

Dark forces from right-handed neutrinos

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<https://xunjiexu.github.io/>

[1] X. Xu, JHEP 09, 105, (2020), [2007.01893]

[2] Garv Chauhan, X. Xu, [2012.xxxxx]

The SM and fundamental forces

1. Gravity

graviton: 

2. Electromagnetic force

photon: 

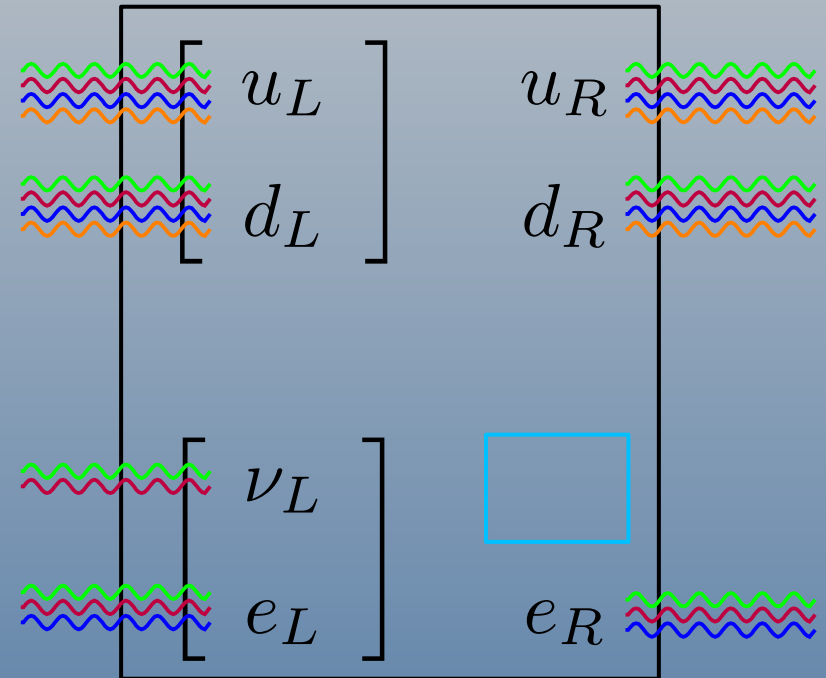
3. Weak force

W^\pm/Z : 

4. Strong force

gluons: 

5. ...



Fundamental long-range forces

Long-range forces in spin-2? Gravity

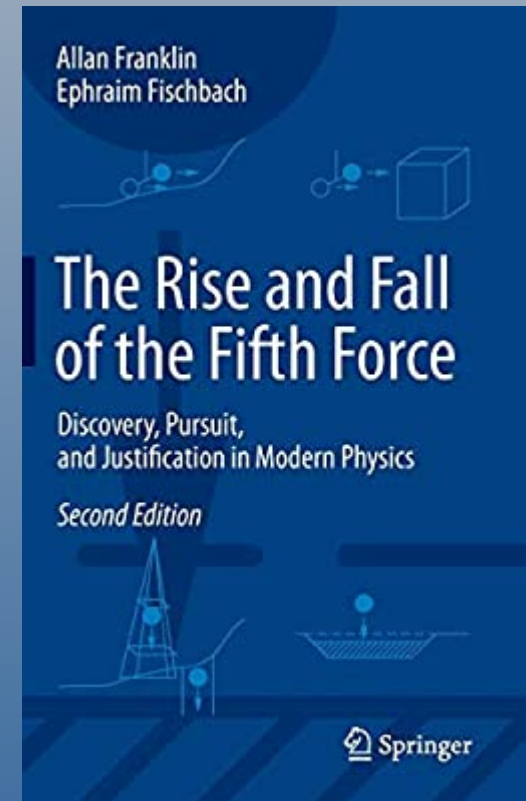
spin-1? Photon

spin-0?

Search for long-range Yukawa forces?

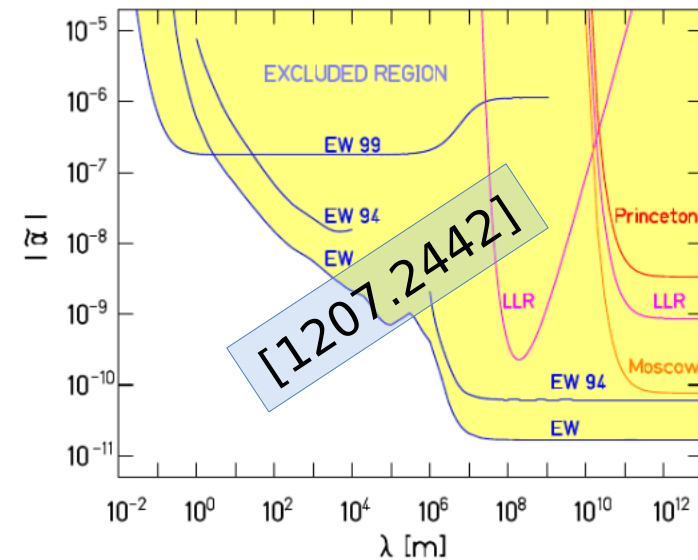
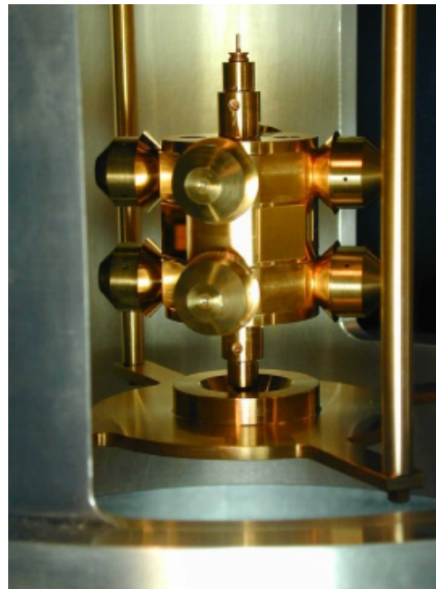
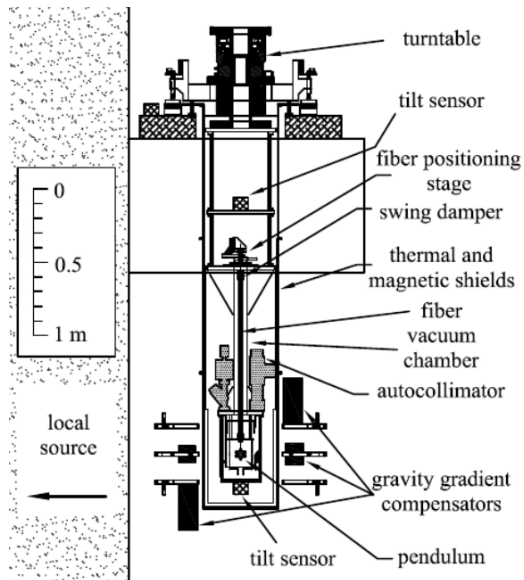
... not new ...

from 1980s to 1990s, already ...



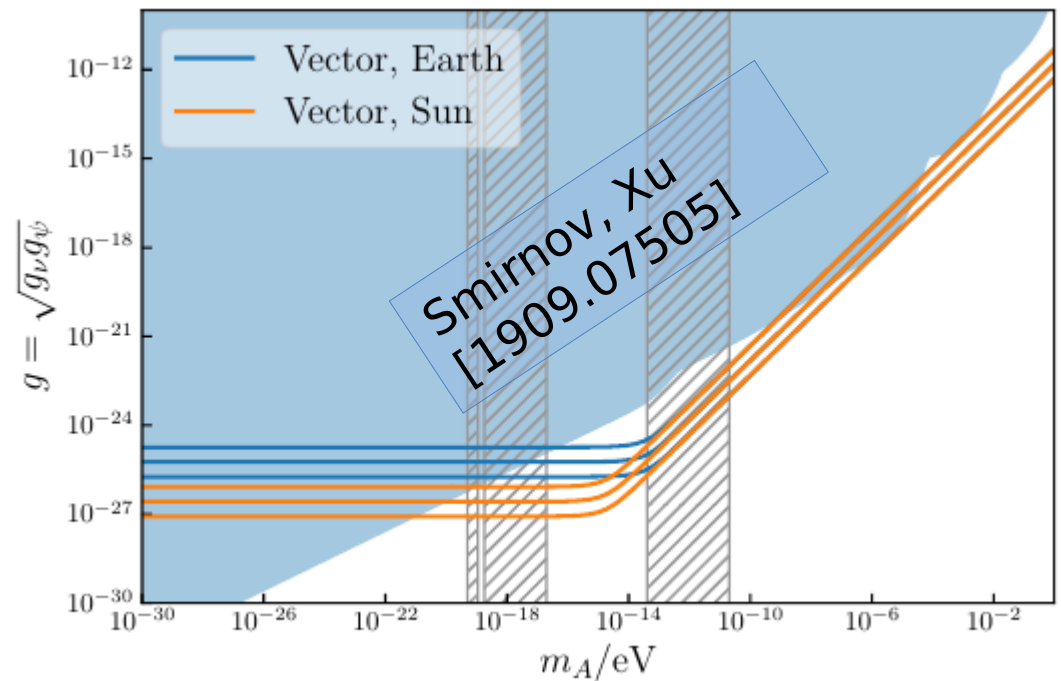
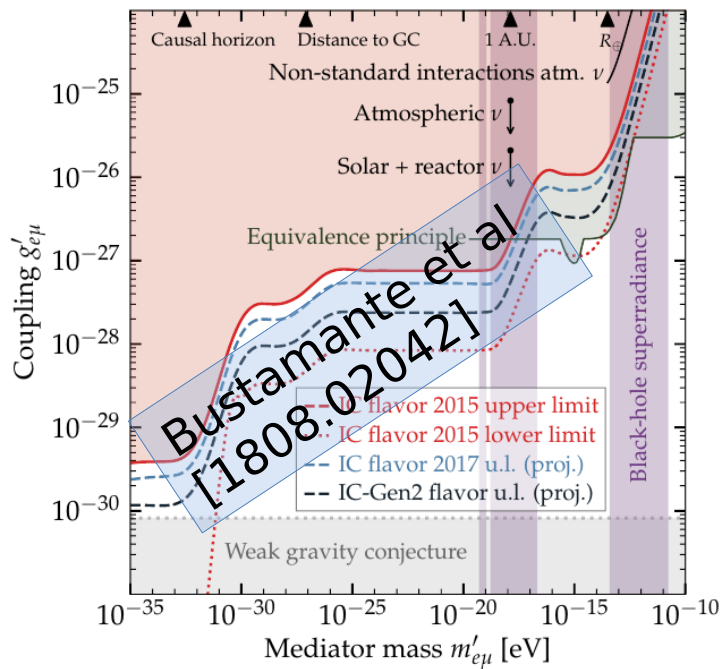
Progress still going on ...

Torsion-balance tests of Equivalence Principle
Lunar Laser-Ranging (LLR) technology
Long-range forces in neutrino oscillation
And, of course, LIGO/VIRGO ...



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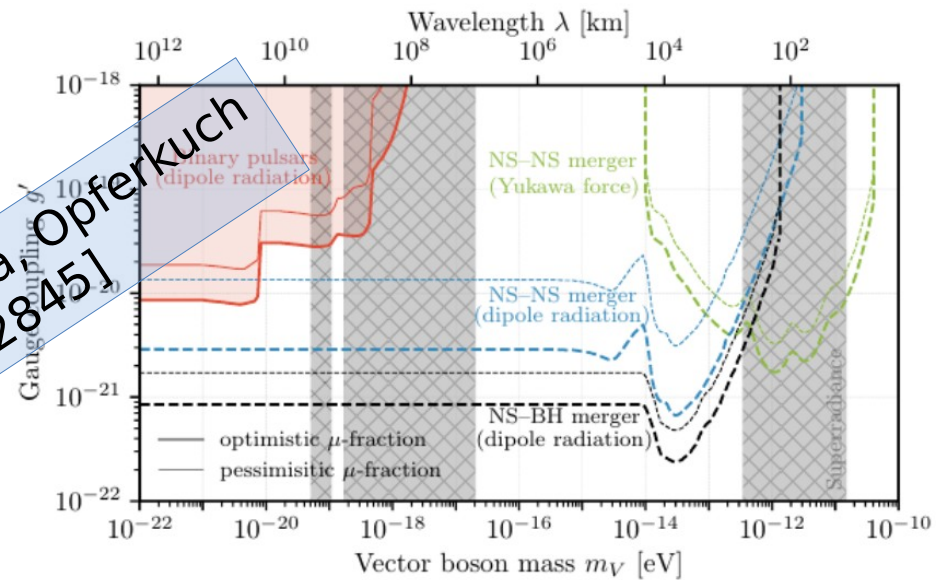
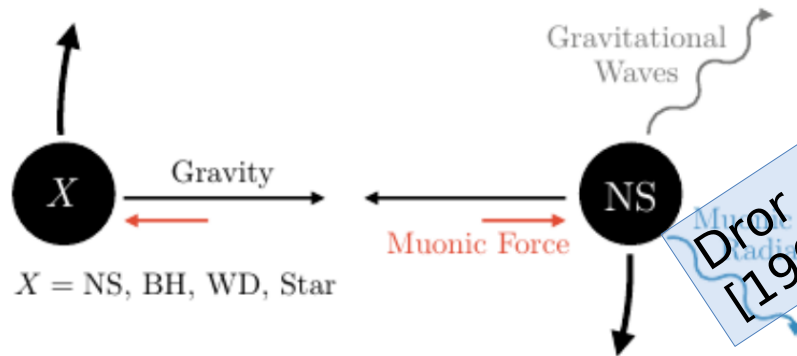
Progress still going on ...

Torsion-balance tests of Equivalence Principle

Lunar Laser-Ranging (LLR) technology

Long-range forces in neutrino oscillation

And, of course, LIGO/VIRGO ...



Experimental bounds:

..., 10^{-10} , 10^{-15} , 10^{-20} , ...

How weak do you expect the new force to be?

How weak do you expect the new force to be?

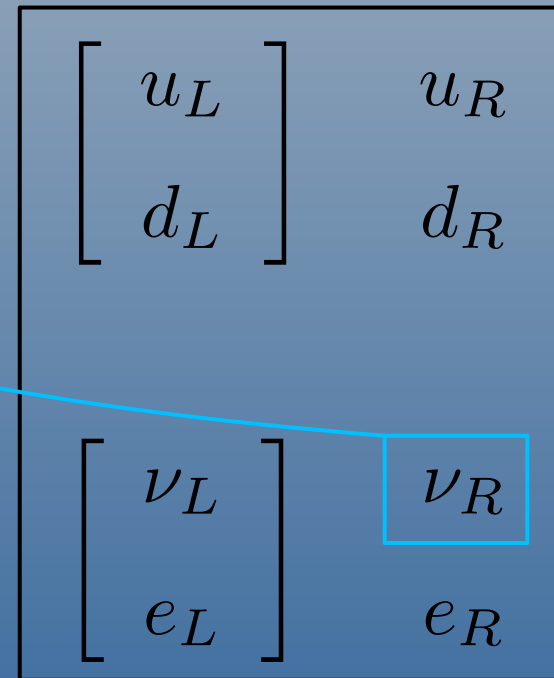
An answer from ν_R -philic scalar:

$$y_e \sim \frac{G_F m_e m_\nu}{16\pi^2} \sim \mathcal{O}(10^{-21}), y_\mu \sim \frac{G_F m_\mu m_\nu}{16\pi^2} \sim \mathcal{O}(10^{-19})$$

Adding the missing piece of SM,
we get a ν_R -portal:

$$\mathcal{L} \supset \nu_R \nu_R \phi$$

and hence a new dark force

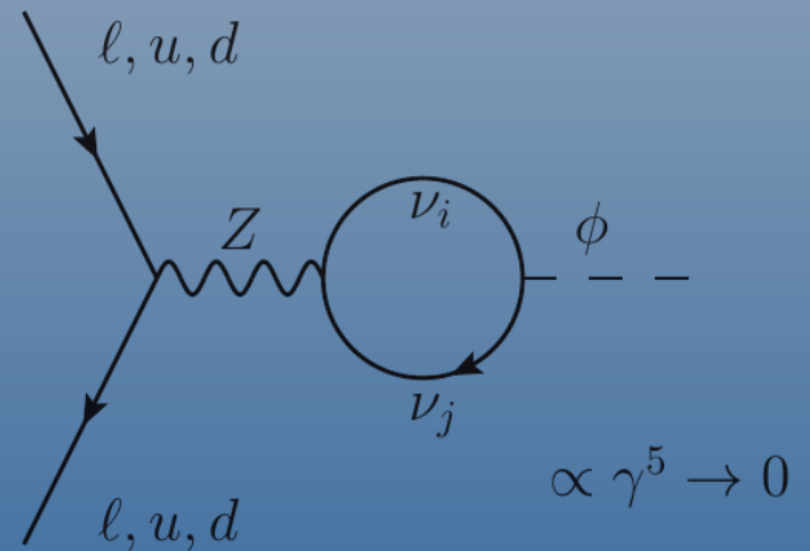
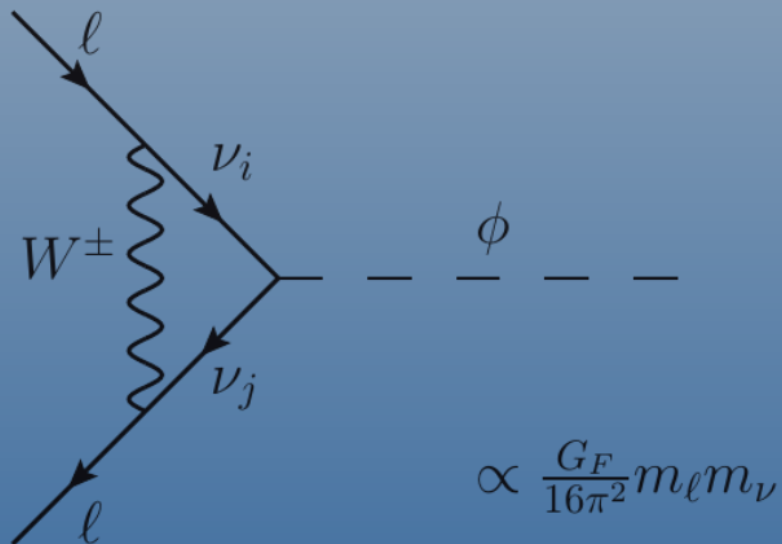


The new dark force from

$$\mathcal{L} \supset \nu_R \nu_R \phi$$

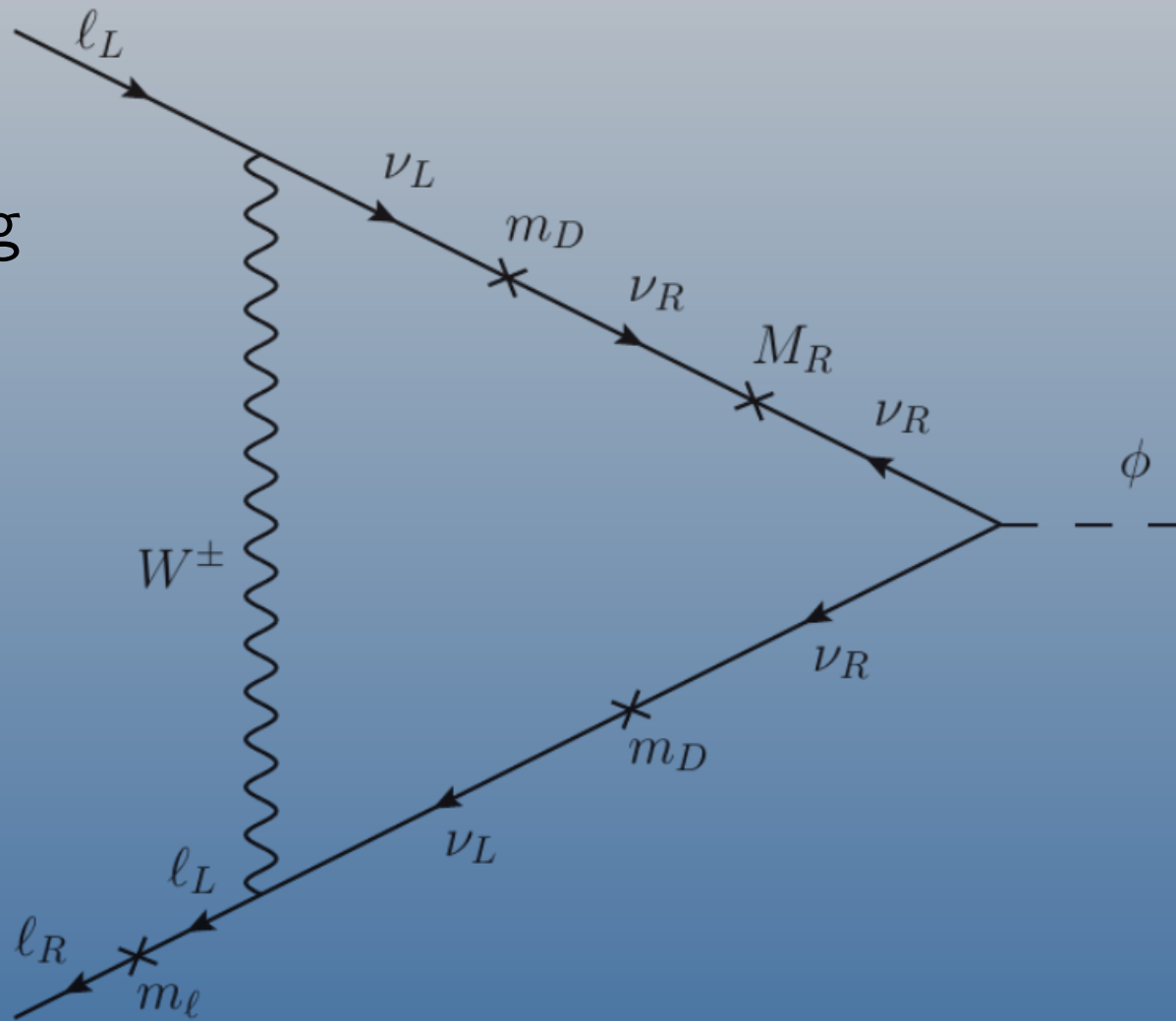
does not couple to normal matter directly.

But, ...



Why $\frac{G_F}{16\pi^2} m_\nu m_\ell$?

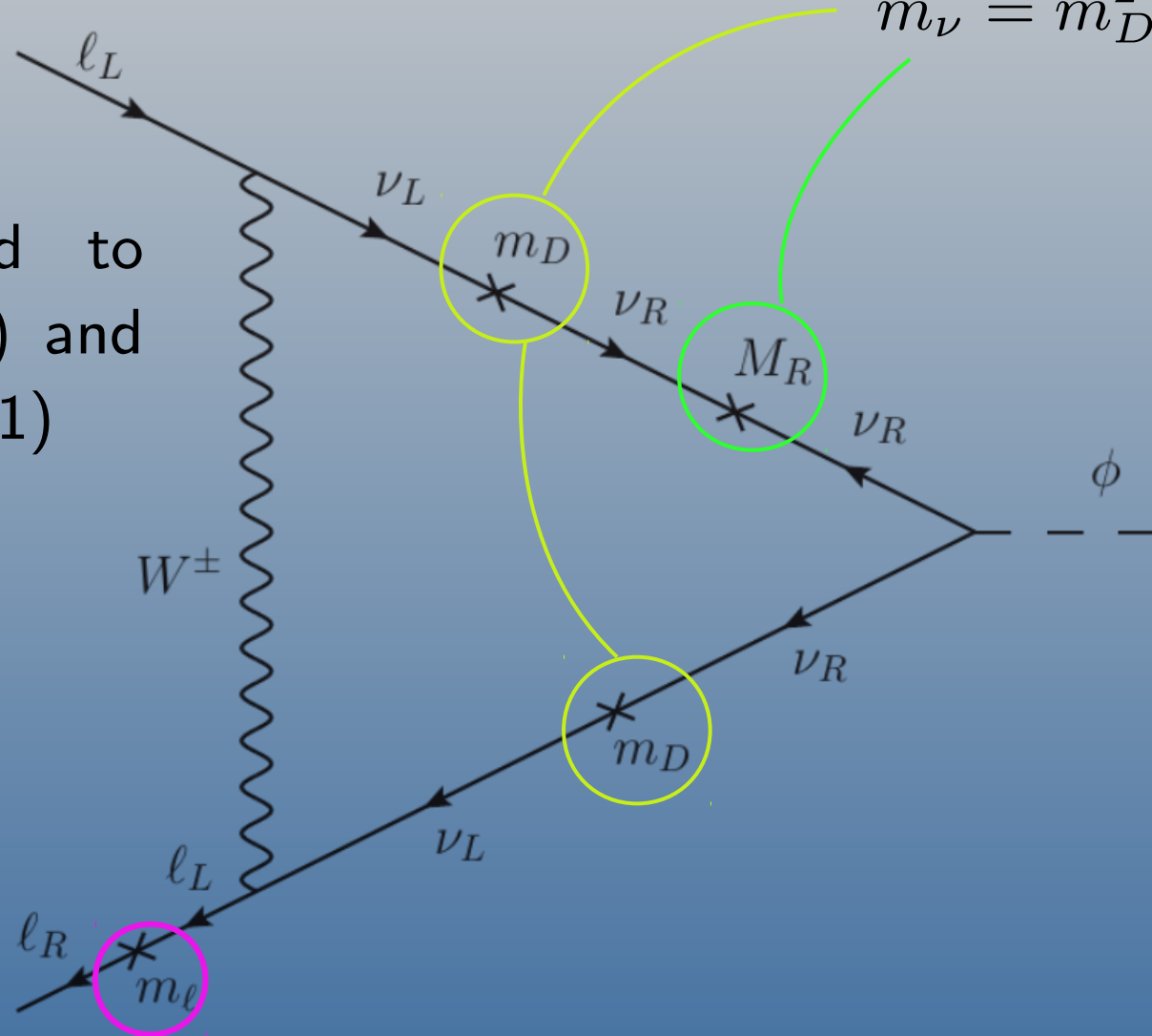
Because of
chirality flipping



Why $\frac{G_F}{16\pi^2} m_\nu m_\ell$?

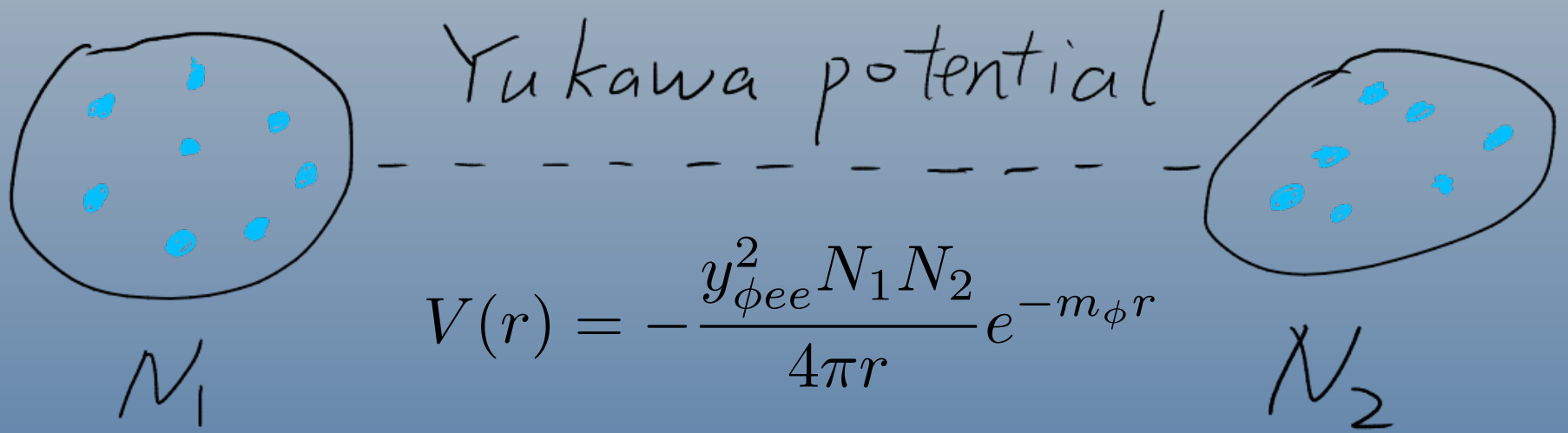
Because we need to flip chirality ($\times 3$) and lepton number ($\times 1$)

Seesaw formula:
 $m_\nu = m_D^2 / M_R$



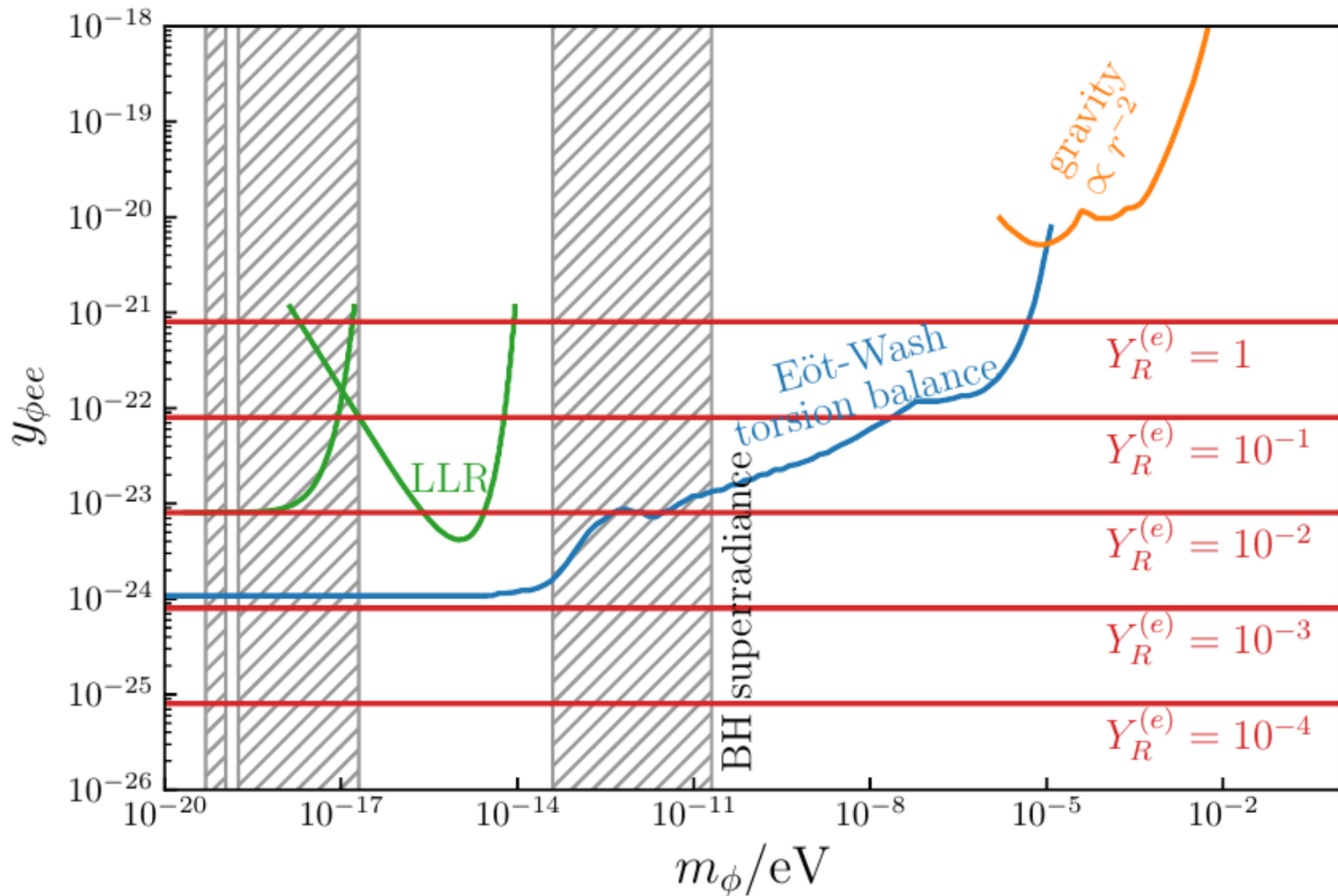
$$\mathcal{L} = \frac{y_R}{2} \nu_R \nu_R \phi + m_D \nu_L \nu_R + \frac{M_R}{2} \nu_R \nu_R + m_\ell \ell_L \ell_R + \dots$$

Now let's compare it to experiments



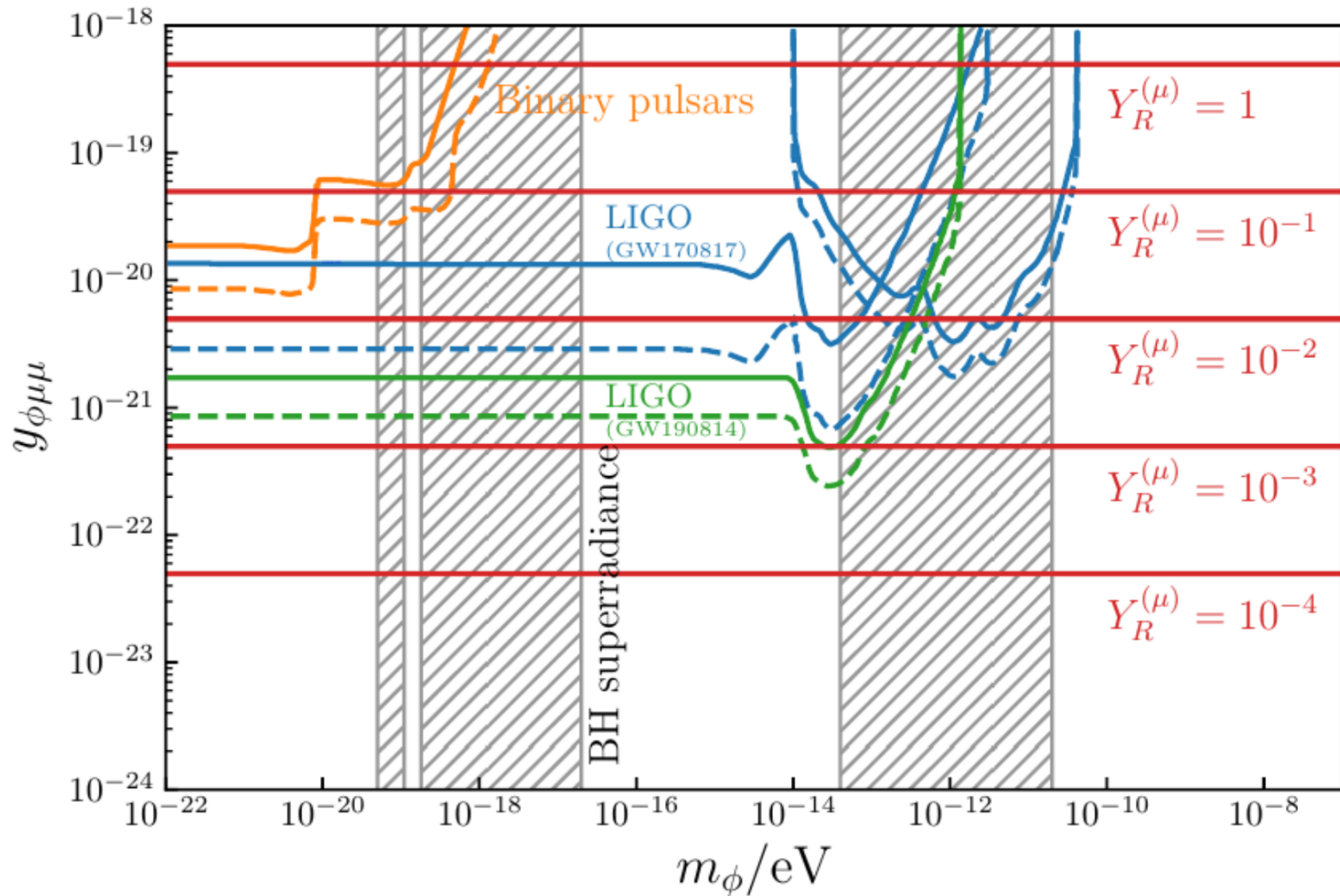
$$y_{\phi ee} = \frac{3G_F m_e Y_R^{(e)} m_\nu^{(e)}}{16\sqrt{2}\pi^2} \approx 8.0 \times 10^{-22} Y_R^{(e)} \left(\frac{m_\nu^{(e)}}{0.01 \text{ eV}} \right)$$

In normal matter, ϕ only couples to electrons



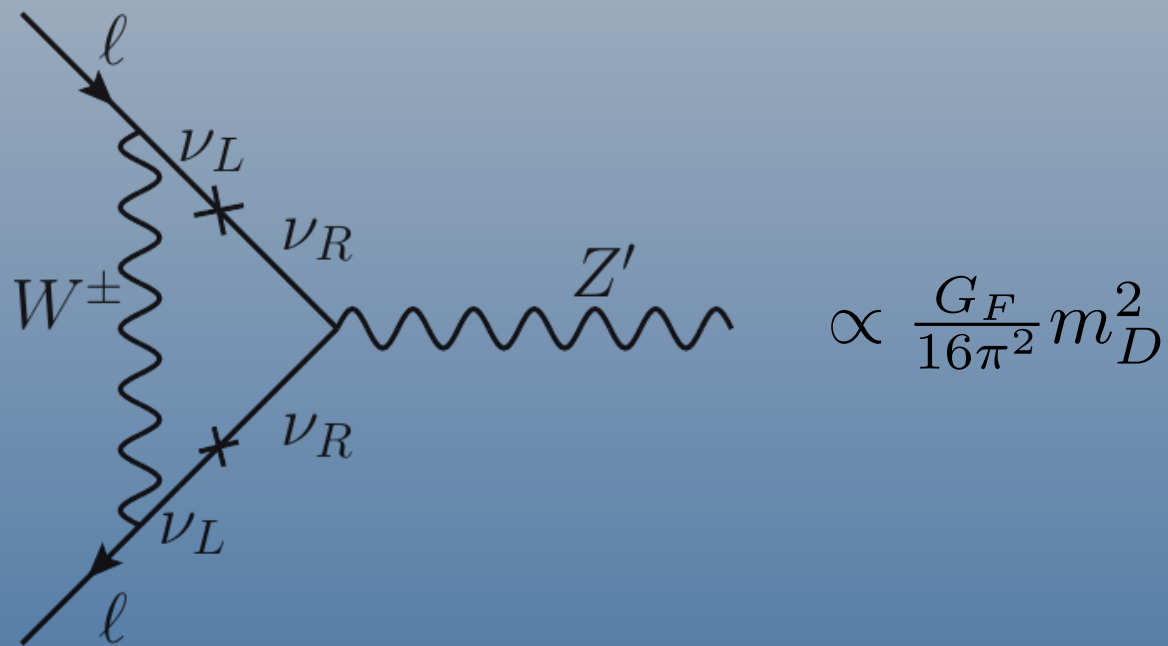
In neutron stars, we can probe the muonic force because

$$F^{(\mu)} \sim (m_{\mu}/m_e)^2 F^{(e)} \gg F^{(e)}$$

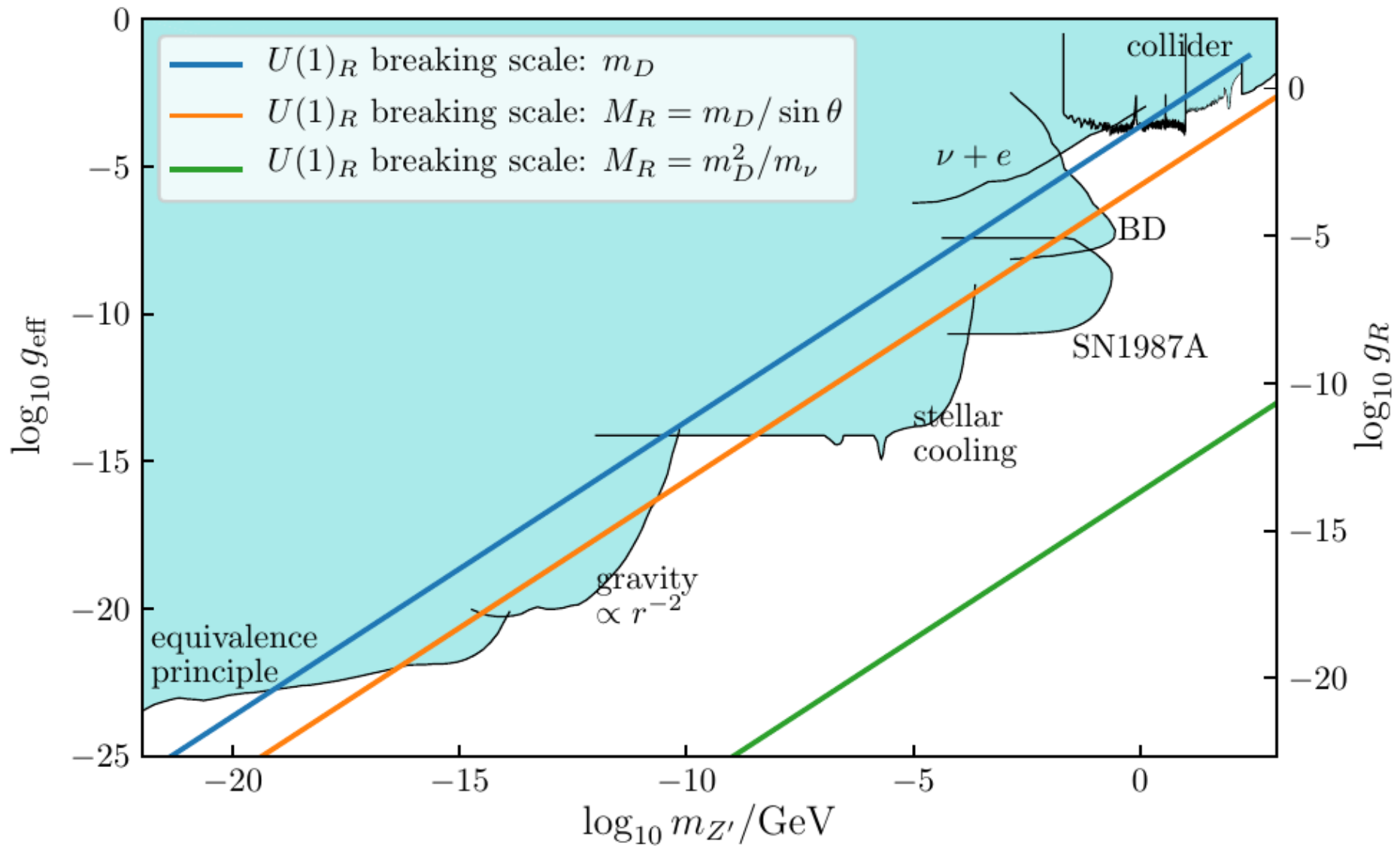


What if the dark force is mediated by a vector?

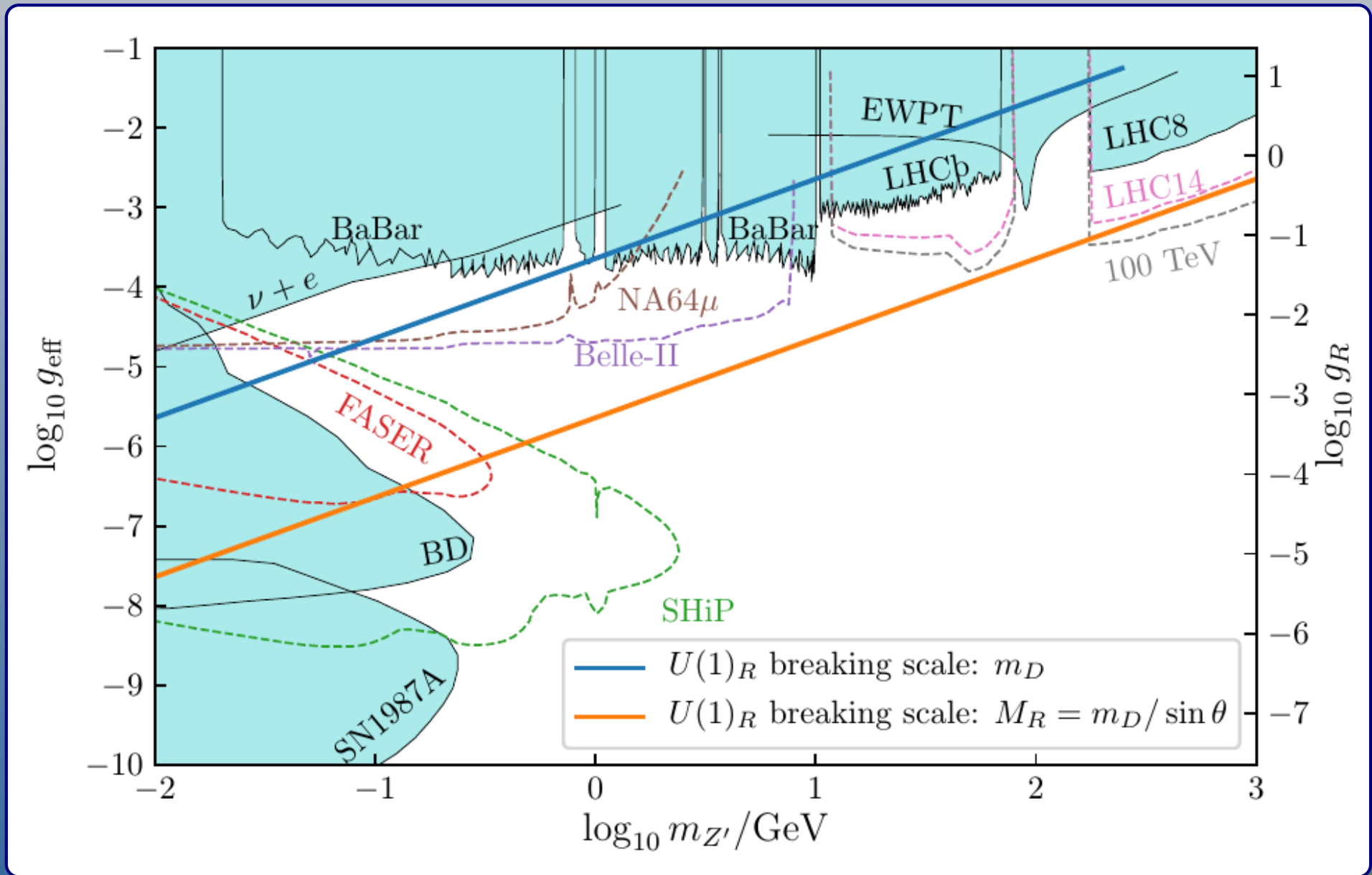
$$\mathcal{L} \supset \nu_R \nu_R \phi \rightarrow \mathcal{L} \supset \bar{\nu}_R \not{Z}' \nu_R$$



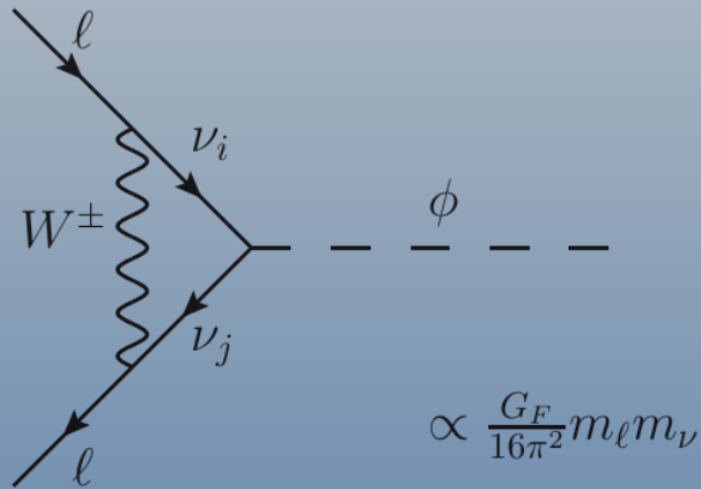
Answer: coupling no longer suppressed by m_ν



ν_R -philic dark photon in future experiments

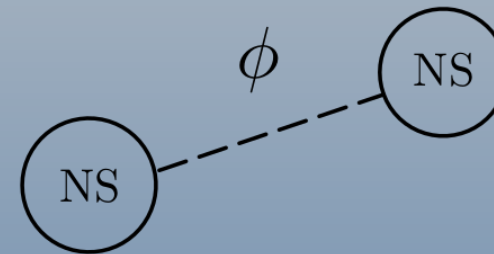


Summary



$$\ell = e: y \sim 10^{-21}$$

$$\ell = \mu: y \sim 10^{-19}$$



GW170817

$$y \sim [10^{-19}, 10^{-20}]$$

Note: in NS, $n_\mu \sim n_e$,
since $m_\mu \gg m_e$, μ - μ force \gg e - e force

Backup

Unitarity gauge ?

R_ξ gauges ?

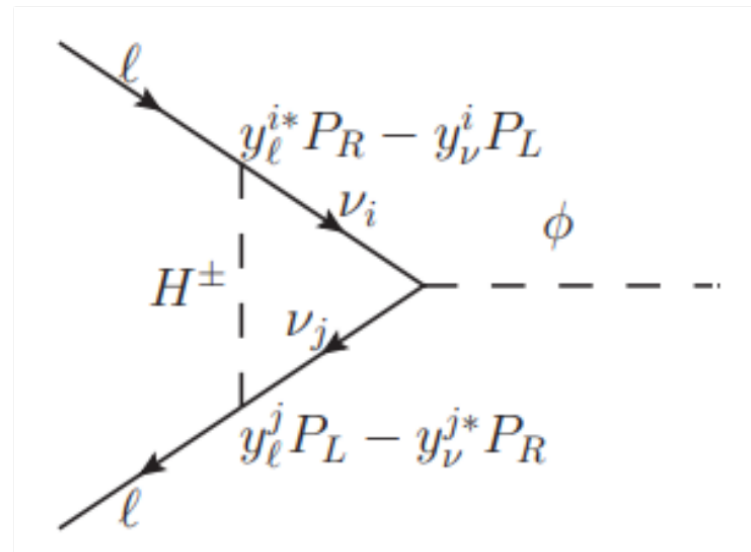
Feynman-'t Hooft gauge?

$$\Delta_{\mu\nu}^W(k) = -i \frac{g_{\mu\nu} - k_\mu k_\nu / m_W^2}{k^2 - m_W^2} - i \frac{k_\mu k_\nu / m_W^2}{k^2 - \xi m_W^2},$$



$k \rightarrow \infty$

vs $\xi \rightarrow \infty$



$$i\mathcal{M}_W = i \frac{m_\ell G^{ij}}{256\pi^2 m_W^2} [F_1(m_i, m_j) + F_2(m_i, m_j)] \overline{u(p_2)} u(p_1)$$

$$F_1 = 6 \left(\frac{1}{\epsilon} + \log \frac{\mu^2}{m_W^2} \right) - 2 \frac{m_j^2 + 2m_W^2}{m_j^2 - m_W^2} \log \frac{m_j^2}{m_W^2} \\ + \frac{5m_i^2 m_j^2 - 5m_i^2 m_W^2 - 5m_j^2 m_W^2 + 11m_W^4}{(m_i^2 - m_W^2)(m_j^2 - m_W^2)} \\ - \frac{2(m_i^2 m_j^2 m_W^2 + m_i^2 m_j^4 - 2m_i^2 m_W^4 - 7m_j^4 m_W^2 + 2m_j^2 m_W^4 + 2m_j^6) \log \frac{m_i^2}{m_j^2}}{(m_i^2 - m_j^2)(m_j^2 - m_W^2)^2} \\ - \frac{2m_W^4 (17m_i^2 m_j^2 - 10m_i^2 m_W^2 + 5m_i^4 - 7m_j^2 m_W^2 + 2m_j^4 + 2m_W^4) \log \frac{m_i^2}{m_W^2}}{(m_i^2 - m_W^2)^2 (m_j^2 - m_W^2)^2} \\ - \frac{2m_i^2 m_j^2 (2m_i^2 m_j^2 - 7m_i^2 m_W^2 - 4m_j^2 m_W^2) \log \frac{m_j^2}{m_W^2}}{(m_i^2 - m_W^2)^2 (m_j^2 - m_W^2)^2},$$

$$F_2 = -6 \left(\frac{1}{\epsilon} + \log \frac{\mu^2}{\xi m_W^2} \right) + 2 \frac{m_j^2}{m_j^2 - \xi m_W^2} \log \frac{m_j^2}{\xi m_W^2} \\ - \frac{5m_i^2 (m_j^2 - \xi m_W^2) + \xi m_W^2 (7\xi m_W^2 - 5m_j^2)}{(m_i^2 - \xi m_W^2)(m_j^2 - \xi m_W^2)} \\ + \frac{2m_j^2 [m_i^2 (m_j^2 - \xi m_W^2) - 3\xi m_j^2 m_W^2 + 2m_j^4] \log \frac{m_i^2}{m_j^2}}{(m_i^2 - m_j^2)(m_j^2 - \xi m_W^2)^2} \\ - \frac{2\xi^2 m_j^2 m_W^4 (2m_j^2 - 3\xi m_W^2) \log \frac{m_i^2}{\xi m_W^2}}{(m_i^2 - \xi m_W^2)^2 (m_j^2 - \xi m_W^2)^2} \\ + \frac{2\xi m_i^2 m_W^2 (9\xi m_j^2 m_W^2 - 4m_j^4 - 4\xi^2 m_W^4) \log \frac{m_i^2}{\xi m_W^2}}{(m_i^2 - \xi m_W^2)^2 (m_j^2 - \xi m_W^2)^2} \\ + \frac{2m_i^4 (2m_j^4 - 5\xi m_j^2 m_W^2 + 3\xi^2 m_W^4) \log \frac{m_i^2}{\xi m_W^2}}{(m_i^2 - \xi m_W^2)^2 (m_j^2 - \xi m_W^2)^2}.$$

UV divergence ?

$$F_1 = \frac{6}{\epsilon} + \dots$$

$$F_2 = -\frac{6}{\epsilon} + \dots$$

Cancellation of ξ - dependence (in the limit of large ξ)

$$i\mathcal{M}_W = i \frac{m_\ell G^{ij}}{256\pi^2 m_W^2} [F_1(m_i, m_j) + F_2(m_i, m_j)] \overline{u(p_2)} u(p_1)$$

Each single diagram
is gauge dependent!

$$\xi \rightarrow \infty$$

$$F_2 \rightarrow 6 \log \xi$$

But,

$$\sum_{ij} G^{ij} \log \xi = 0$$

$$F_2 = -6 \left(\frac{1}{\epsilon} + \log \frac{\mu^2}{\xi m_W^2} \right) + 2 \frac{m_j^2}{m_j^2 - \xi m_W^2} \log \frac{m_j^2}{\xi m_W^2}$$

$$- \frac{5m_i^2 (m_j^2 - \xi m_W^2) + \xi m_W^2 (7\xi m_W^2 - 5m_j^2)}{(m_i^2 - \xi m_W^2) (m_j^2 - \xi m_W^2)}$$

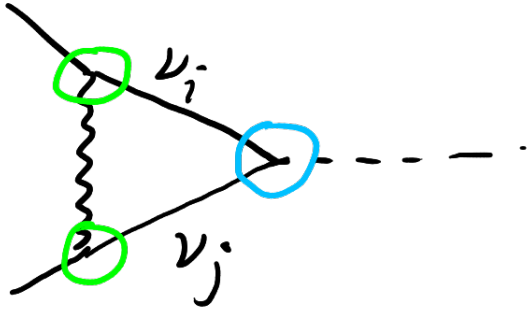
$$+ \frac{2m_j^2 \left[m_i^2 (m_j^2 - \xi m_W^2) - 3\xi m_j^2 m_W^2 + 2m_j^4 \right] \log \frac{m_i^2}{m_j^2}}{(m_i^2 - m_j^2) (m_j^2 - \xi m_W^2)^2}$$

$$- \frac{2\xi^2 m_j^2 m_W^4 (2m_j^2 - 3\xi m_W^2) \log \frac{m_i^2}{\xi m_W^2}}{(m_i^2 - \xi m_W^2)^2 (m_j^2 - \xi m_W^2)^2}$$

$$+ \frac{2\xi m_i^2 m_W^2 (9\xi m_j^2 m_W^2 - 4m_j^4 - 4\xi^2 m_W^4) \log \frac{m_i^2}{\xi m_W^2}}{(m_i^2 - \xi m_W^2)^2 (m_j^2 - \xi m_W^2)^2}$$

$$+ \frac{2m_i^4 (2m_j^4 - 5\xi m_j^2 m_W^2 + 3\xi^2 m_W^4) \log \frac{m_i^2}{\xi m_W^2}}{(m_i^2 - \xi m_W^2)^2 (m_j^2 - \xi m_W^2)^2}.$$

$$G^{ij} = \text{O} \times \text{O} \times \text{O}$$



Chiral basis (original basis)

$$\mathcal{L} \supset \frac{g}{2c_W} Z_\mu \nu_L^\dagger \bar{\sigma}^\mu \nu_L + \left[\frac{g}{\sqrt{2}} W_\mu^- \ell_L^\dagger \bar{\sigma}^\mu \nu_L - y_\nu H^+ \ell_{L\nu R} + y_\ell H^- \nu_{L\ell R} + \frac{y_R}{2} \nu_R \nu_R \phi + \text{h.c.} \right],$$



$$\begin{pmatrix} \nu_L \\ \nu_R \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_4 \end{pmatrix}, \quad U^T \begin{pmatrix} 0 & m_D \\ m_D & M_R \end{pmatrix} U = \begin{pmatrix} m_1 & \\ & m_4 \end{pmatrix}$$

mass basis

$$\mathcal{L} \supset g_Z^{ij} Z_\mu \nu_i^\dagger \bar{\sigma}^\mu \nu_j + \left[\underset{\downarrow}{g_W^i} W_\mu^- \ell_L^\dagger \bar{\sigma}^\mu \nu_i - y_\nu^i H^+ \ell_{L\nu i} + y_\ell^i H^- \nu_{i\ell R} + \frac{y_R^{ij}}{2} \nu_i \nu_j \phi + \text{h.c.} \right].$$

$$\begin{pmatrix} 1 \\ 0 \end{pmatrix} \rightarrow U^T \begin{pmatrix} 1 \\ 0 \end{pmatrix}$$

$$U^T \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} U$$

$$\begin{matrix} \text{|||} \\ g_W \end{matrix}$$

$$\begin{matrix} \text{|||} \\ Y_R \end{matrix}$$

$$g_W^+ Y_R = 0!$$

$$G^{ij} \equiv \underbrace{g_W^{i*} g_W^j (m_j y_R^{ij} + m_i y_R^{ij*})}$$

$$\downarrow \sum_{i,j} G_{ij}$$

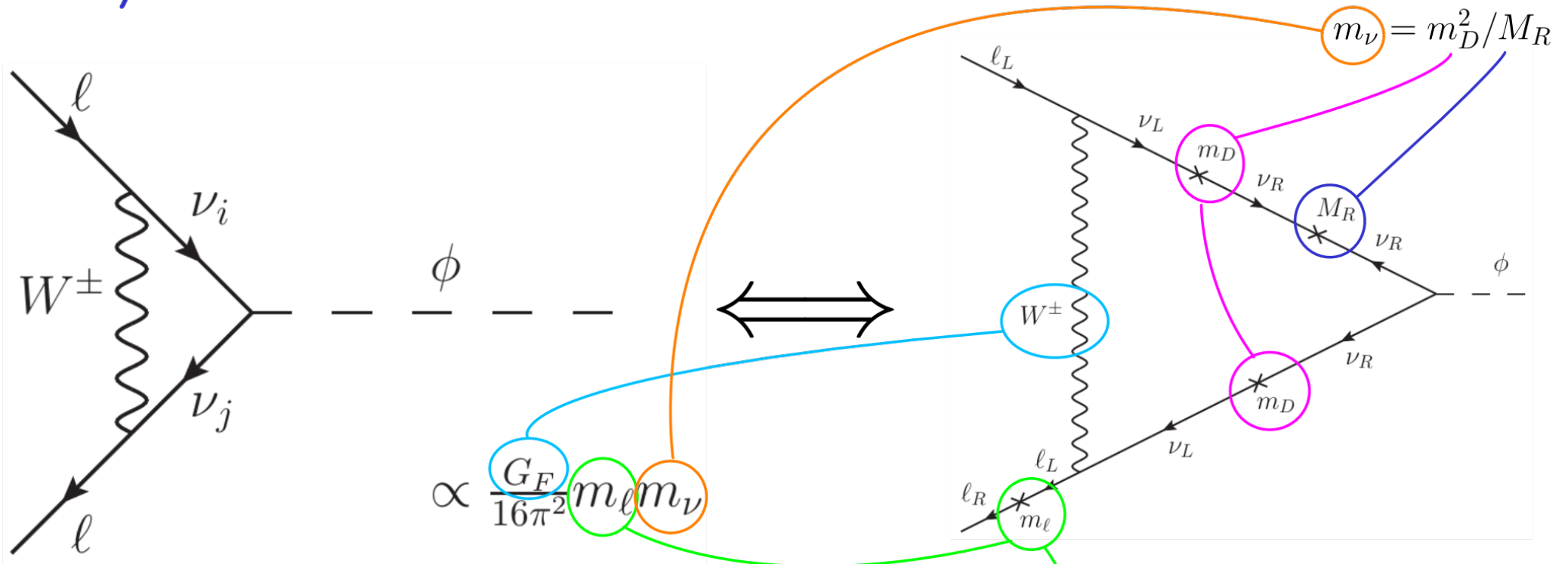
g_W^+	Y_R	m	g_W
	m		

An implication of $\sum_{ij} G_{ij} = 0$

even if $\frac{1}{\epsilon}$ didn't cancel out in $F_1 + F_2$

it would disappear after \sum_{ij}

Why this result?



$$\phi \bar{\psi}_e \psi_e = \phi l_L l_R + \text{h.c.}$$

$$\mathcal{L} = \frac{y_R}{2} \nu_R \nu_R \phi + m_D \nu_L \nu_R + \frac{M_R}{2} \nu_R \nu_R + m_\ell l_L l_R + \dots$$