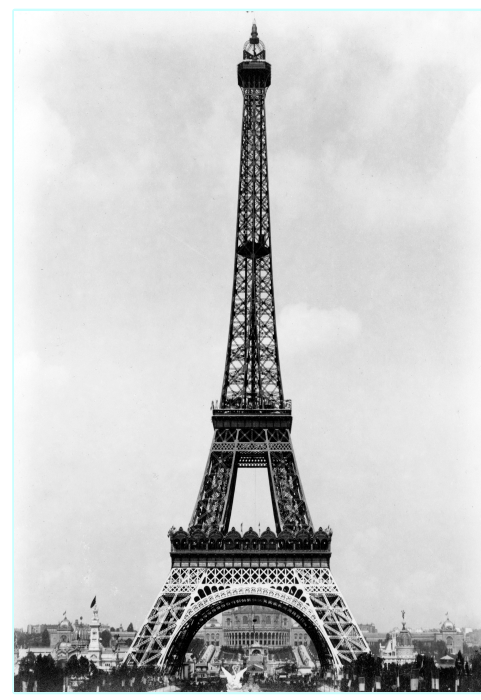
1 per m² per billion years



T. Wulf (1909)



V. Hess 1912



- 1932 Positron
- 1936 Muon
- 1947 Pions : π^0, π^+, π^-
- 1949 Kaons (K)
- 1949 Lambda (Λ)
- 1952 Cascade (Ξ)
- 1953 Sigma (Σ)

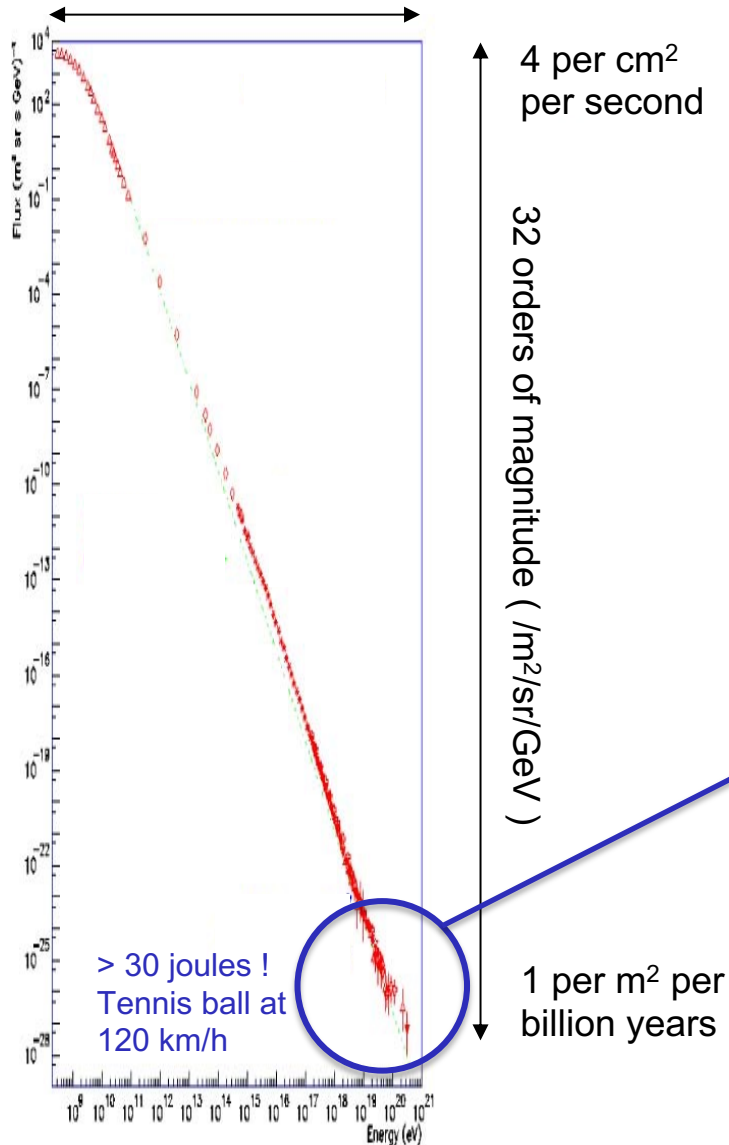


1953 towards the accelerators

CONGRÈS INTERNATIONAL SUR LE RAYONNEMENT COSMIQUE
BAD NAEUM-DE-BAD NAEUM, 6-12 Juillet 1953
Paris AIX

Cosmic rays, challenging since 1900

> 12 orders of magnitude!



✧ Major role in Galactic ecosystem !

- ✧ Energy density \sim star light, thermal, B field
- ✧ Regulate the equilibrium between the different phases of the interstellar medium
- ✧ Control ionisation, heating
- ✧ Regulate star formation
- ✧ Produce Li, Be and B!

✧ Major unknowns

- ✧ Sources are unknown (Galactic and Extragal.)
- ✧ Acceleration processes are uncertain

✧ How does Nature proceed ? to produce them

- ✧ What, where and how ?

✧ How does Nature behave ? at such energies

- ✧ Lorentz factors beyond all tests of Relativity
- ✧ Cross section beyond LHC reach

NB: GZK effect = interaction of the UHECRs with the ambient photons

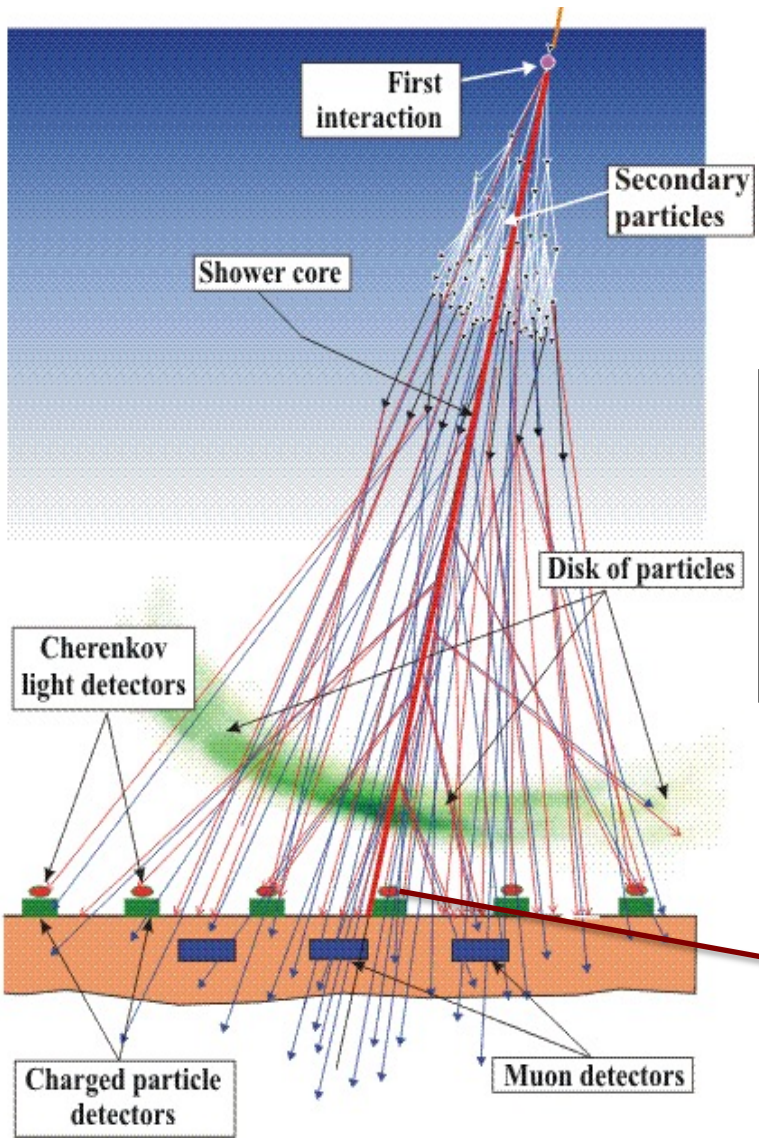
Indirect detection The atmosphere is the primary cosmic ray detector!

Detection of the shower particles

(sampling of the "shower front")



Volcano Ranch (1959–1963)
February, 22nd 1962: 10^{20} eV !



Detection of the induced effect of shower particles

(fluorescence light)



Fly's Eye (1981–1995)
October, 15th 1993: 3×10^{20} eV !



Current generation ground detectors

✧ Hi-Res (1993–1997–2006–2010)

✧ The Telescope array

700 km² (Utah, USA)



✧ The Pierre Auger Observatory
3000 km² (Argentina)



- GZK-like attenuation: established!
- Composition getting heavier above a few EeV
- Departure from isotropy (first order: dipole) at "low" energies (≥ 8 EeV, 6%, 6σ)
- Correlation with matter (but not discriminating) at intermediate energies ($> 3\sigma$)
- Warm spots at intermediate angular scales at the highest energies
- Shower physics: "muon excess" (indirect)
- Declination-dependent energy spectrum (4.3σ)

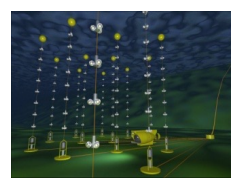
However, no clear progress regarding sources and acceleration mechanisms

Outline

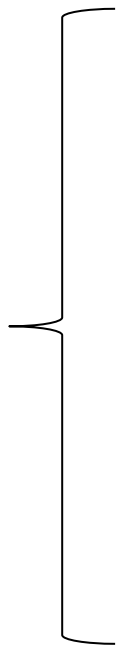


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High-Energy Gamma Rays (MeV-TeV)

The last spectral domain in photonic astrophysics

- Covers large energy range with different observatories
- Satellites (Fermi, AMEGO (launch 2029), ASTROGAM)
- Imaging Air Cherenkov Telescopes (H.E.S.S., Veritas, MAGIC)
- Ground-based arrays (GRAPES, TAIGA, HAWC, LHAASO, SWGO)
- Main future project within APPEC: [CTA](#) (ESFRI)



FERMI

Veritas



MAGIC



H.E.S.S.



LHAASO

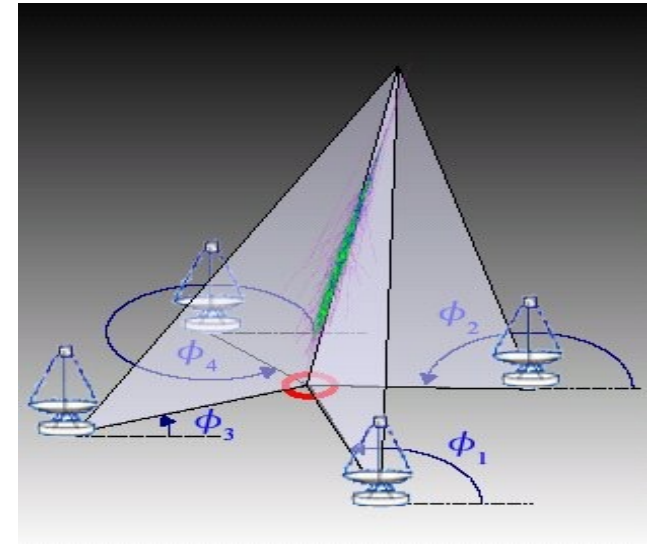


HAWC



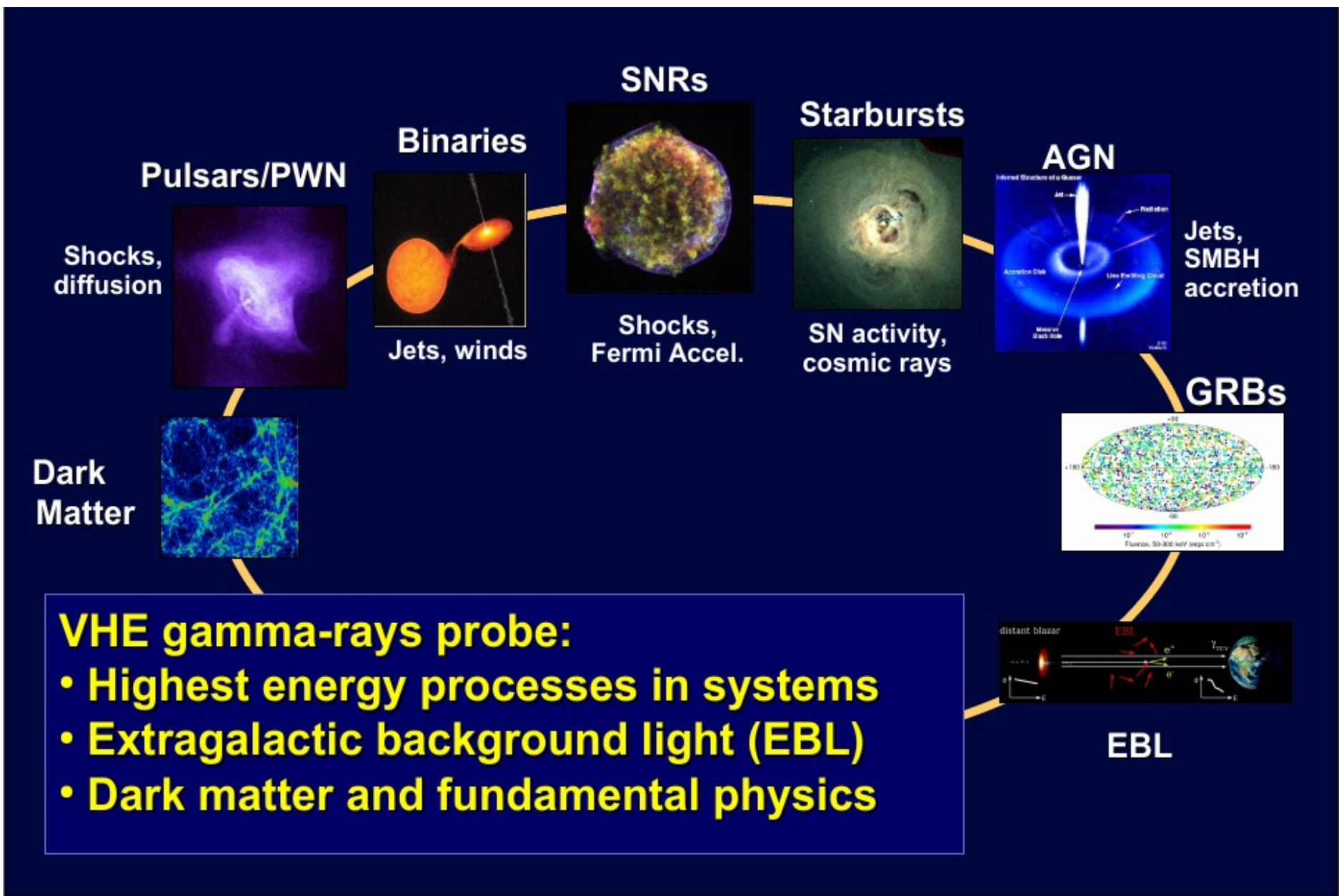
Imaging Atmospheric Cherenkov Telescopes

- Cherenkov light is emitted on a very narrow cone ($\theta < 1^\circ$) illuminating an area of about 300 m diameter at 1800 m a.s.l. on the ground.
- A telescope located within the light pool detects the shower if it collects enough Cherenkov photons \rightarrow effective detection area $\sim 10^5 \text{ m}^2$
- With an array of several telescopes, the shower can be reconstructed in 3D (stereoscopy)
 - \rightarrow total number of photons (energy estimator)
 - \rightarrow better angular resolution

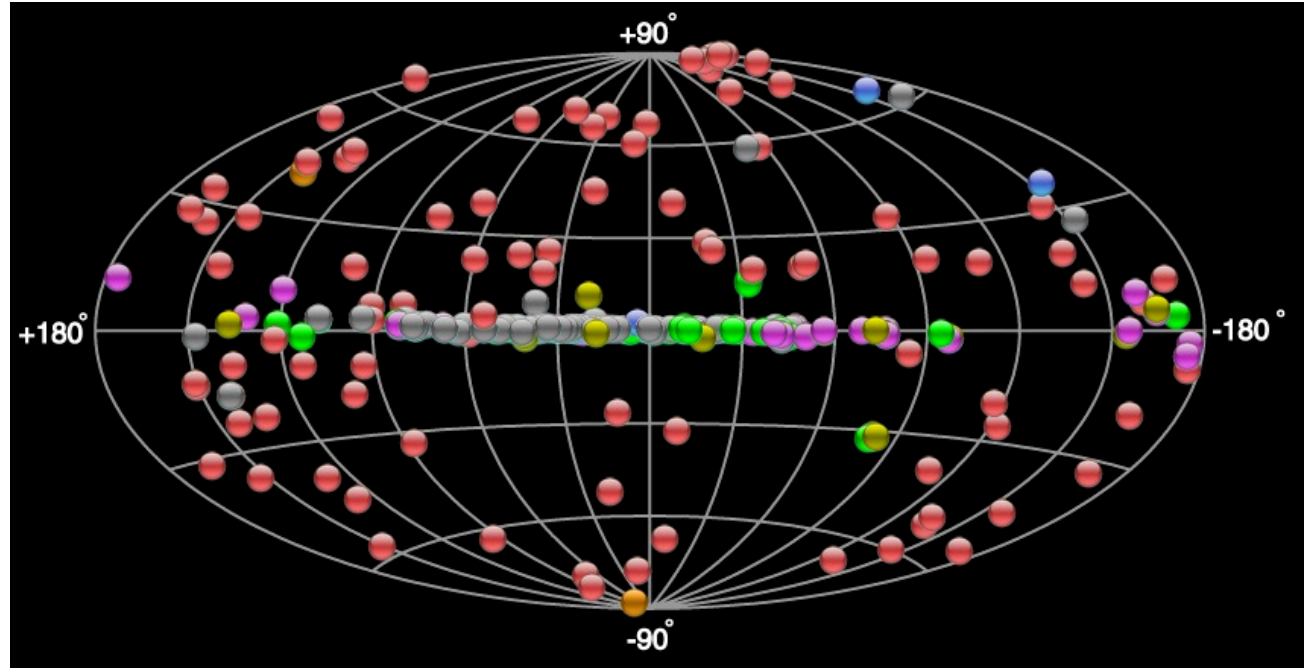


Mor than 30 years ago, the Crab nebula was the first γ -ray source firmly detected (9σ) at very high energies by The Whipple Observatory

Astronomy with IACTs



The VHE gamma sky



- ### Source Types
- PWN TeV Halo
PWN/TeV Halo TeV Halo Candidate
 - XRB Nova Gamma BIN
Binary PSR
 - HBL IBL GRB FSRQ LBL
AGN (unknown type) FRI
Blazar
 - Shell Giant Molecular
Cloud SNR/Molec. Cloud
Composite SNR
Superbubble SNR
 - Starburst
 - DARK UNID Other
 - Star Forming Region
Globular Cluster Massive
Star Cluster BIN
uQuasar Cat. Var. BL
Lac (class unclear) WR

Black Holes, Jets, and the History of Star-Formation

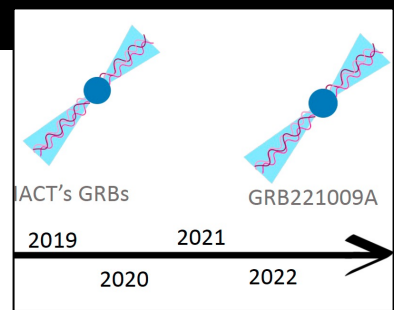
How do black holes make jets and accelerate particles?
Spectrum and redshift distribution of extragal. background light ?

Cosmic Rays

How and where are particles accelerated ?
What is the connection between star formation and cosmic rays ?

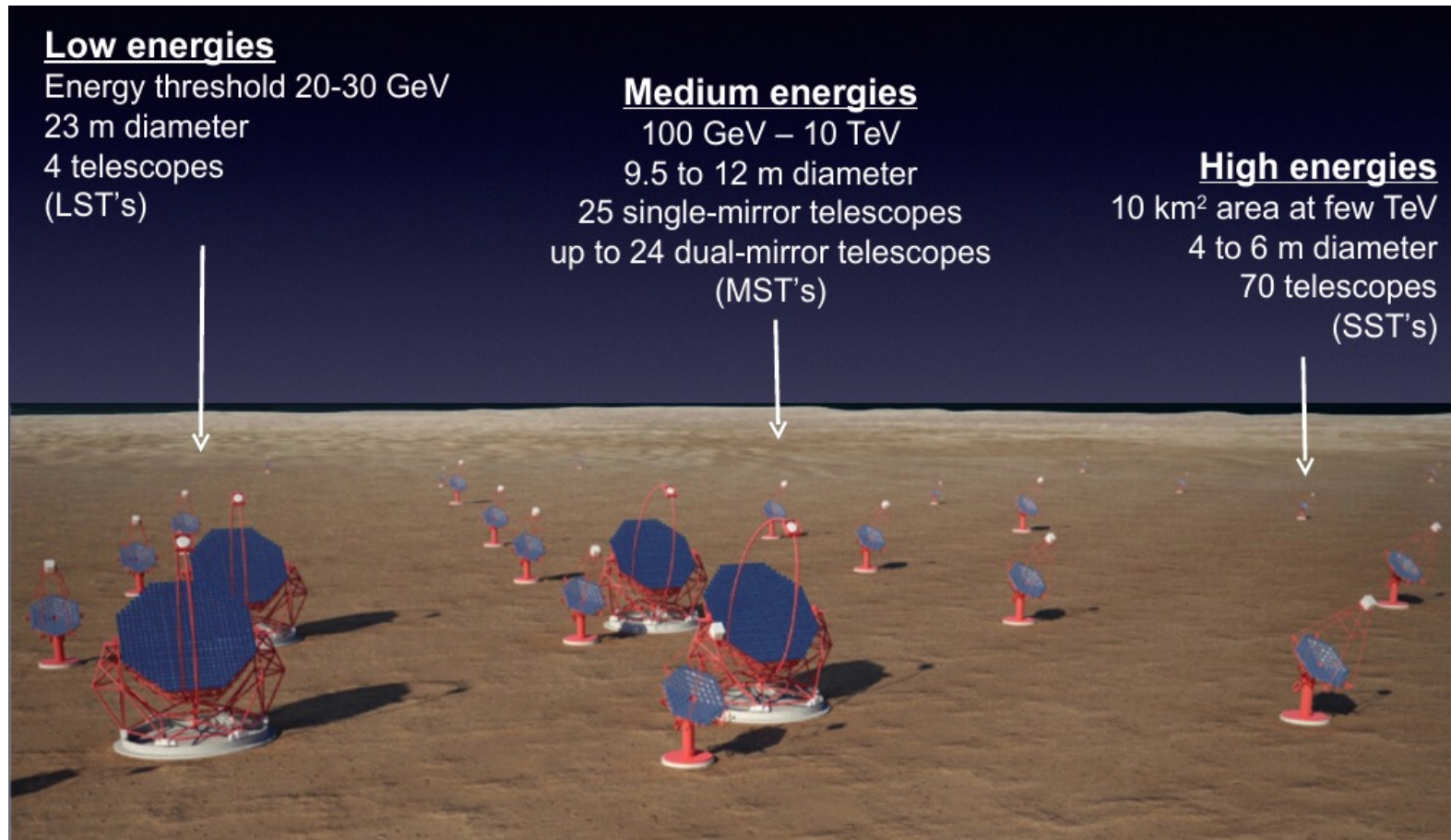
Dark Matter and Lorentz Invariance

What is the nature of Dark Matter ? How is it distributed ?
Is the speed of light a constant for high energy photons ?



Next: Cherenkov Telescope Array

- ESFRI Project
- Open, proposal-driven observatory
- 3 telescope types: LST, MST, SST
- 2 sites: La Palma + Chile
- Governance: ERIC (established 2022)
- 31 countries, >200 institutes, ~1400 scientists
- Construction next 3-5 years
- 10 x more sensitive than precursors

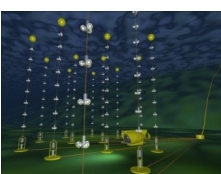


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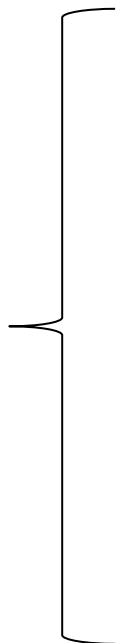


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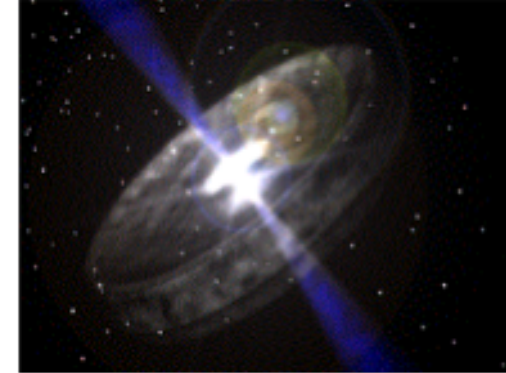
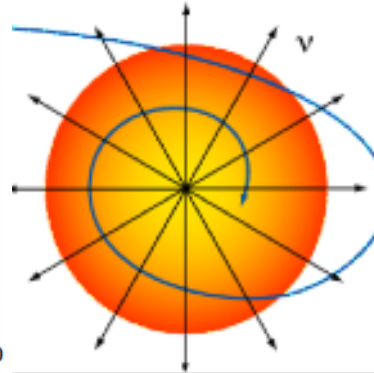
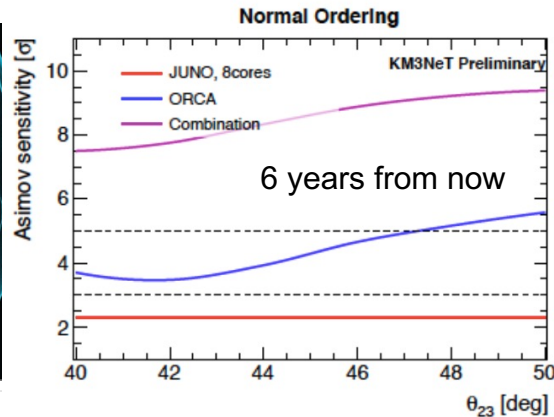
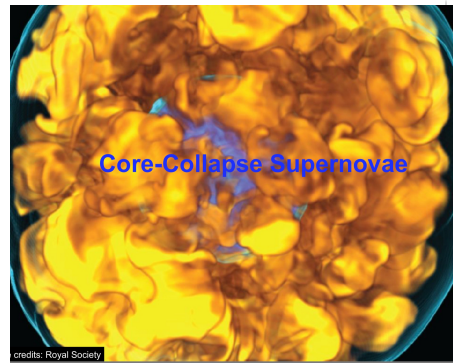


Cosmic-rays
VHE gamma-rays
Neutrinos
Gravitational waves
→ Livia Conti

Concluding remark

Neutrino Telescopes

The Science scope



MeV Energy
No reco. in HE NT

Low Energy
 $\text{GeV} < E < 50 \text{ GeV}$

Medium Energy
 $10 \text{ GeV} < E < 1 \text{ TeV}$

High Energy
 $E > 1 \text{ TeV}$

CCSNe

Oscillation

Dark Matter

HE Astrophysics

Full Galactic coverage
All mass progenitors
Triangulations

PMNS Unitarity KM3NeT & IC
Neutrino Mass Ordering
with KM3NeT (ORCA $\geq 3\sigma$ 3yrs)

Not covered
here

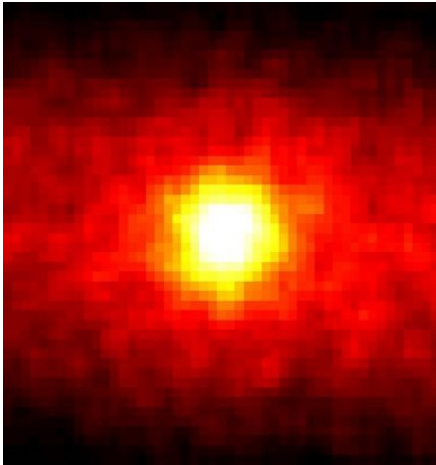
Focus of this talk

+ Exotics (Monopoles, Nuclearites, etc.)

+ Environmental Sciences

First extraterrestrial neutrinos – Multi-messengers

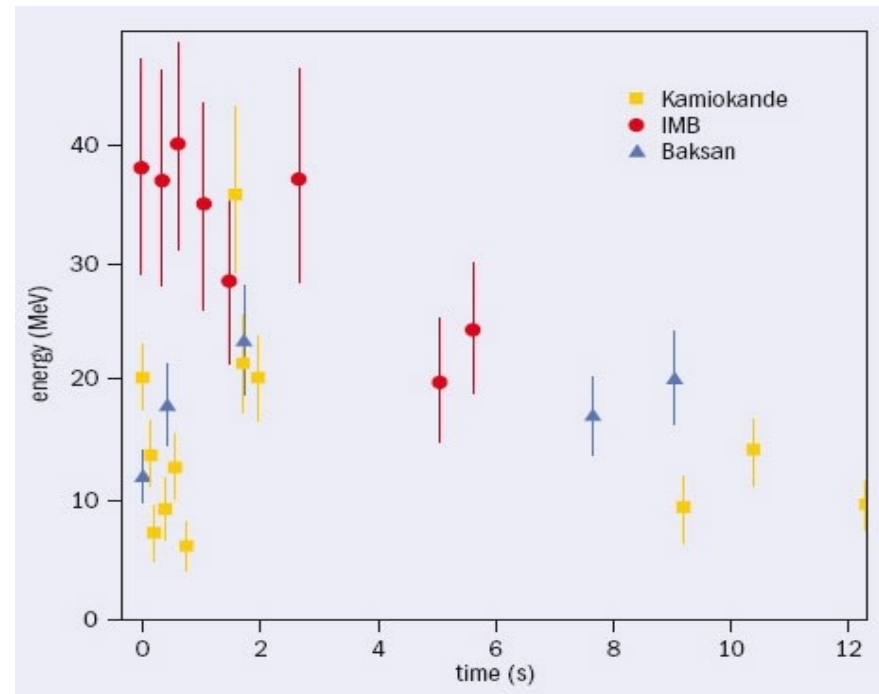
- ❖ 1960's: SUN seen by Homestake
- ❖ 1988 : Kamiokande



→ Confirmation of deficit of ν_e already observed in radiochemical experiments

Neutrino Oscillate

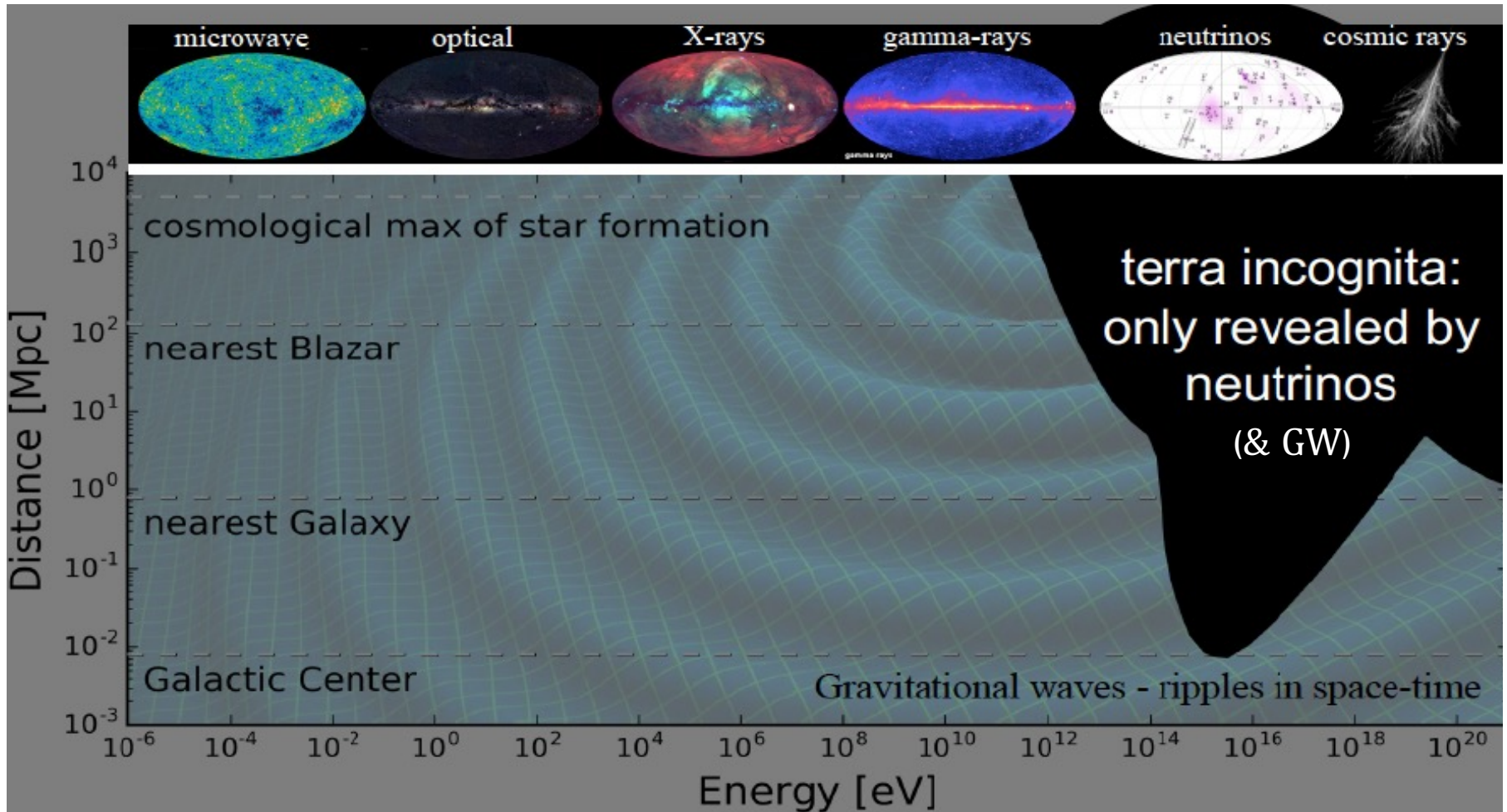
- ❖ 1987: Observation of a neutrino burst from the supernova SN1987A in the Large Magellanic Cloud



24 neutrinos detected in ~10 seconds about 3 hours before the electromagnetic emission

Typical energy ~10 MeV

Neutrinos as cosmic messengers



From F. Halzen

Neutrinos

✓ Transient sources

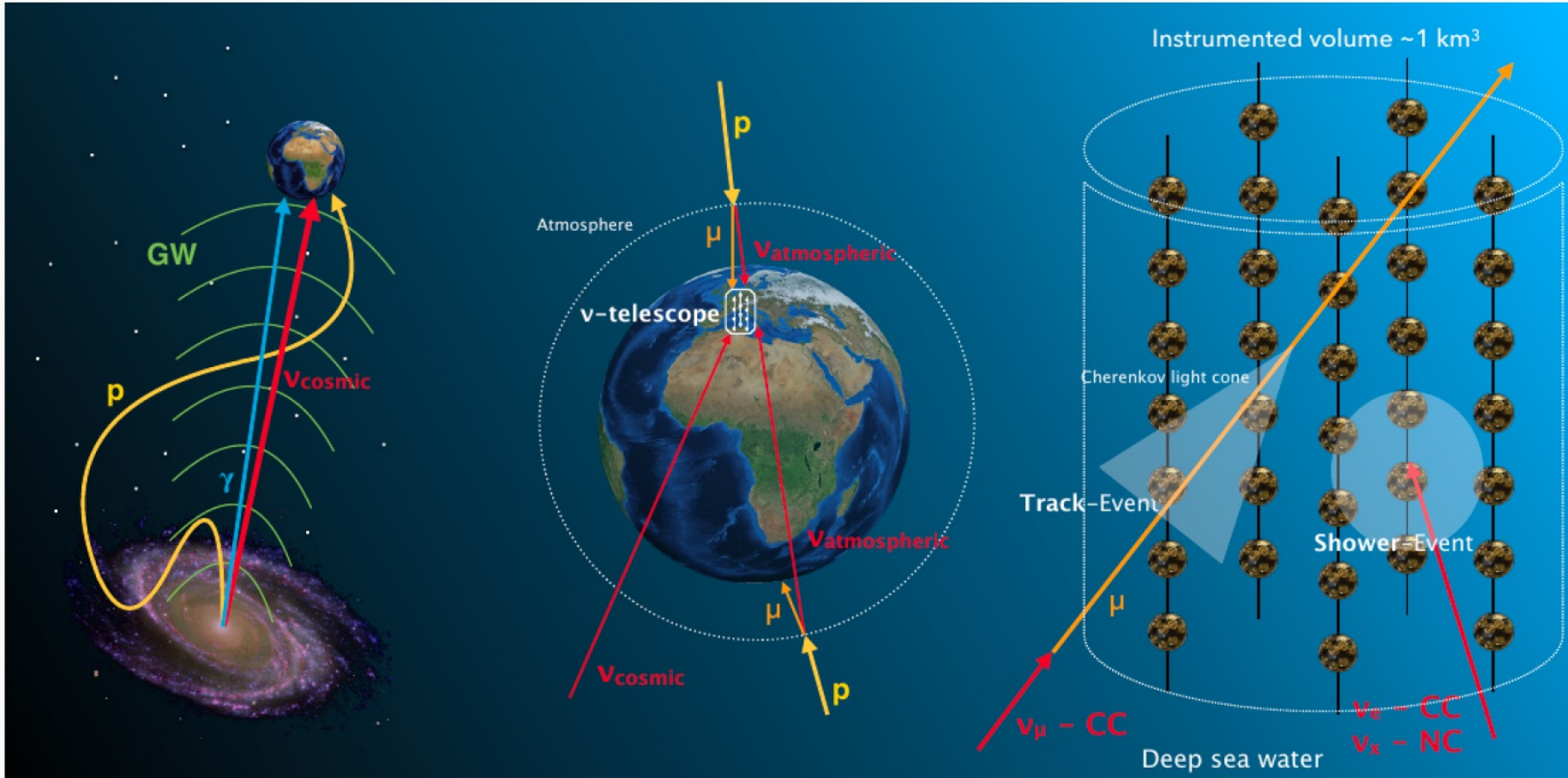
✓ Core of astrophysical bodies

✓ Point source

✓ Cosmological distance

⇒ Signature of hadronic acceleration

Detection principles




The neutrino telescope world map



P-ONE
R&D



ANTARES
first in Deep sea
0.01 km³
2008 – 2022



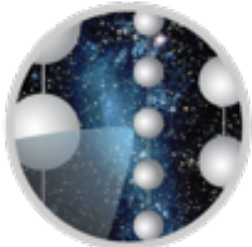
KM3NeT
Deep sea
1 + 0.006 km³
Construction



Baikal/GVD
Deep lake
~1 km³
Construction



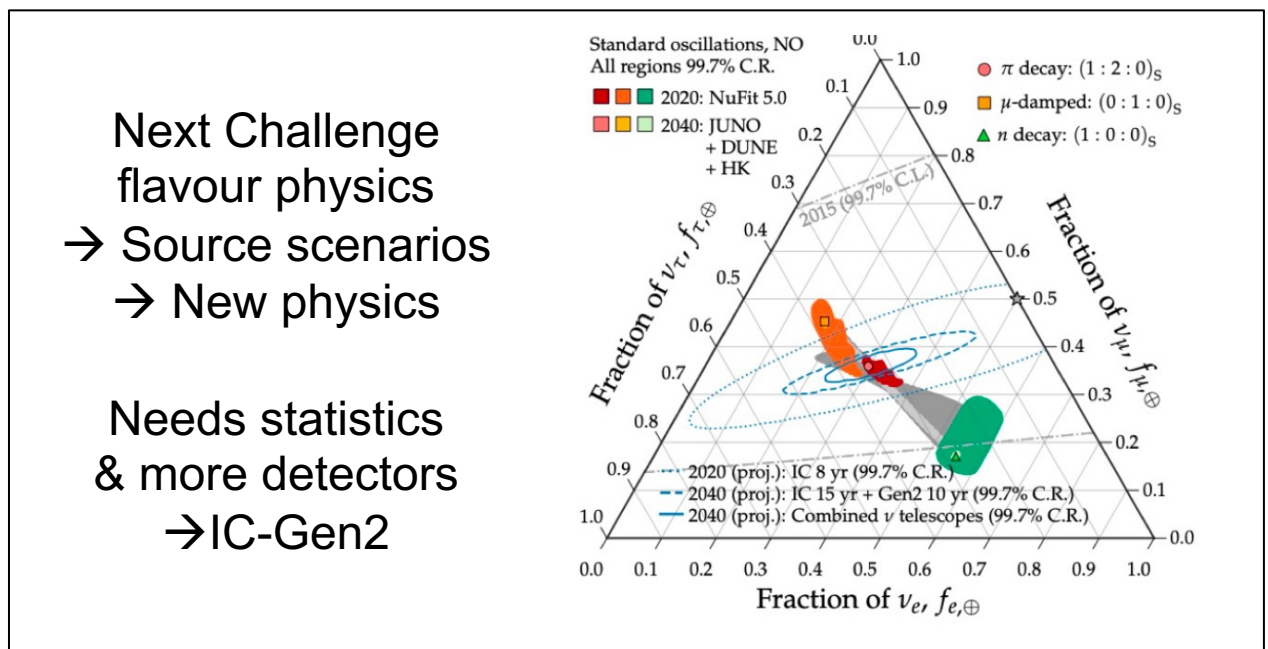
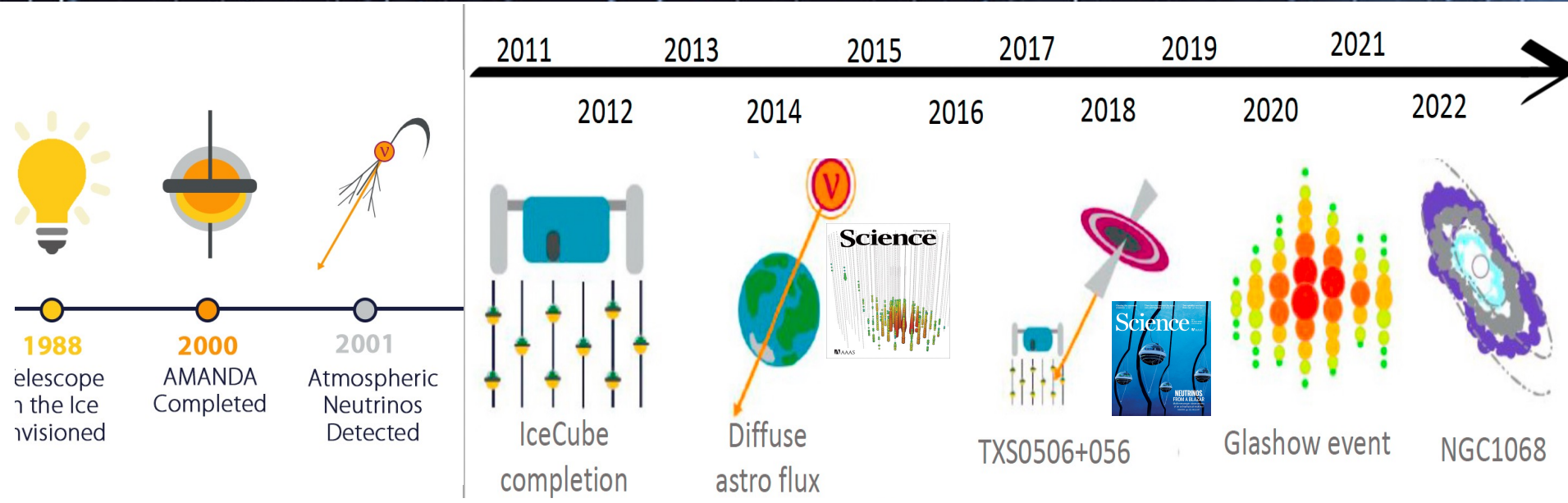
GNN
The GLOBAL NEUTRINO NETWORK



IceCube
Deep ice
1 km³
2011 –
Upgrade
with a lower
energy threshold
Construction
Gen-2 : 10 km³


www.globalneutrino.org
Frame for enhanced cooperation

IceCube opened the field with km-scale detector



Developments in the Mediterranean Sea

ANTARES – the first undersea neutrino telescope

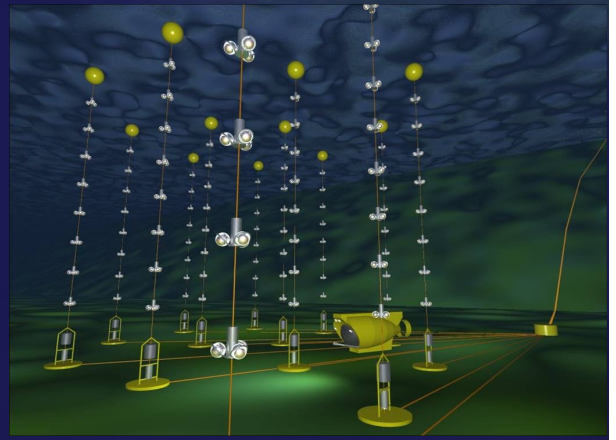


Toulon

Institut M.Pacha

A multi disciplinary observatory with > 15 years of data

- Competitive physics results & intriguing hints
- Constraints on neutrinos as seen by IceCube.
- Extensive multi-messenger program.
- Joint studies with several partners.
- About 100 papers published & 100 PhD students
- **QUITE AN ADVENTURE ! But only the beginning ...**



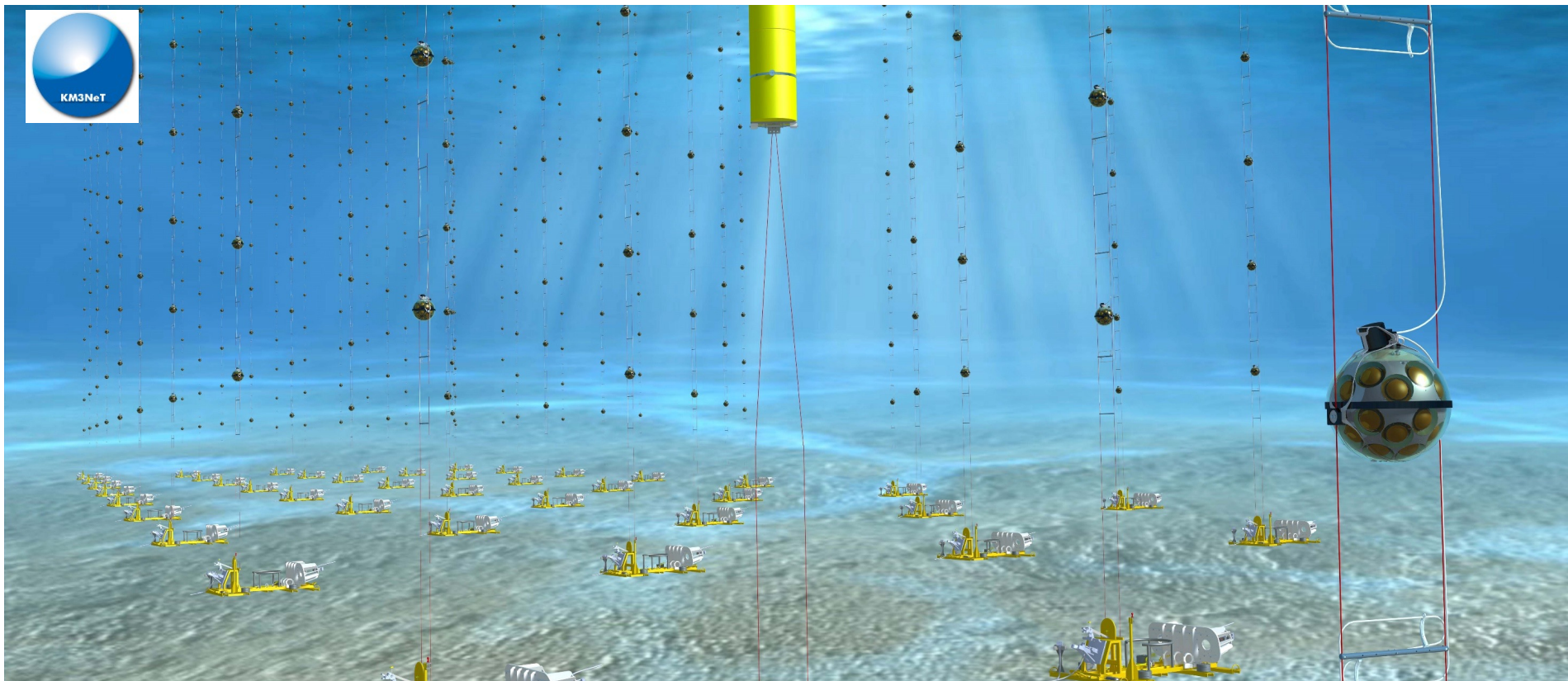
Site ANTARES
42 50'N, 6 10'E

Google™

Developments in the Mediterranean Sea

Next is KM3NeT - ESFRI project

- ARCA (high-energy neutrino astronomy, Italian site)
 - Installation started, completed 2026
 - Discovery and subsequent observation of neutrino sources
- ORCA (low-energy neutrino physics, French site)
 - Installation started, completion 2025
 - Determination of mass ordering of neutrinos

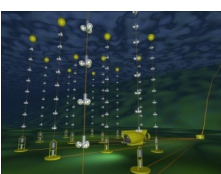


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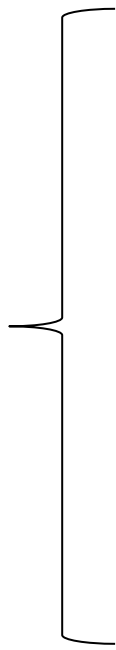


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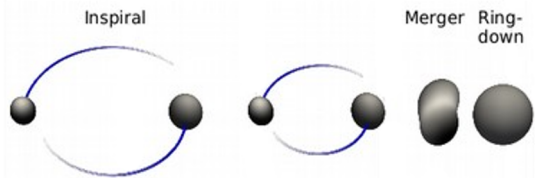
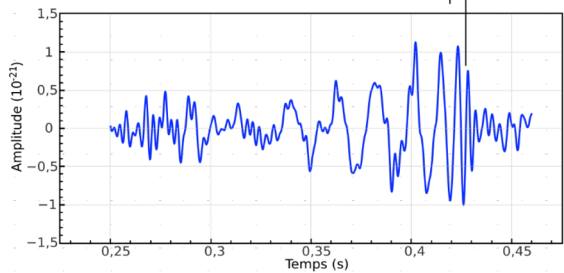
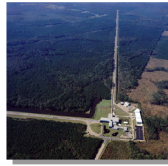
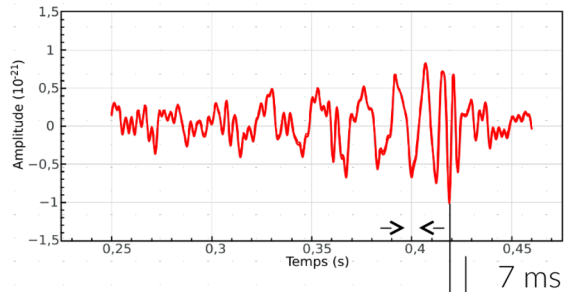


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First Detections

Sep 14, 2015



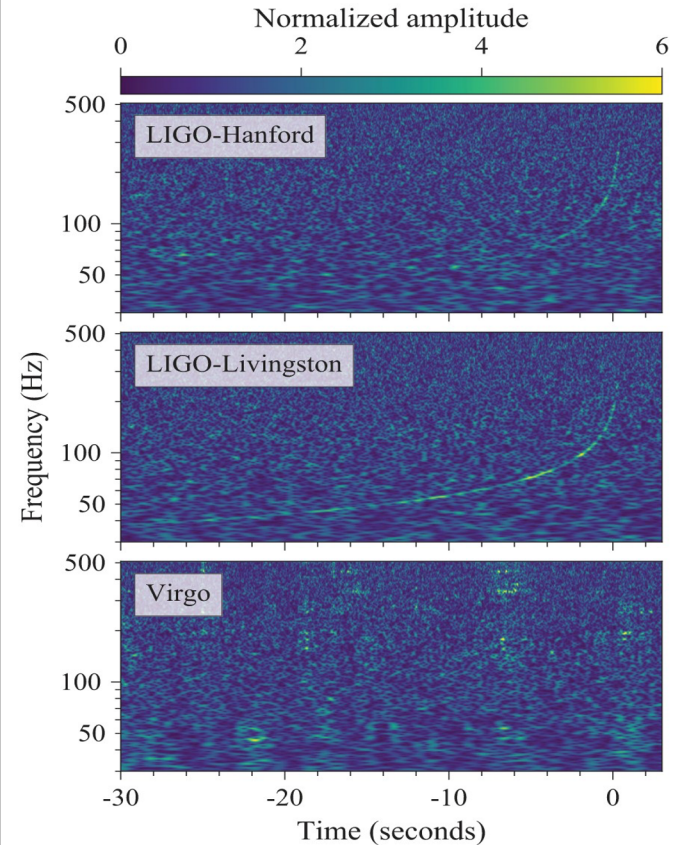
$$m_1 = 36.2_{-3.8}^{+5.2} M_{\odot}$$

$$m_2 = 29.1_{-3.7}^{+4.4} M_{\odot}$$

$$D_L = 420_{-150}^{+180} \text{ Mpc}$$

1.4 billions light-year

Aug 17, 2017



GW170817

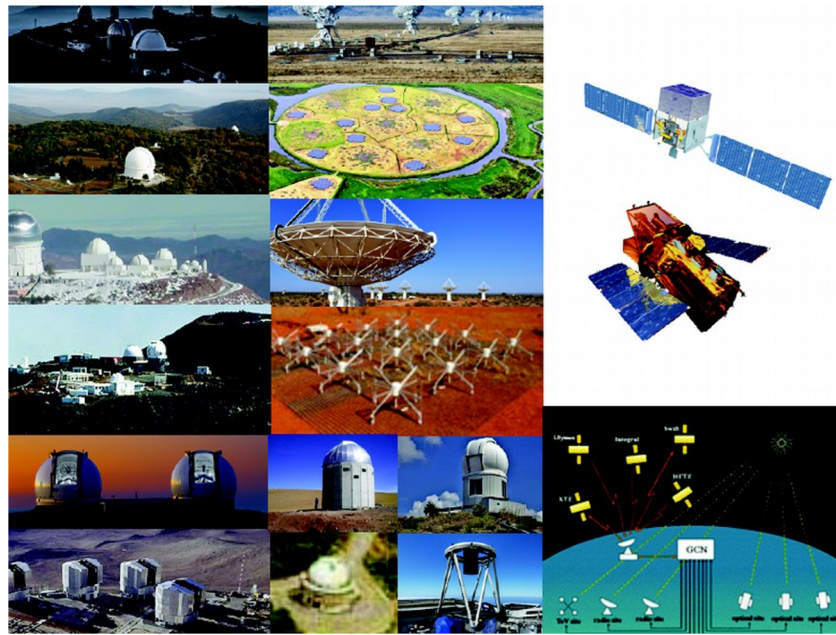
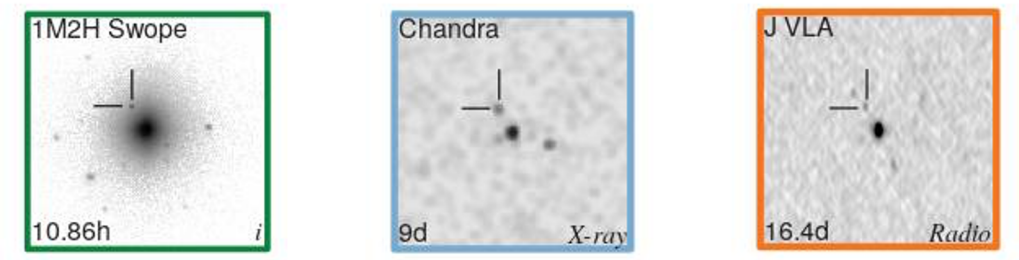
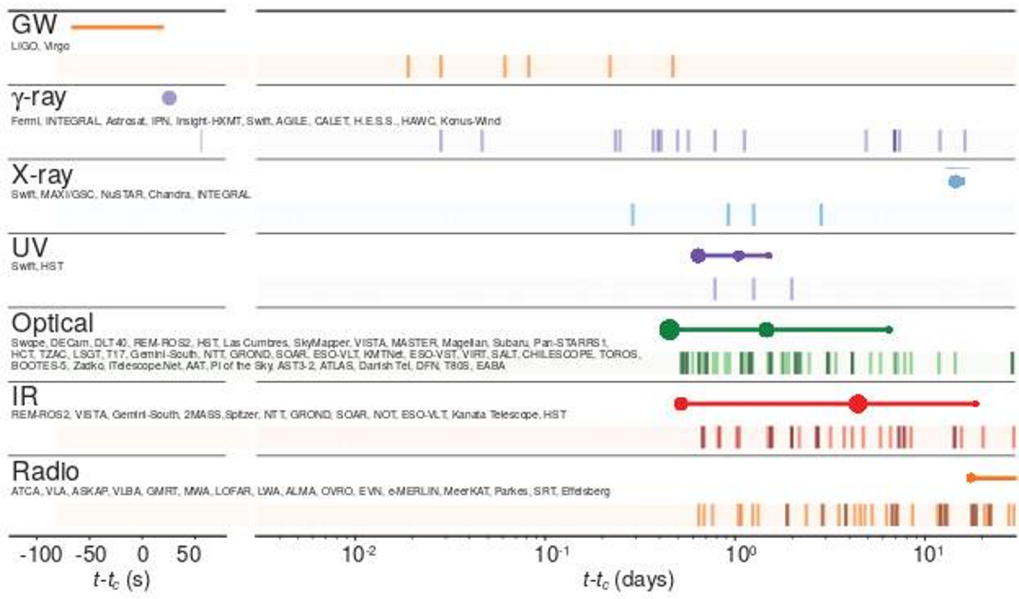
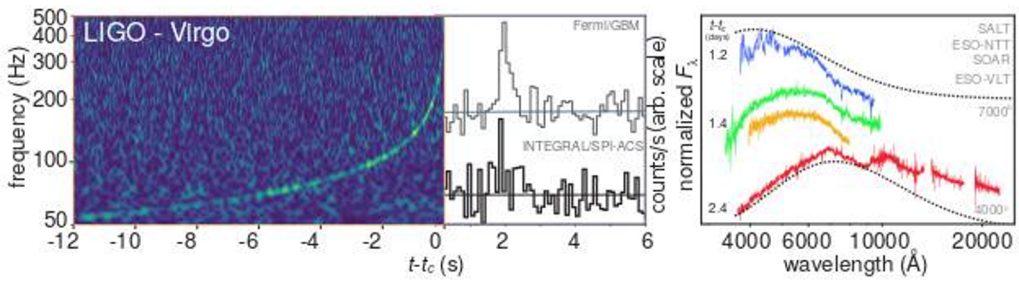
$$m_1 = 1.46_{-0.10}^{+0.12} M_{\odot}$$

$$m_2 = 1.27_{-0.09}^{+0.09} M_{\odot}$$

$$D_L = 40_{-15}^{+7} \text{ Mpc}$$

(130 millions light-years)

GW170817: multi-messengers !



Association with
gamma-ray
bursts
Jet of relativistic
plasma?
 $|c/c_g - 1| < 5 \times 10^{-16}$



Achievements so far

A large population of “heavy” binary black holes, so far unobserved

Raises many questions

How do they form?

In what environment?

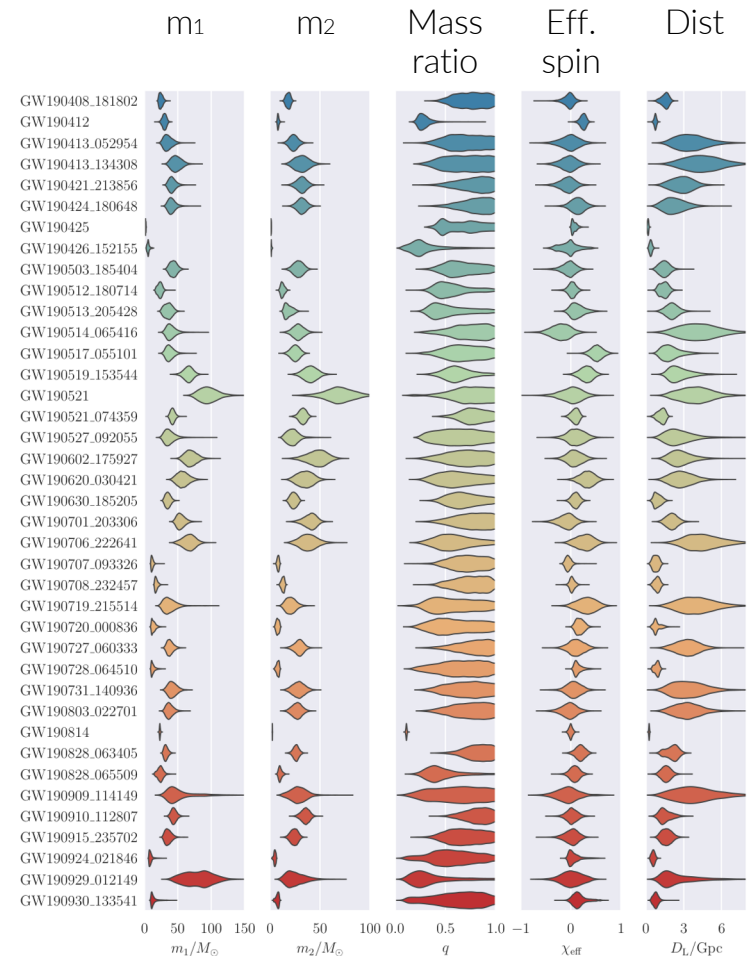
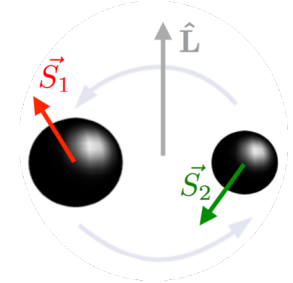
Is there a single formation channel?

Some of the detected binaries are incompatible with the current understanding of black hole formation from massive stars

Other types of binary systems

Binary neutron stars
Possible mixed black hole and neutron star binaries

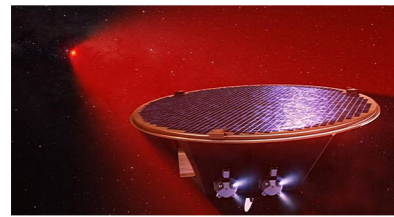
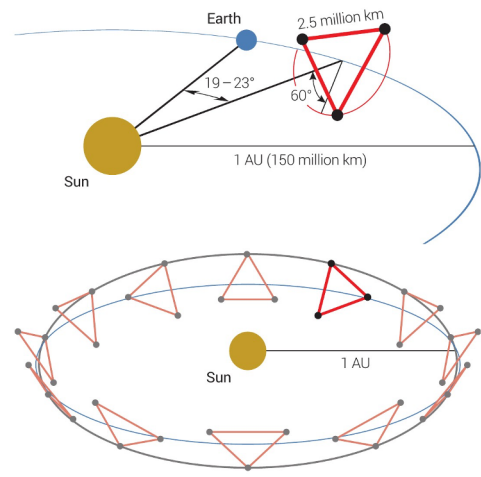
GWTC 2



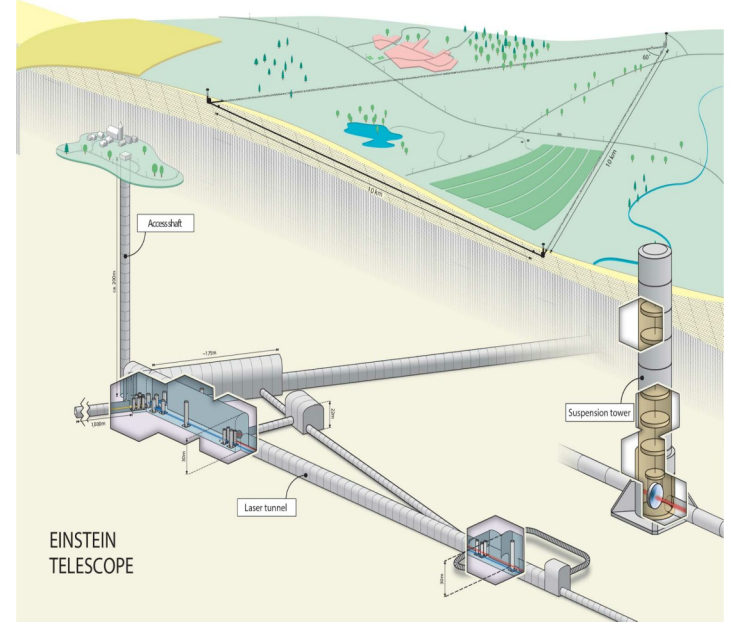
What next ? Cover frequency range



Future ESA mission



Einstein telescope Target : mid-2030
Artist view



Target heaviest and most diverse objects

Trace the history of black holes across all stages of galaxy evolution

Constrain deviation from the Kerr metric of General relativity.

3rd generation detectors
x 10 sensitivity improvement

Exceptional science reach
~90 % of all BBH mergers in the Universe
1 BBH every 30 sec

A Citizen Matter

Le Monde WEEK-END

CULTURE & IDÉES
EUROPE LE SECRET MARCHANDAGE DES NORMES DIESEL

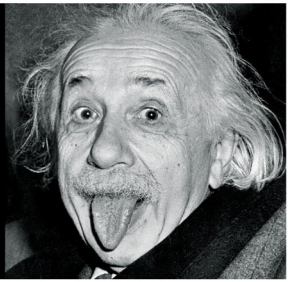
Samedi 19 février 2016 - 77^e année - N° 2208 - 4,20 € - France métropolitaine - www.lemonde.fr - Fondateur : Hubert Beuve-Méry



Les petits calculs d'un remaniement

- François Hollande a choisi un remaniement tactique pour neutraliser les différentes composantes de la majorité avant 2017
- L'arrivée de trois ministres écologistes, dont Emmanuel Coste, porte un coup sévère au parti de Cécile Duflot
- Le retour de l'ancien premier ministre Jean-Marc Ayrault, nommé au Quai d'Orsay, tient à la volonté de rassurer la gauche du PS
- Le chef de l'Etat a choisi l'une de ses proches, Audrey Azoulay, pour remplacer Fleur Pellerin au ministère de la culture
- Plus qu'un gouvernement de combat, c'est un gouvernement de contrats qu'a choisi le président

Einstein avait raison



Le physicien avait prédit l'existence des ondes gravitationnelles en 1916, la preuve directe en est aujourd'hui apportée

Cette découverte majeure est à l'angor parmi les plus grandes percées de la connaissance

Les trous noirs existent et l'espace-temps est bien un contenant élastique susceptible d'onduler

UN AIR DE LEHMAN BROTHERS

Pourquoi les banques sont attaquées en Bourse

Les banques mondiales, détachées de la réalité économique et de la confiance des investisseurs, font peser les risques systémiques et compromettent l'ensemble de l'économie mondiale. Les investisseurs ne sont pas prêts à accepter une rentabilité insuffisante et des engagements trop élevés, notamment dans le secteur financier. Les investisseurs de responsabilité politique ont une vision plus réaliste des conditions économiques et font peser le spectre d'un mouvement de sortie de capital.

UN AIR DE LEHMAN BROTHERS

- UN PRICE E 2,00
- COMMERCE LE GRAND BAZAR DU TRAVAIL DU DIMANCHE
- CULTURE LE THÉÂTRE DE L'EUROPE AUX EUROPÉENS
- CULTURE & IDÉES LIVRES POLITIQUES: FAUT-IL ÊTRE LU POUR ÊTRE LU ?
- SPORT & FORME PROFESSION GLADIATEUR

Plan russo-américain pour une trêve en Syrie

Sergueï Lavrov et John Kerry s'accordent sur un accès humanitaire aux villes assiégées



MIGRANTS L'ÉTAT VA RASER LA MOITIE DE LA JUNGLE DE CALAIS

PAR MARCOFINO BARRAUD

La partie du Pas-de-Calais où sévissent de manière chronique les flux de la « jungle » de Calais, l'ancienne décharge désaffectée, va être rasée. Les migrants, nombreux à la fin de l'été, ont été évacués dans des containers et des tentes. Les autorités ont décidé de raser la moitié de la jungle de Calais, ce qui permettra de créer un espace de vie pour les migrants. Les autorités ont également décidé de créer un espace de vie pour les migrants.

BACCHANALES MODERNES !

Le nu, l'ivresse et la danse dans l'art français du XIX^e siècle



12 février - 23 mai 2016

"All the News That's Fit to Print"

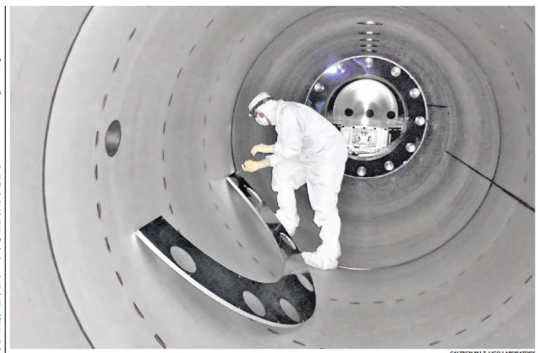
The New York Times

VOL. CLXV... No. 57,140 + NEW YORK, FRIDAY, FEBRUARY 12, 2016 \$2.50

Clinton Paints Sanders Plans As Unrealistic

New Lines of Attack at Milwaukee Debate

By AMY CHOIZICK and PATRICK HEALY
MILWAUKEE — Hillary Clinton, scrambling to recover from her double-digit defeat in the New Hampshire primary, repeatedly challenged the trillion-dollar policy plans of Bernie Sanders at their presidential debate on Thursday night and portrayed him as a big talker who needed to "level" with voters about the difficulty of accomplishing his agenda.



A worker installed a baffle in 2010 to control light in the Laser Interferometer Gravitational-Wave Observatory in Hanford, Wash.

WITH FAINT CHIRP, SCIENTISTS PROVE EINSTEIN CORRECT

A RIPPLE IN SPACE-TIME

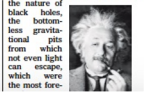
By DENNIS OVERBYE
A team of scientists announced on Thursday that they had heard and recorded the sound of two black holes colliding a billion light-years away, a fleeting chirp that fulfilled the last prediction of Einstein's general theory of relativity.

Long in Clinton's Corner, Blacks Notice Sanders

By RICHARD FUSSETT
ORANGEBURG, S.C. — When Helen Duley was asked whom she would vote for in the South...

Last Occupier In Rural Oregon Is Coaxed Out

candidate she barely knew. "It makes me feel good," she said, chuckling, "that young people are listening to the elderly people." She now said she was an un-



She now said she was an un-

Additional Challenges: citizen science



Minimizing the knowledge gap
between Large Research Infrastructures
and Society through Citizen Science

DISCOVER OUR FOUR DEMONSTRATORS

<https://www.reinforceeu.eu>



GRAVITATIONAL
WAVE NOISE
HUNTING



DEEP SEA HUNTERS



SEARCH FOR
NEW PARTICLES
AT THE LHC



COSMIC MUONS
IMAGES

Initiated by S. Katsanevas

Stavros Katsanevas (1953 - 2022)



Photo credits: «Stavros Katsanevas» Grèce 2020 - Nikos Aliagas

Several testimonies related to his various activities including fRound table on “art & sciences”.

Confirmed speakers

- **Astroparticle convergence in Europe** / Frank Linde - Christian Spiering
- **Early times in CERN** / Francois Richard
- **EGO - environmental aspects** / Irene Fiori - Maria Tringali
- **Onset of astroparticle in France** / Michel Spiro
- **Astrocent** / Leszek Roszkowski
- **Sonification** / Wanda Diaz Merced
- **Early career in Lyon** / Imad Laktineh
- **Muography** / Jacques Marteau
- **Gravitational wave detection on the Moon** / Philippe Lognonné
- **Geoscience and astrophysics** / Claude Jaupart
- **Creation of the DiIP** / Themis Palpanas
- **APOGEIA** / Veronique Van Elewick

SYMPOSIUM STAVROS KATSANEVAS

Thursday, June 1st, 2023
 Université Paris Cité - Laboratoire APC
 Paris, France

This symposium is organized in memory of Stavros Katsanevas (1953-2022) by his colleagues and friends to honor his multiple contributions, from astroparticle physics to art in science.

Registration is free but mandatory

More info:

https://indico.in2p3.fr/e/stavros_katsanevas

Tomás Saraceno

Collage towards Gravitational Waves, 2017 - Ongoing
 Z33 House for Contemporary Art, Genk, Belgium. Curated by Jan Boelen.
 Courtesy the artist and Z33 House for Contemporary Art.

©Tomás Saraceno



Just a beginning ...

ACAME

The logo for ACAME is displayed in large, stylized letters. The 'A' and 'C' are filled with a bright, fiery orange and yellow glow, while the 'M' and 'E' are hollow. A satellite with blue solar panels is positioned inside the top bar of the letter 'E'. The background is a dark space filled with stars and a complex network of glowing purple and red lines.

Astrophysics Center for Multimessenger studies in Europe

Gravitational waves, Cosmic rays, Neutrinos
VHE gamma-rays, X-rays, Optical, Radio

