## Status of the SoLid experiment

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Universiteit Antwerpen

SoLid

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#### Overview

- Introducing the SoLid experiment
- Background description
- Signal selection
- Alternative analysis: Heavy Neutral Leptons
- Upgrade to the Phase 2 detector

## Introducing the SoLid experiment

#### Physics motivation



- Reactor antineutrino anomaly
  - Consistent deficit observed at short (< 1 km) baselines compared to predictions
  - Deficit could be explained by an additional (sterile) neutrino of  $△m^2 \cong 1-10 \text{ eV}^2$
  - Sterile neutrino hypothesis given more weight by unrelated anomalies (Gallium, LSND)
- Reactor antineutrino spectrum distortion (a.k.a. the 5 MeV bump)
  - Excess observed at 5 MeV by most large reactor experiments
  - Among the fissile isotopes in commercial reactors, <sup>235</sup>U is considered most likely
- Also an anti-proliferation component

#### The BR2 reactor



- Belgian Research reactor 2
- Located on the SCK-CEN site in Mol, Belgium
- Rated for 50 100 MW<sub>Th</sub>
  - Typically 60 MW<sub>Th</sub>
  - 5 or 6 month-long reactor cycles per year
- Highly enriched <sup>235</sup>U
- Compact conical core
  - Ø ~ 0.5 m
  - h ~ 1 m
  - Experimental hall starts as close as 5.5 m from the core
  - Low neutron and gamma backgrounds in experimental hall
- 37 m above sea level, 6-8 m MWE overburden

#### The SoLid detector

#### A highly segmented modular antineutrino detector using dual solid scintillators and multiplexed readout

- Highly segmented
  - Built from 12.800 optically isolated cells
  - Each cell measures 5 x 5 x 5 cm<sup>3</sup>
- Modular
  - 16 x 16 cells make a plane
  - 10 planes make a module
  - Detector consists of 5 modules
- Dual solid scintillators
  - PVT cube as neutrino target and for positron and gamma detection
  - <sup>6</sup>LiF:ZnS(Ag) layers for neutron capture and detection
- Multiplexed readout
  - 64 WLS fibres bring light from the cells to the edge of the detector
  - Each fibre is read out by a SiPM
- 3200 fibre-SiPMs pairs service 12800 cells Simon Vercaemer - EOS winter solstice meeting 2020

#### Antineutrino detector



#### The SoLid detector

A highly segmented modular antineutrino detector using dual solid scintillators and multiplexed readout



Antineutrino detector

- Inverse beta decay reaction:  $v + p \rightarrow n + e^+$
- Prompt signal from e<sup>+</sup> scintillation and annihilation gammas in PVT
  - Fast scintillator, very brief pulse (few ns) 0
  - Provides v interaction cube 0
- Delayed signal from capture of thermalised neutron on <sup>6</sup>Li
  - $n + {}^{6}Li \rightarrow \alpha + {}^{3}H$  $\bigcirc$
  - $\alpha$  and <sup>3</sup>H scintillate in ZnS(Ag)  $\bigcirc$
  - Slow scintillator, extended pulse (10s  $\mu$ s) 0
  - Neutron cube close to v cube
  - Neutron capture time:  $\tau = 68 \, \mu s$ 0

### Trigger system

- Random trigger
  - Operates at 1 Hz
- Threshold trigger
  - Triggers signal above 2 MeV threshold
  - Coincidence required between horizontal and vertical fibre, within 75 ns
- Neutron trigger
  - Targets neutron scintillation in ZnS(Ag)
  - Counts peaks over threshold in rolling time window



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### Data collection

- Random trigger
  - Reads full detector for 13.6  $\mu$ s
  - Saves Raw waveform 0
- Threshold trigger
  - Reads triggering plane for 13.6  $\mu$ s 0
  - Suppresses signal below ~100 keV 0
- Neutron trigger
  - Reads triggering plane and 3 or 4 neighbouring planes
  - Reads 500  $\mu$ s before and 200  $\mu$ s after trigger 0

<sup>0</sup> <sub>2 4</sub>

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X[cube]

Suppresses signal below ~100 keV 0

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# **Background description**

#### Backgrounds: BiPo



Radioactive decay sequence in the Uranium series:

1.  ${}^{214}\text{Bi} \rightarrow {}^{214}\text{Po} + e^-$  Q = 3.27 MeV

#### e<sup>-</sup> mimics prompt signal

- 2.  ${}^{214}\text{Po} \rightarrow {}^{210}\text{Pb} + \alpha$   $t_{1/2} = 168 \,\mu\text{s}$  $\alpha$  mimics delayed signal when in ZnS
- Internal constant contamination in ZnS layers
- External variable source: <sup>222</sup>Rn release from concrete

#### Backgrounds: BiPo



- Exploit difference between IBD's  $\alpha$  + <sup>3</sup>H and BiPo's single  $\alpha$ 
  - Different energies (4.8 MeV vs 7.8 MeV)
  - 2 particles vs only 1
  - > Slightly different scintillation pattern: 'BiPonisher'
- Lack of annihilation gammas
- Limited energy range
- Differences in  $\Delta T$  and topology



#### Backgrounds: cosmic neutrons

High energy neutrons created by cosmic rays:

- 1. Recoil on nuclei in the detector Recoil mimics prompt
- 2. Neutron thermalises and captures Identical to IBD neutron capture
- High rate due to low overburden
- Pressure dependent rate
- Main source of background
  - Low overburden
  - Exponentially decreasing energy spectrum over IBD energy range (and beyond)
  - High variety of topologies
  - $\circ$  Virtually identical  $\Delta T$  to IBD



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# Signal selection

#### **Basic sequential selections**



- Prompt requirements
  - Energy
  - Energy balance
  - Spatial spread
- Delayed requirements
  - BiPonisher
- Coincidence requirements
  - **⊿**T
  - $\circ$   $\Delta X$ ,  $\Delta Y$ ,  $\Delta Z$ ,  $\Delta R$
- → ~ 10% IBD efficiency, S/B ≅ 0.06

### Neutrino signal in data

- BiPo varies with Rn releases
  - Can be determined in situ from high  $\Delta T$ and low BiPonisher coincidence data



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SoLid Preliminary

- Cosmic neutron rate varies with atmospheric pressure
  - Pressure dependence established during reactor off period
  - Extrapolated rate subtracted from reactor on period



#### Neutrino excess



- Observed excess consistent with IBD simulation
- We are sitting on a lot more data
- Improvements to the reconstruction in the pipeline (next slide)
- A major detector upgrade took place over summer (last slide)

### Moving forwards

Annihilation gammas (Topology):

- Brings new variables
  - Number of gammas reconstructed 0
  - Energy
  - Opening angle
  - Distances 0
- Requires lower offline thresholds
  - Better understanding needed
  - Balancing efficiency and contamination from dark counts
- Improved background rejection  $\rightarrow$



Machine learning:

- Improved background rejection dual approach (no cutting edge)
  - UBDT 0
  - TMVA Neural Network 0
- Improve BiPonisher to BiPonator
  - Simple ratio to CNN
  - Expected 2-3 x improvement in alpha tagging



## Alternative analysis: Heavy Neutral Leptons

#### Heavy neutral leptons

- nuMSM introduces 3 right handed neutrinos
- Virtually no mass limit HNLs
- Resolves significant issues
  - Neutrino masses (seesaw)
  - Universe's baryon asymmetry
  - Dark matter candidates
- Sterile, only produced via mixing
- Small mixing angle, low production rate
- Unstable, detect decay products
  - $\circ \quad \text{Radiative} \quad \begin{array}{l} N_i \rightarrow \nu_j + \gamma \\ N_i \rightarrow \nu_j + \gamma + \gamma & \text{if } m(N_i) > m(\nu_j) \end{array}$

$$\circ$$
 Invisible  $N_i \rightarrow \nu_j + \nu_k + \overline{\nu}$ 

 $\circ$  e<sup>+</sup>e<sup>-</sup> mode  $N_i \rightarrow \nu_j + e^+ + e^-$  if  $m(N) > 2m_e = 1.022$  MeV



#### SoLid as HNL detector

- BR2 as neutrino source
  - ~ 60 MW<sub>Th</sub> →  $12x10^{18} \nu/s$  (12 EBq)
  - 12 EBq is isotropic, need to apply geometric efficiency (~ 0.13 %) → 16 PBq
  - $\circ \qquad \text{Small mixing angle} \rightarrow \text{Less Bq}$
  - Long decay time  $\rightarrow$  O( 10<sup>-4</sup> Bq)
- Mass range limited by <sup>235</sup>U v spectrum and e<sup>+</sup>e<sup>-</sup> decay mode requirement
  - $\circ$  1.022 MeV < m(N)  $\lesssim$  9 MeV
- No neutron in e<sup>+</sup>e<sup>-</sup> decay
  - $\rightarrow$  rely on threshold trigger
    - Minimum 2 MeV visible energy
    - Single plane only



#### Simulation work

- Background simulations from the IBD analysis can be recycled
  - Cosmic background most challenging
- Signal simulation using Pythia
  - Samples with several HNL masses have been created
  - Energy spectrum from reactor simulation
- Trigger studies have been performed
  - Near perfect trigger efficiency for higher masses, also lowest statistics
  - Reduces a bit when reconstruction is taken into account
  - $\circ$  v energy more important than HNL mass



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### Signal selection

#### • Preliminary list of variables composed

- Neutron/muon/alpha veto
- Fiducialization
- ES energy
- ES spatial spread
- Still being refined/expanded
- Manual and TMVA optimizations under way
- Hypothesis testing scripts are being set up



# Upgrade to the Phase 2 detector

#### Detector upgrade







• July: brought Phase 1 to Antwerp

- Replaced all SiPMs by latest generation devices over summer, modified electronics
- October: commissioning of Phase 2 at BR2

• November: start of datataking





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