Marco Drewes, Université catholique de Louvain

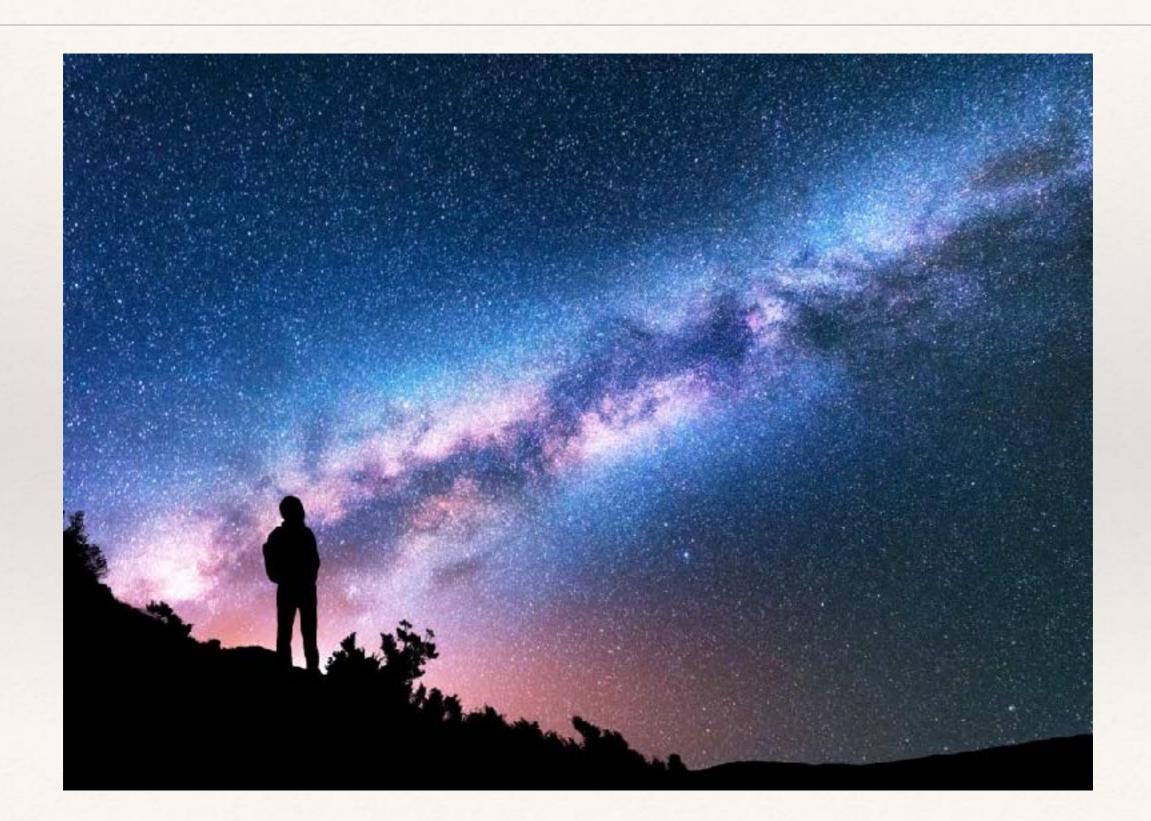
Introduction to Cosmology

11. 09. - 13. 09. 2019

BND Summer School

Spa, Belgium

Lecture I Historical Overview and Basics of GR



And the lord stood upon Tiamat's hinder parts, And with his merciless club he smashed her skull. He cut through the channels of her blood, And he made the North wind bear it away into secret places.

[...]

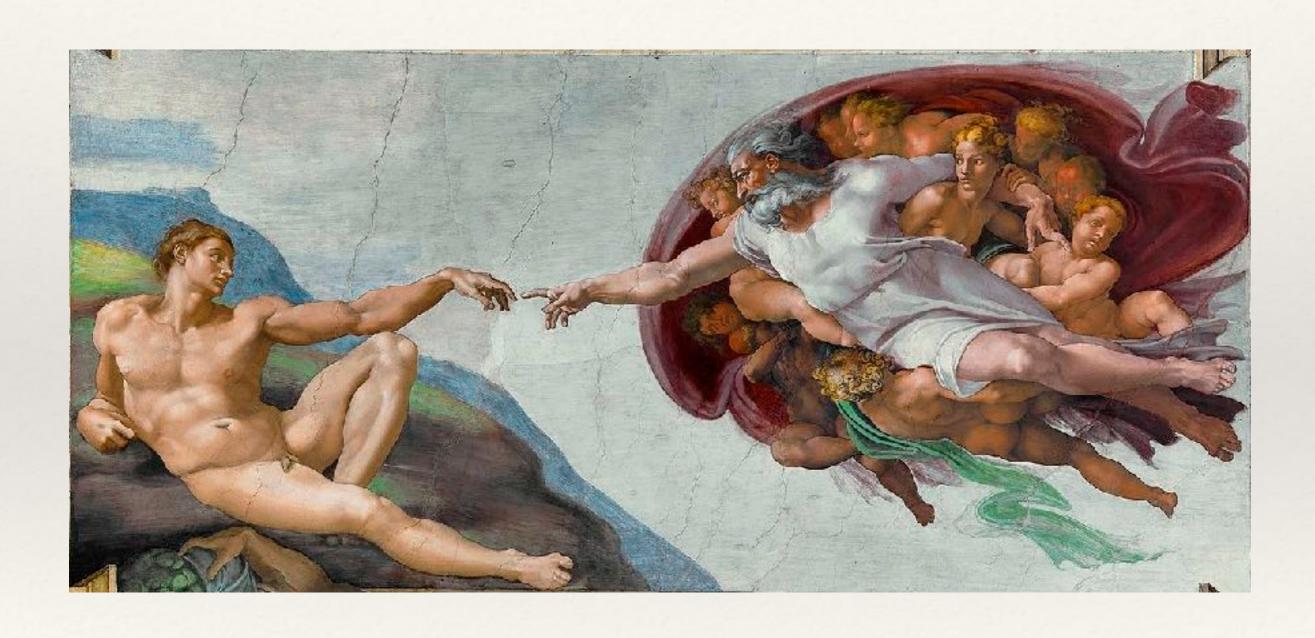
Slicing Tiamat in half, he made from her ribs the vault of heaven and earth.

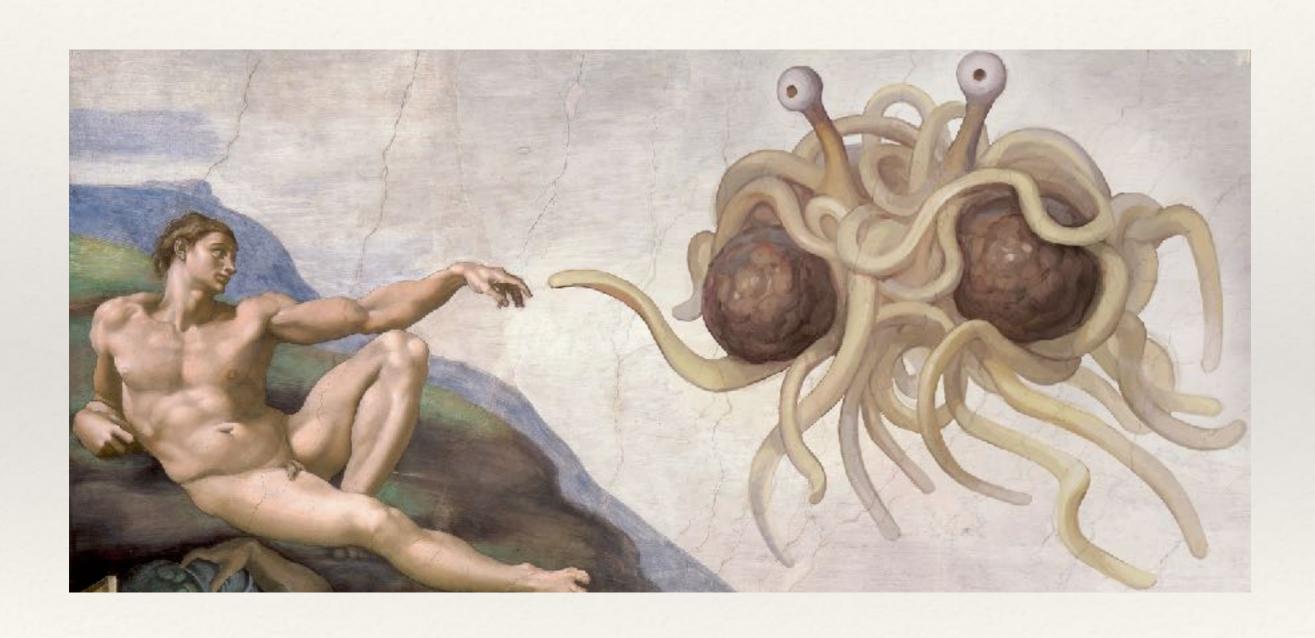
Her weeping eyes became the source of the Tigris and the Euphrates,

her tail became the Milky Way.



Babylonian Enûma Eliš



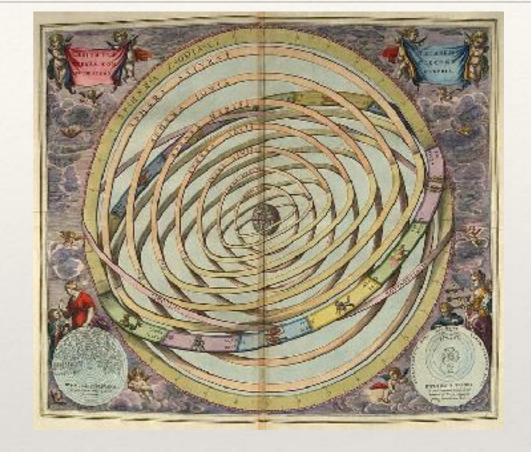


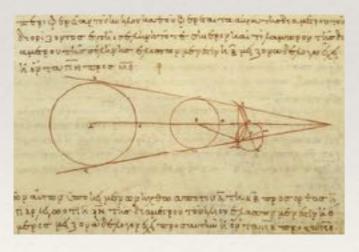
Early Quantitative Models

Aristotle (384 - 322 BC) argued that Earth and Moon are spheres, based on lunar eclipses

Eratosthenes (276-194 BC) computed the radius of the Earth to be 6800 km

...but both assumed geocentric universes





Aristarchus (310 - 230 BC) correctly computed the ratios between the size and distance of the Moon and Sun...

...and computed the distance to the Sun (wrong by a factor 20), concluding that

the Universe is much bigger than Earth, and the Sun is the centre

Mapping the Solar System

In the renaissance age, many astronomers explored the dimensions of the solar system, including **Galileo**, **Coppernicus**, **Brahe**, **Keppler**...

...but a quantitative understanding of the laws that govern it had to wait until **Newton**'s Philosophiae Naturalis Principia Mathematica (1687), which paved the way for

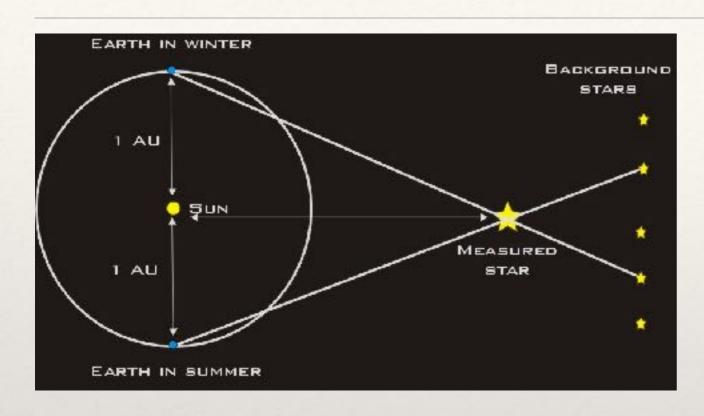


connecting cosmology to fundamental physics!



Around the same time **Rømer** (1644-1710) and **Huygens** (1629-1695) first measured the speech of light, using modulations in the period of Jupiter's moons orbit

To the Stars!



In 1838 **Bessel** first measured the distance to other stars, using the parallax method...

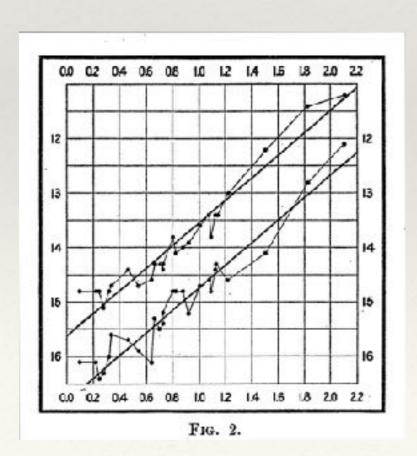
...making the "universe" almost a million times larger!

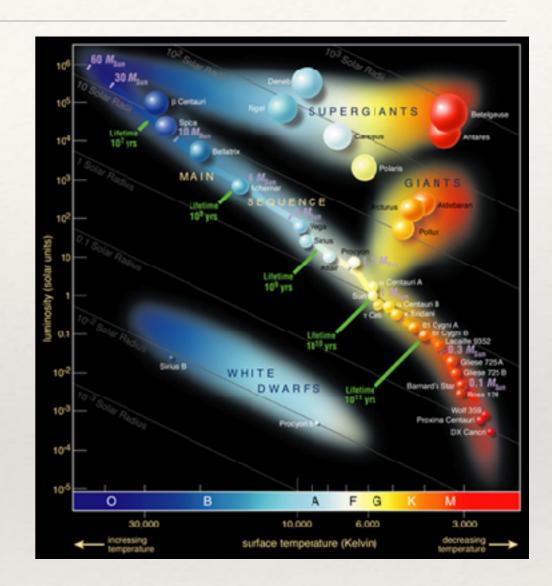
Fun fact: Greek scholars rejected Aristarchus heliocentric model over 2000yrs before because they did not see parallaxes.

To the Stars!

In the early 20th century **Hertzsprung** and **Russel** discovered a relation between colour and luminosity of starts, while **Leavitt** discovered a relation between the period and brightness of variable cepheid stars in 1912.

The cosmic distance ladder was born!

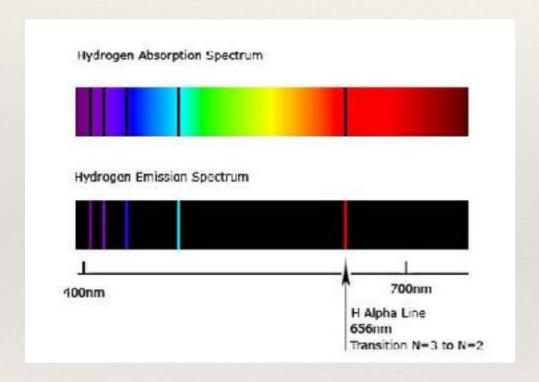


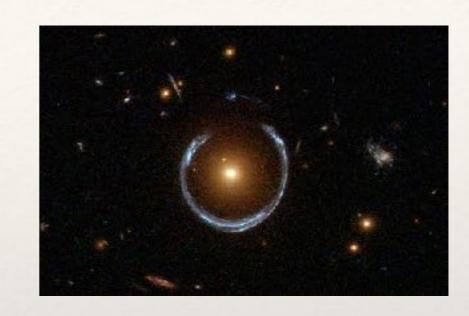


This allows to estimate the distance of far away stars, using distance measurements of nearby ones

The Golden Age of Physics

In 1915 **Einstein** and **Hilbert** developed General Relativity, the mathematical framework to describe the Universe on large scales





The detailed mapping of atomic spectra and their understanding in quantum mechanics proved that

other starts are made of the same atoms that we find on Earth

and

allows to measure their velocity via the Doppler effect

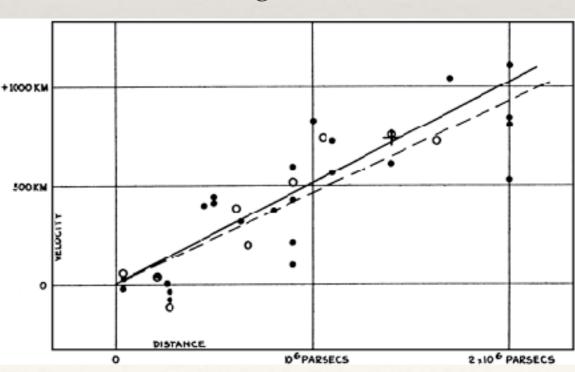
The Big Bang

Between 1922 and 1924 **Friedmann** developed the first cosmological models in General Relativity based on the **cosmological principle**, concluding that the universe expands. **Lemaitre** derived this independently and proposed the "primeval atom", marking the **birth of the Big Bang theory!**

In 1924 Hubble discovered that many "nebulae" are in fact distant galaxies...

making the universe millions of times bigger!

In 1929 he found hat the galaxies move away from us with a "velocity" that is proportional to their distance ("Hubble-Lemaitre law")... ...meaning that we are either the centre of the world of Friedmann and Lemaitre were right!

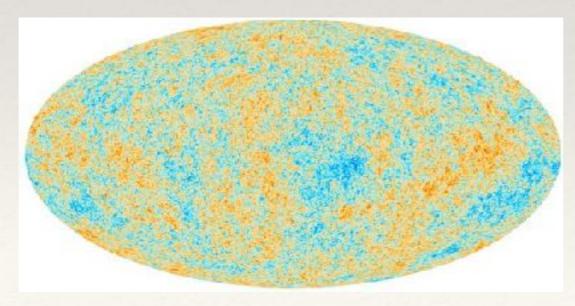


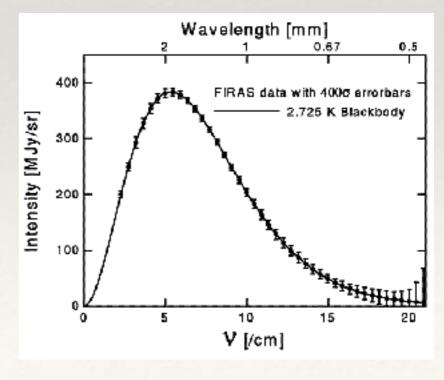
The Primordial Plasma

In the 1940s **Gamov**, **Alpher**, **Herman**, **Bethe** and others explored the implication of the high pressure, density and temperature in the unifant universe predicted by Friedmann and Lemaitre., concluding that

- The temperature was so high that thermonuclear reactions would occur, giving rise to **Big Bang Nucleosynthesis**
- Redshifted photons from the hot plasma should still be present as a

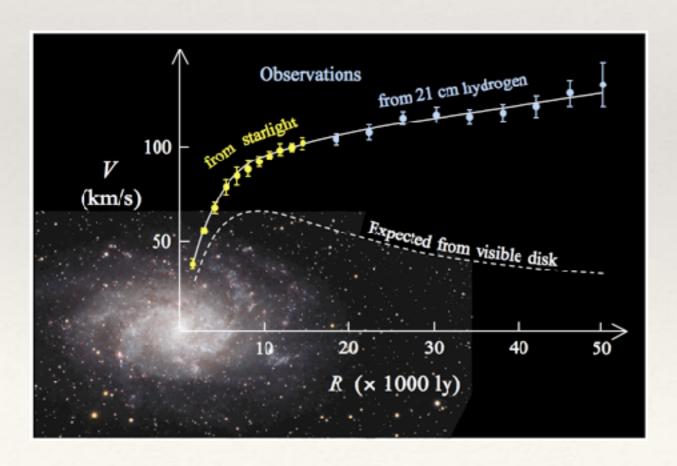
Cosmic Background Radiation





Dark Matter

In 1933 **Zwicky** found that the Coma galaxy cluster cannot be bound together by the gavtity of its visible constituents, suggesting that there is extra **Dark Matter**, but was largely ignored....



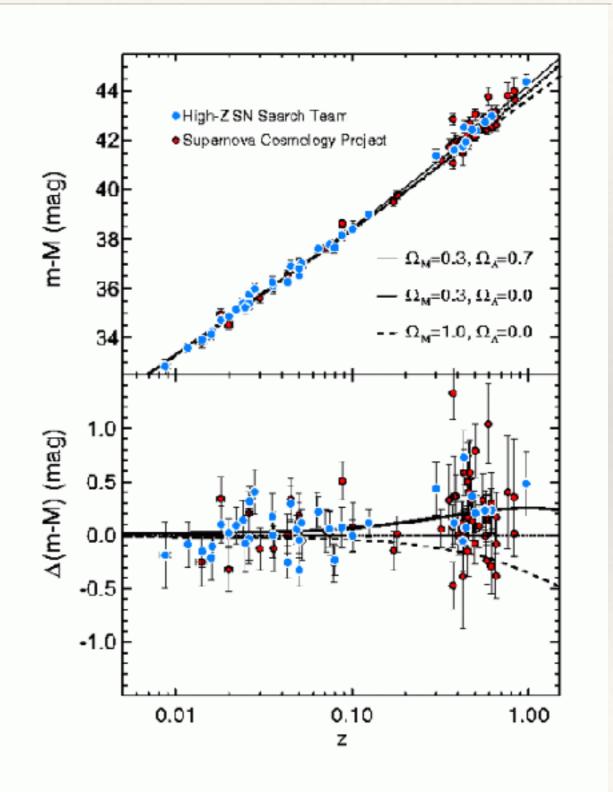


...until **Rubin** provided convincing evidence based on galaxy rotation curves and proved that Dark Matter makes up most of the mass in the universe

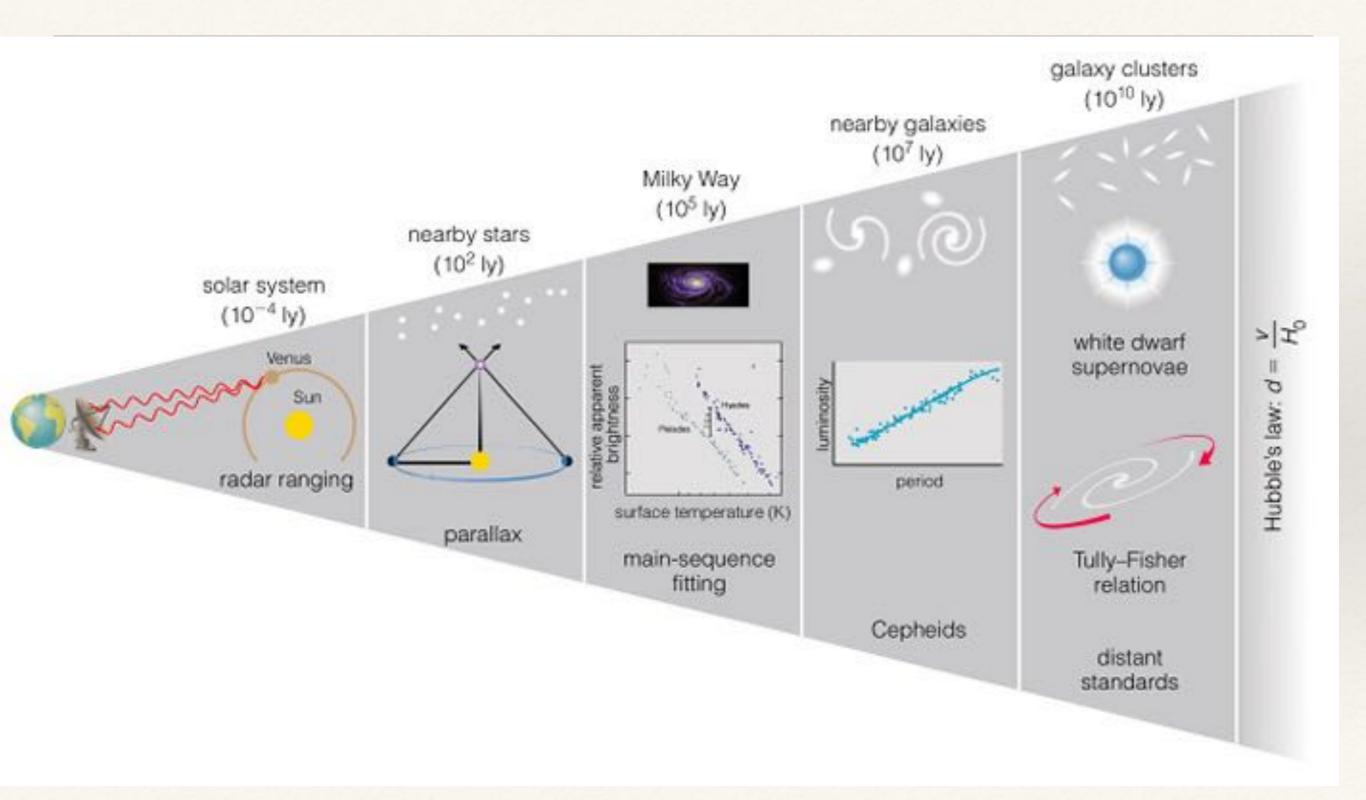
Dark Energy

In 1998 and 1999 two teams of astronomers independently found deviations from Hubble's law in the "wrong direction", indicating that the universe is accelerating!

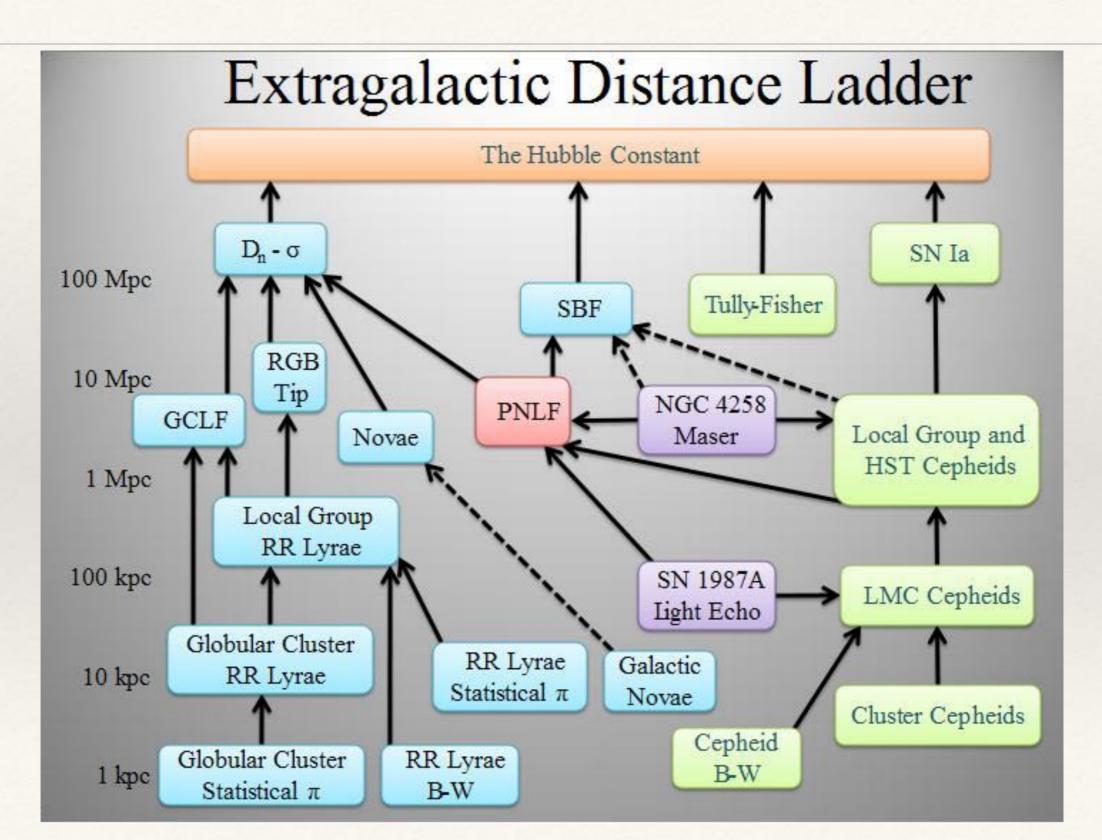




Cosmic Distance Ladder



Actual Cosmic Distance Ladder



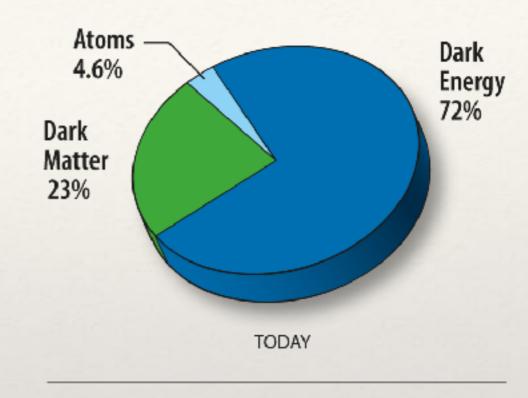
Standard Model of Cosmology

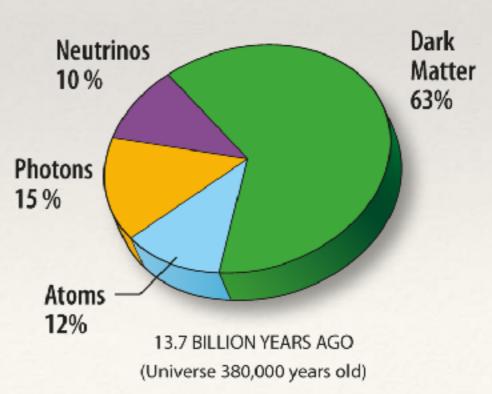
The universe is largely homogeneous and isotropic.

It seems to have vanishing overall spacial curvature.

It is expanding and originated from a high temperature early phase known as "big bang".

The visible constituents are well described by the Standard Model of particle physics, but there is a lot of dark stuff that we do not understand...



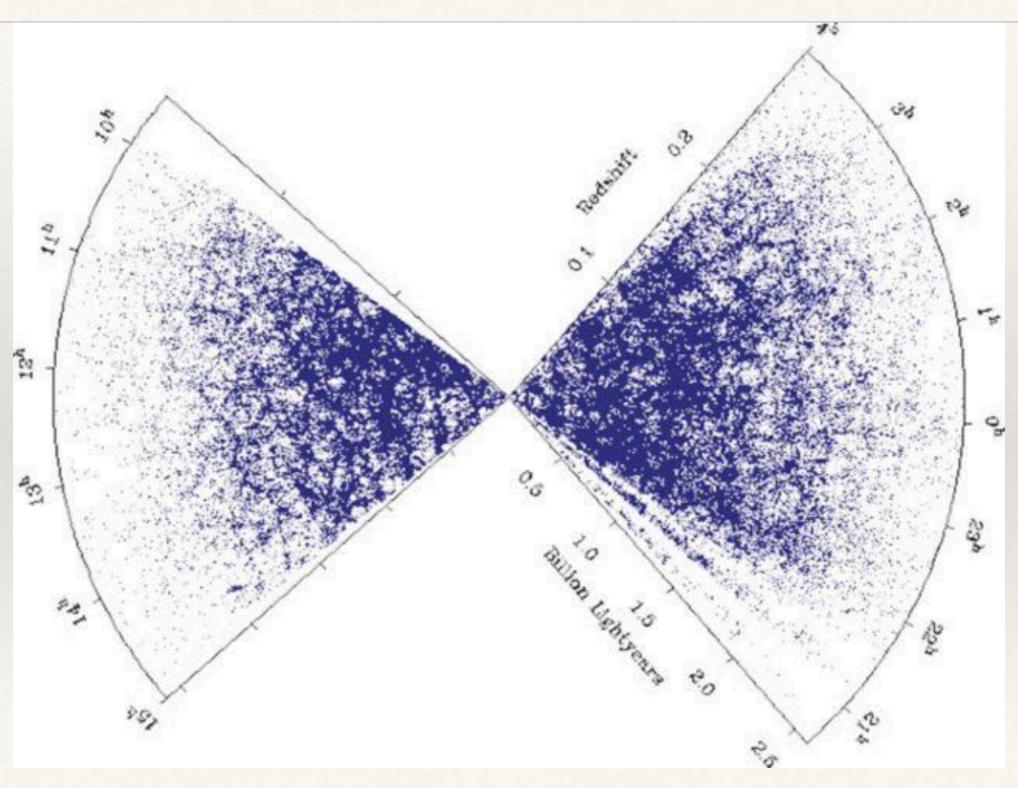


The basics of GR part was based on a combination of different books lecture notes. The main results (geodesic equation, Einstein field equations...) can be found in any textbook of GR

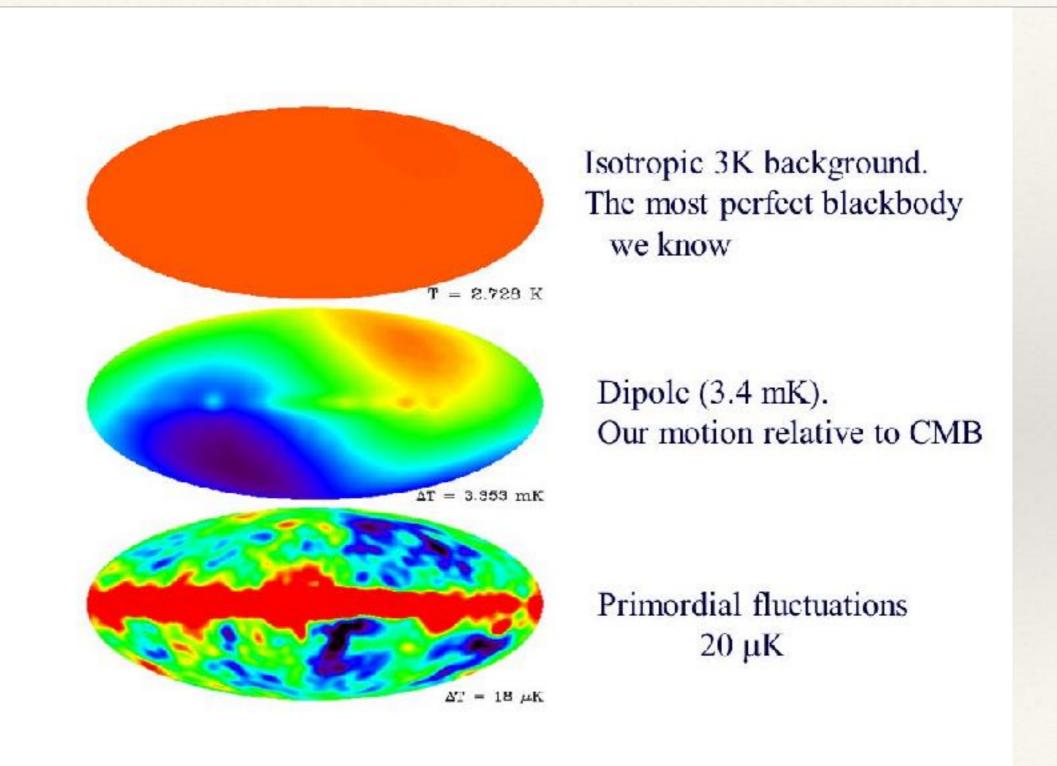
Lecture II The Friedmann - Lemaitre Robertson - Walter Metric

This part was closely following chapter 2 in the book "The Early Universe" by Kolb and Turner.

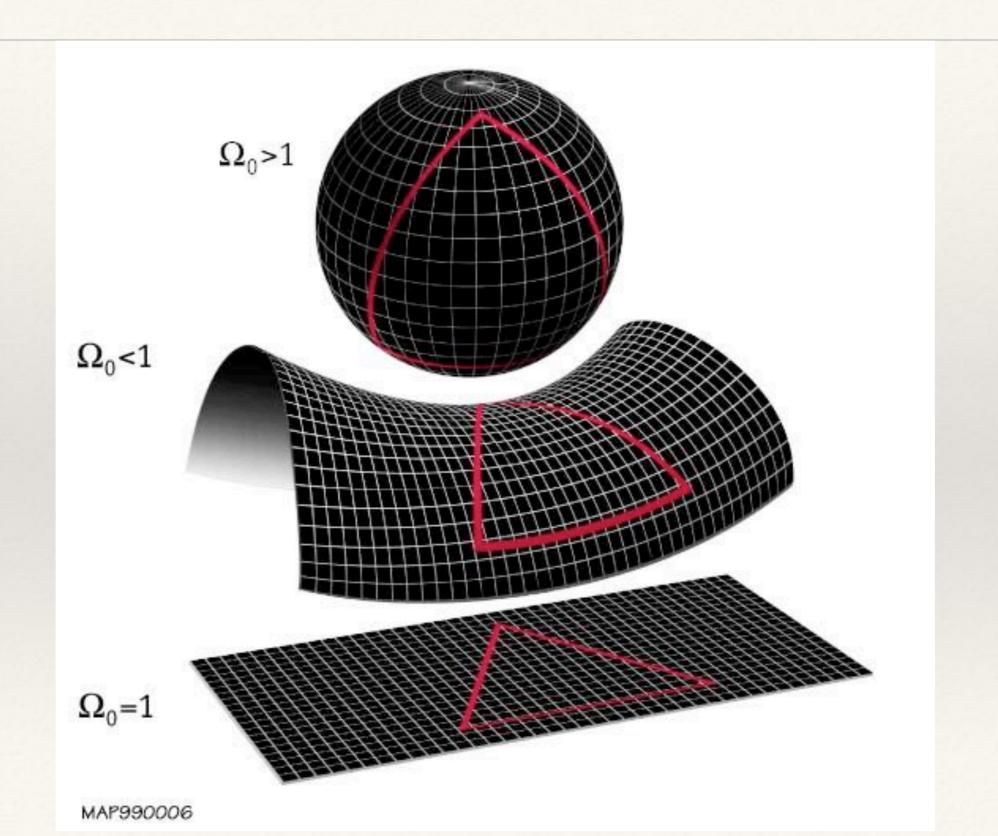
Cosmological Principle Tested



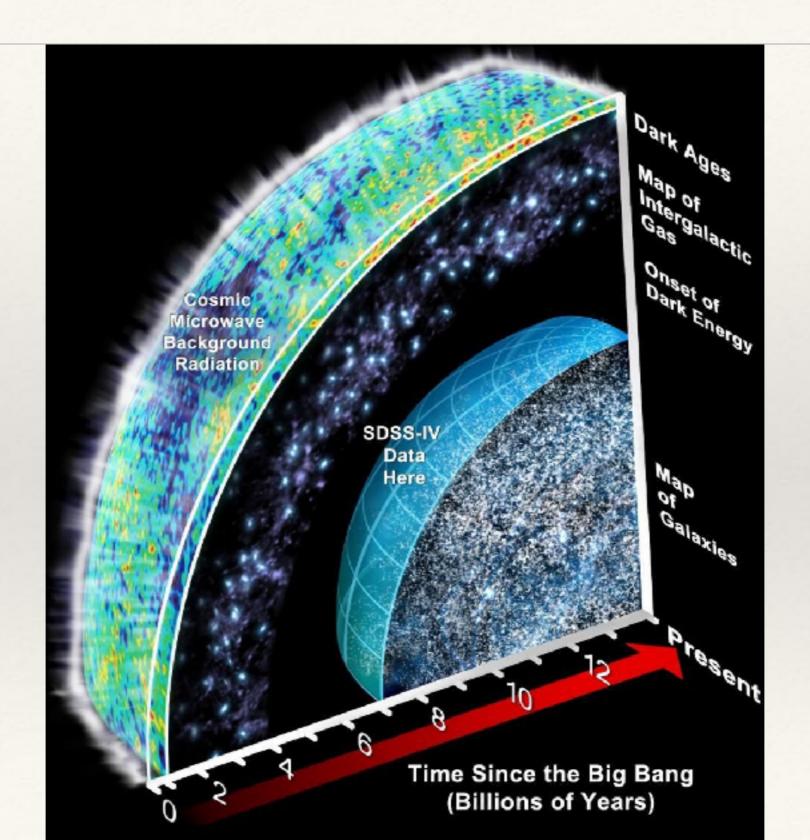
Cosmological Principle Tested



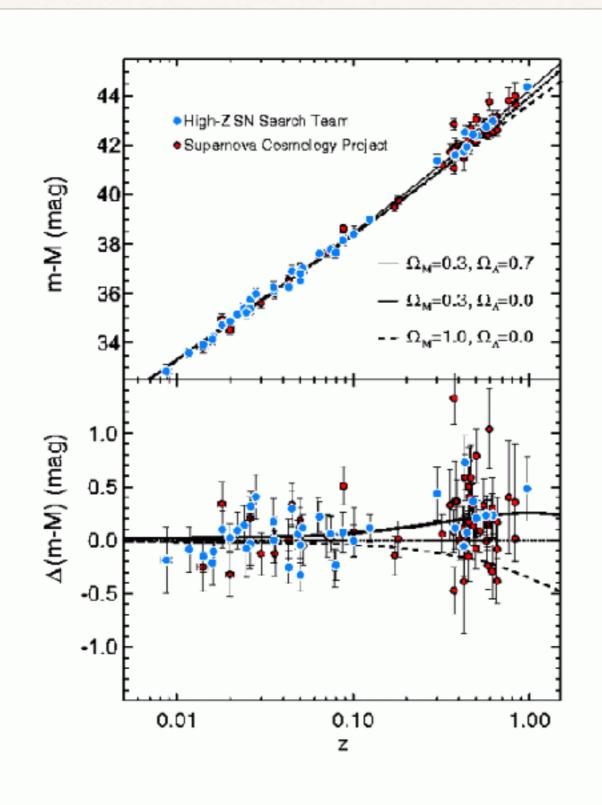
FLRW Universes



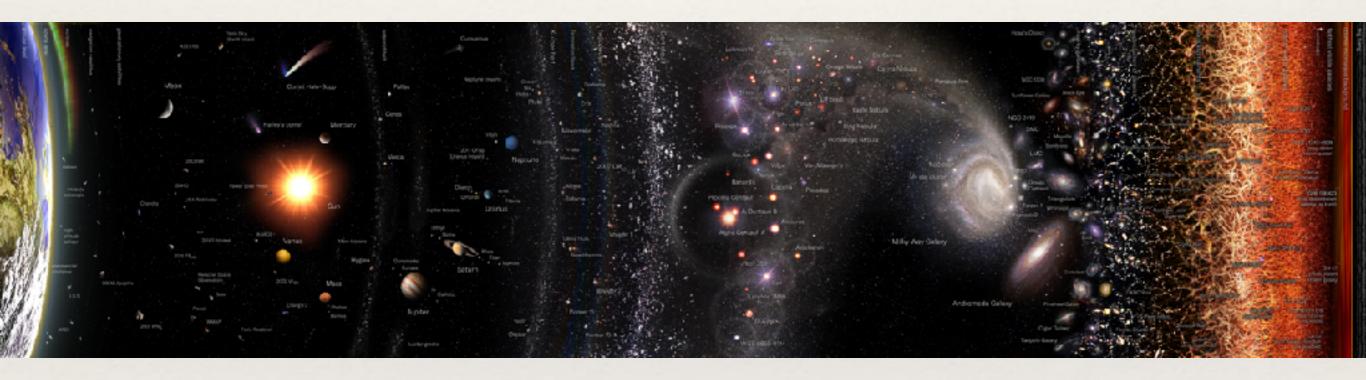
Observable Universe



Redshift - Distance Relation

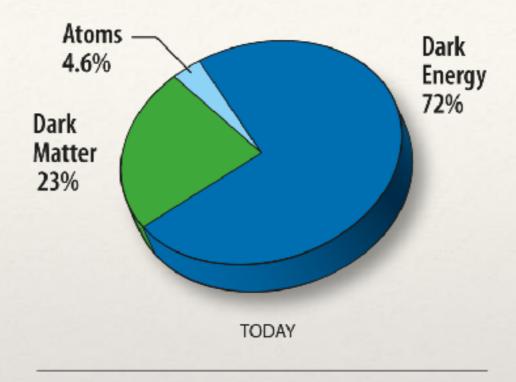


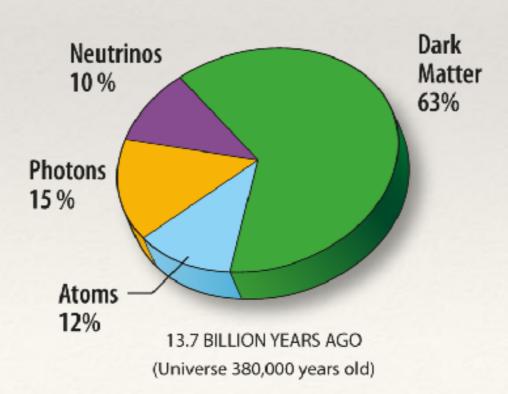
Observable Universe



Size: 46.5 billion light-years

Changing Energy Content



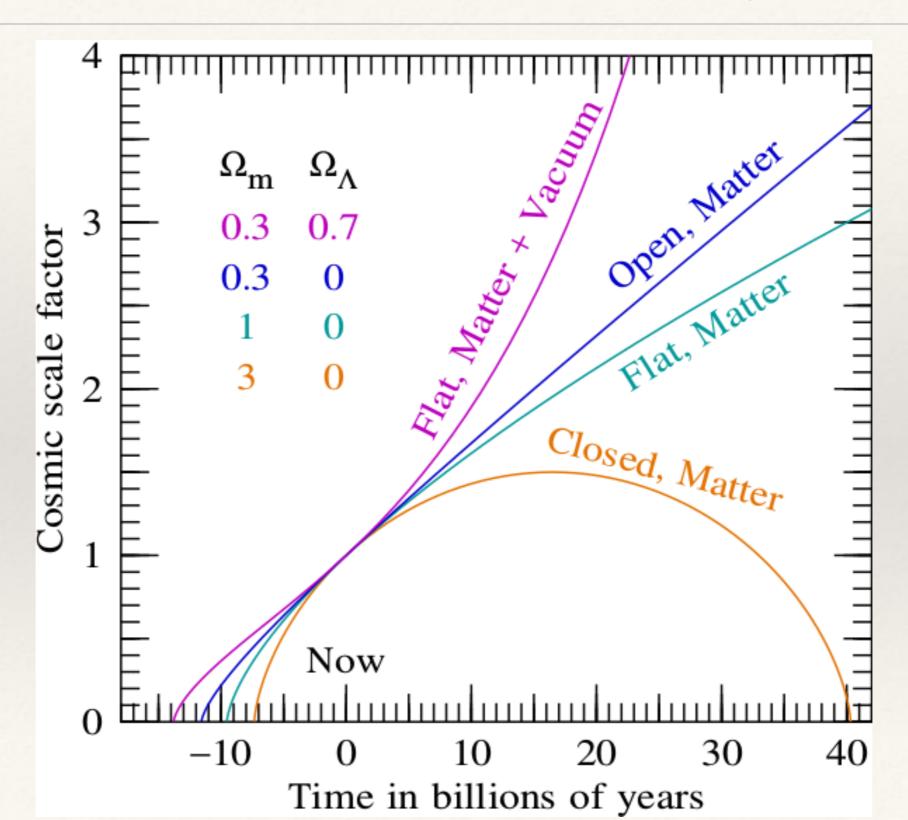


Lecture III Standard Model of Cosmology... ...and its Limits

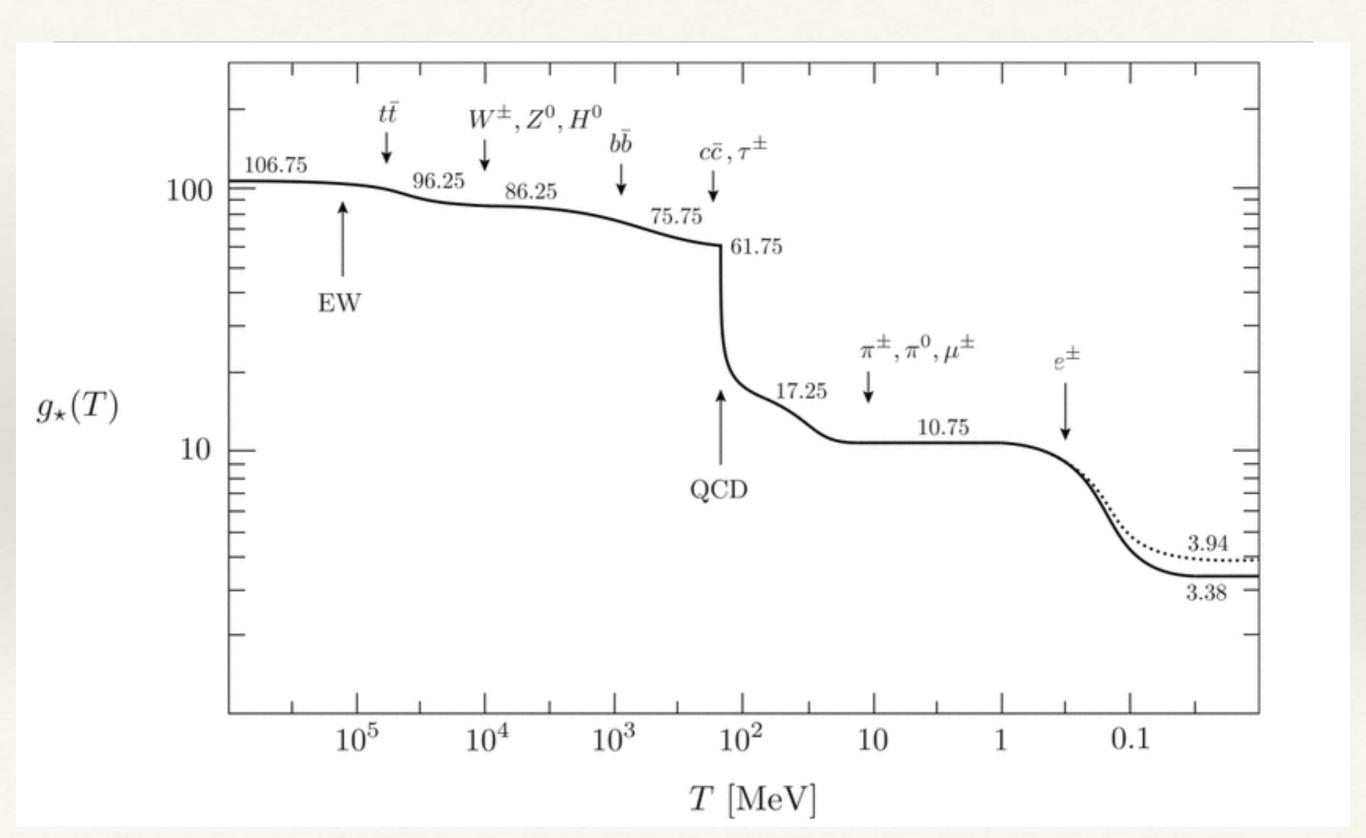
History of the Universe

This part is closely following chapter 3 in the book "The Early Universe" by Kolb and Turner.

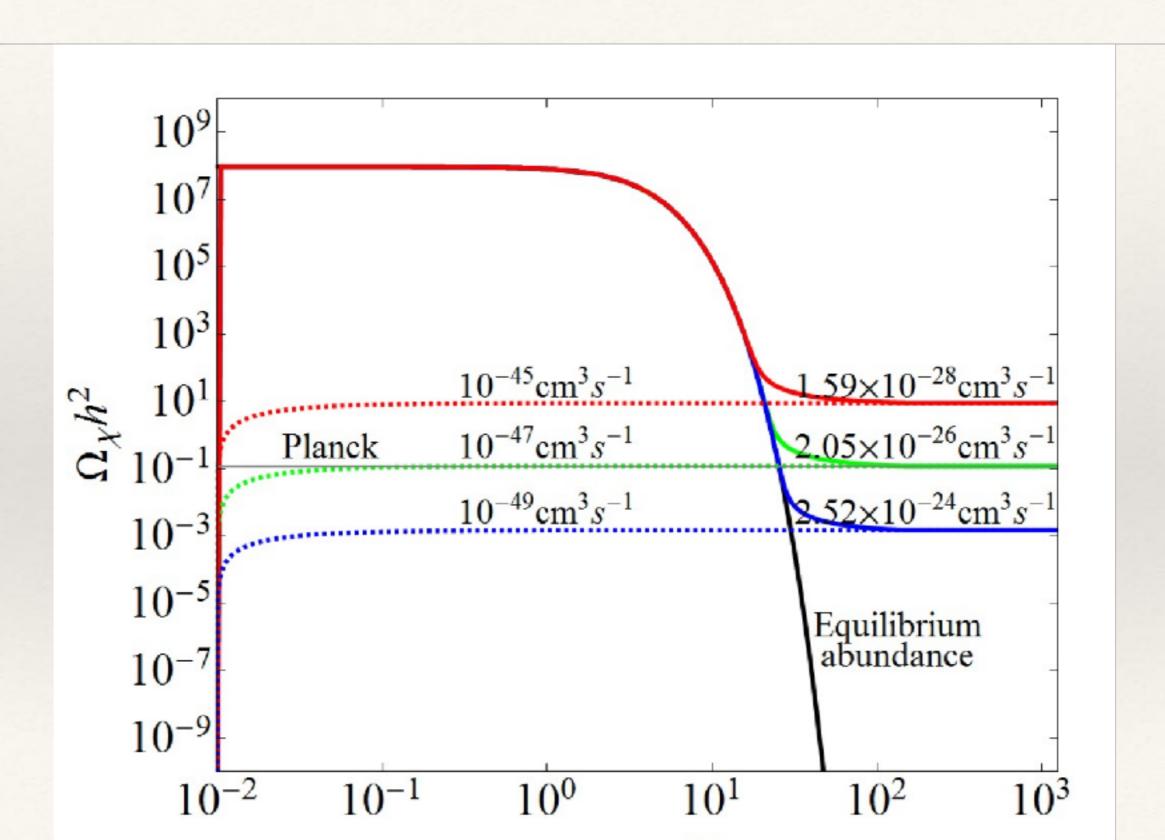
Expansion History

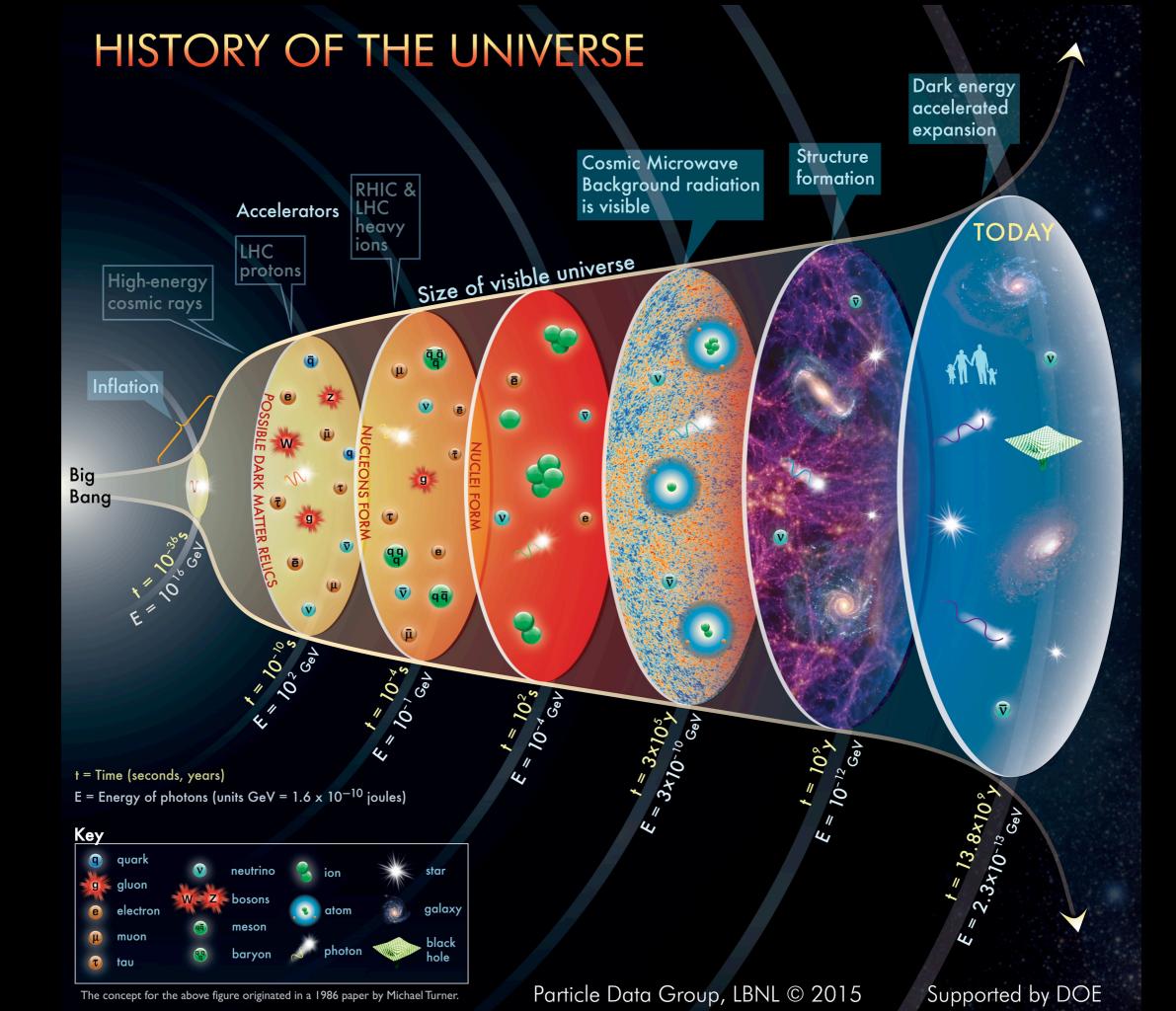


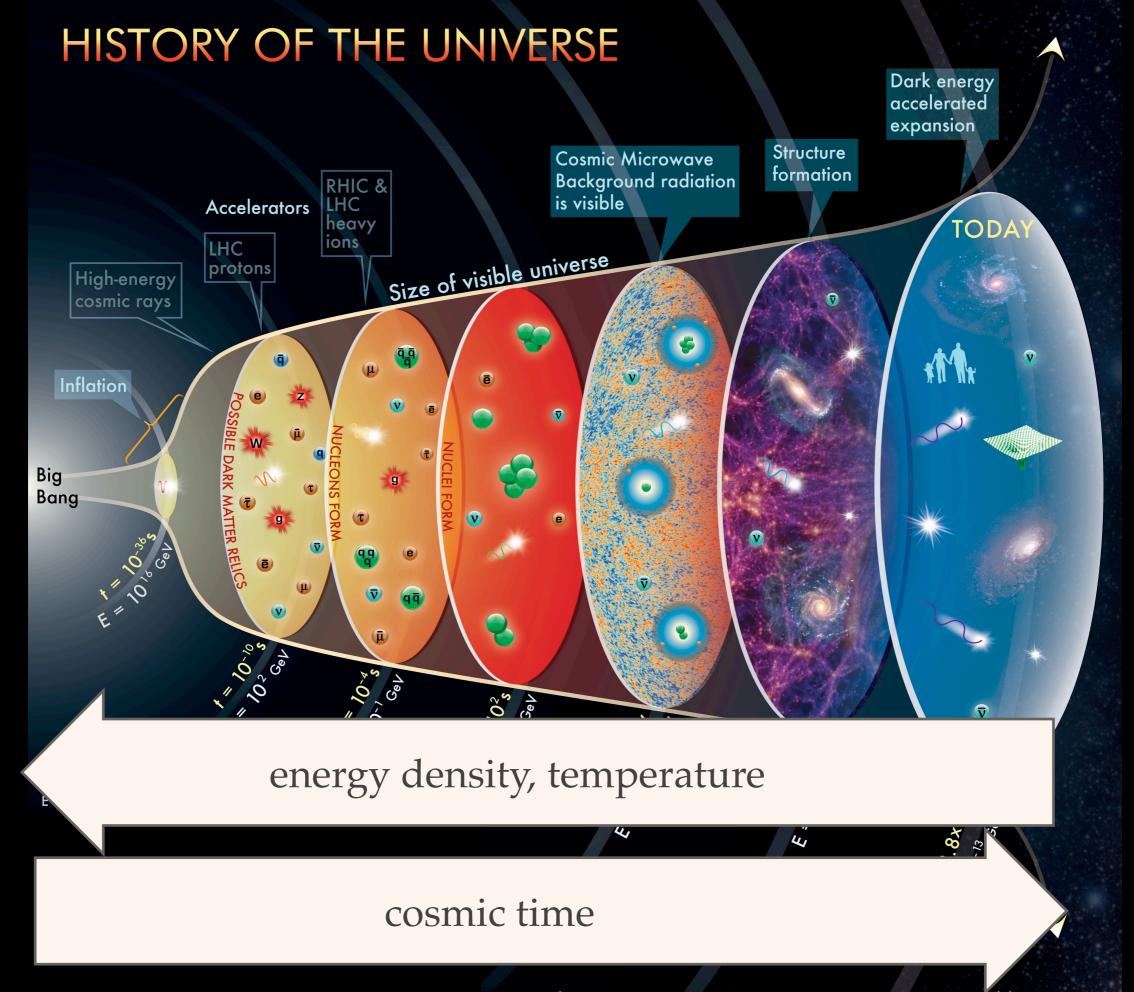
Effective Number of Relativistic Degrees of Freedom



Freeze Out and Freeze In





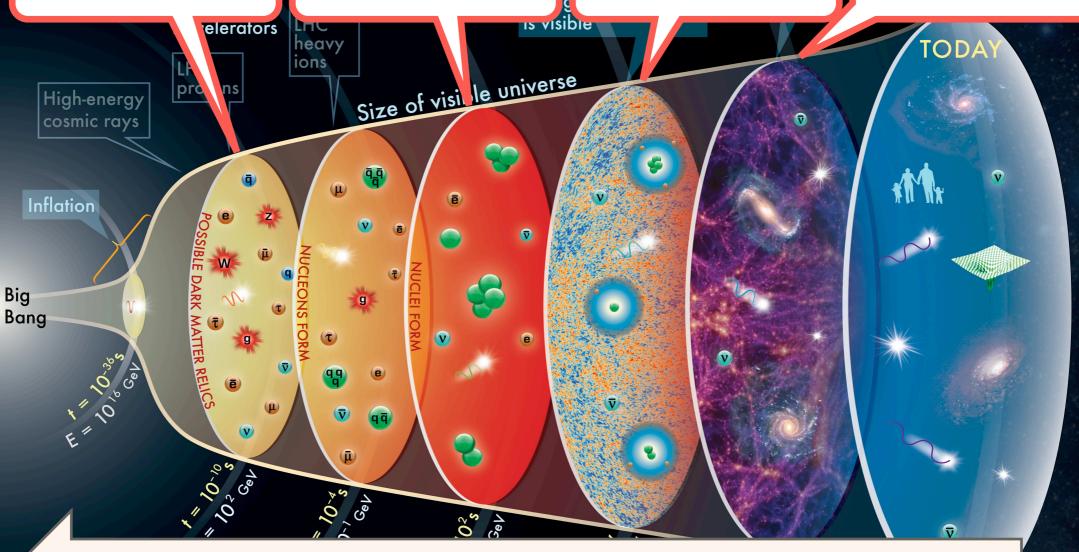


Large Hadron Collider High-energy cosmic rays Inflation

light element abundances

Cosmic Microwave **Background**

optical astronomy

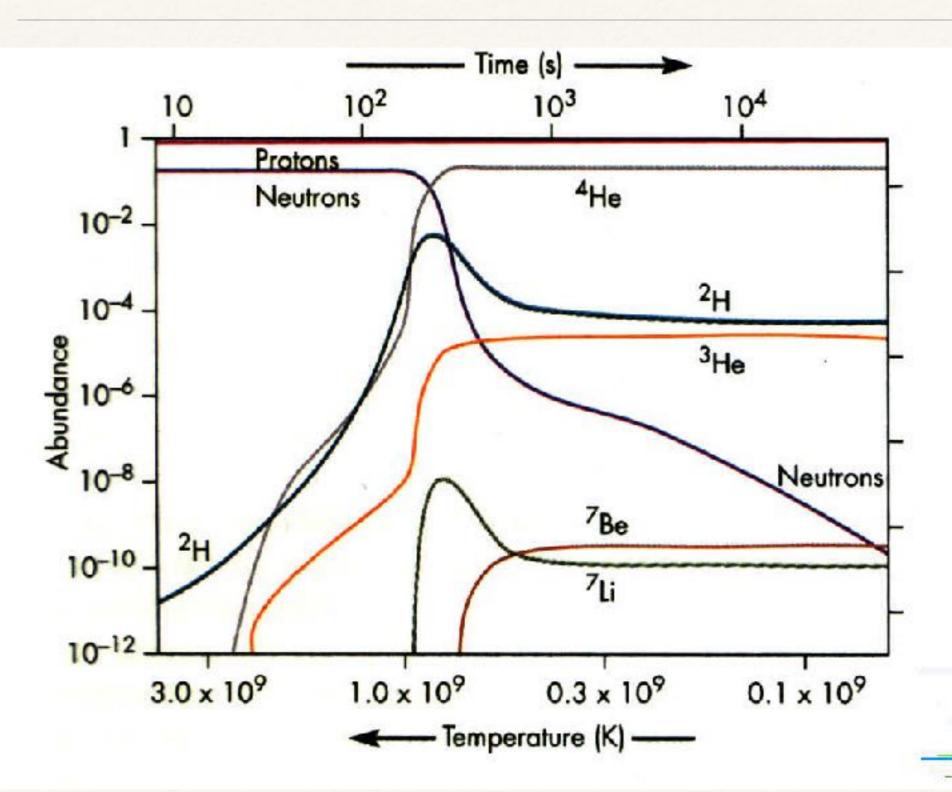


energy density, temperature

cosmic time

W

Big Bang Nucleosynthesis

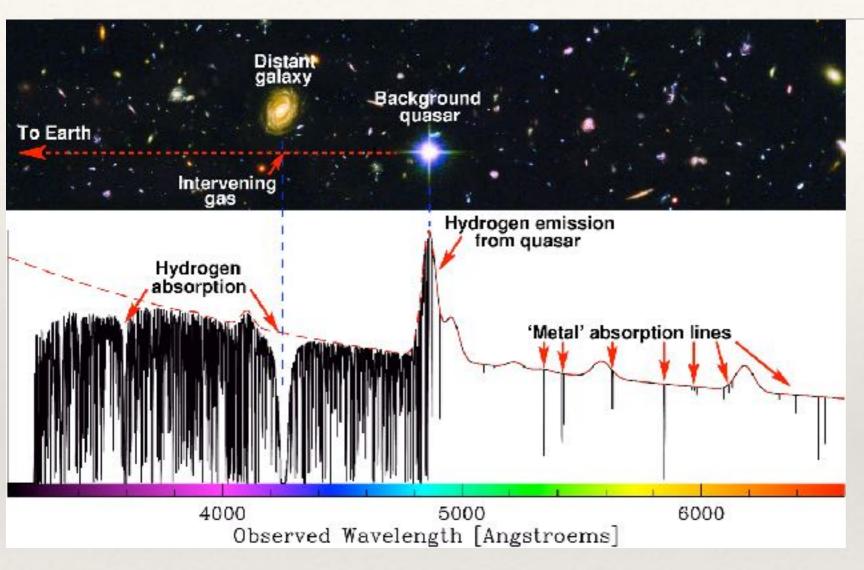


Light elements are produced in a chain of nuclear reactions.

The only unknown parameter is the baryon-to-photon ratio

Primordial light element abundances measure the baryon asymmetry!

Big Bang Nucleosynthesis

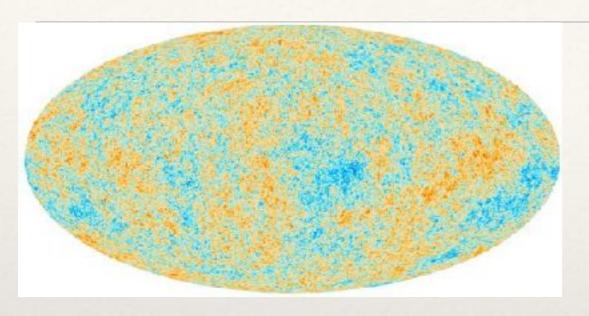


- light element abundances
 in intergalactic medium
 can be measured in
 quasar spectra
- Deuterium is sensitive to baryon asymmetry and not produced in stars

Constraint on baryon-to-photon ratio η:

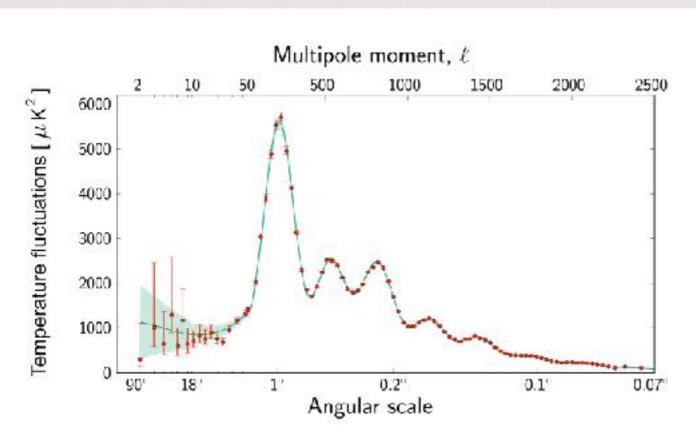
 $5.8 \times 10^{-10} < \eta < 6.6 \times 10^{-10}$ PDG 2016

Cosmic Microwave Background

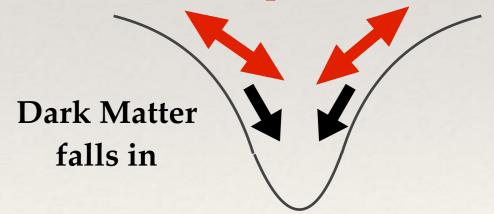


Constraint on baryon-to-photon ratio η:

$$6.03 \times 10^{-10} < \eta < 6.15 \times 10^{-10}$$
PDG 2016

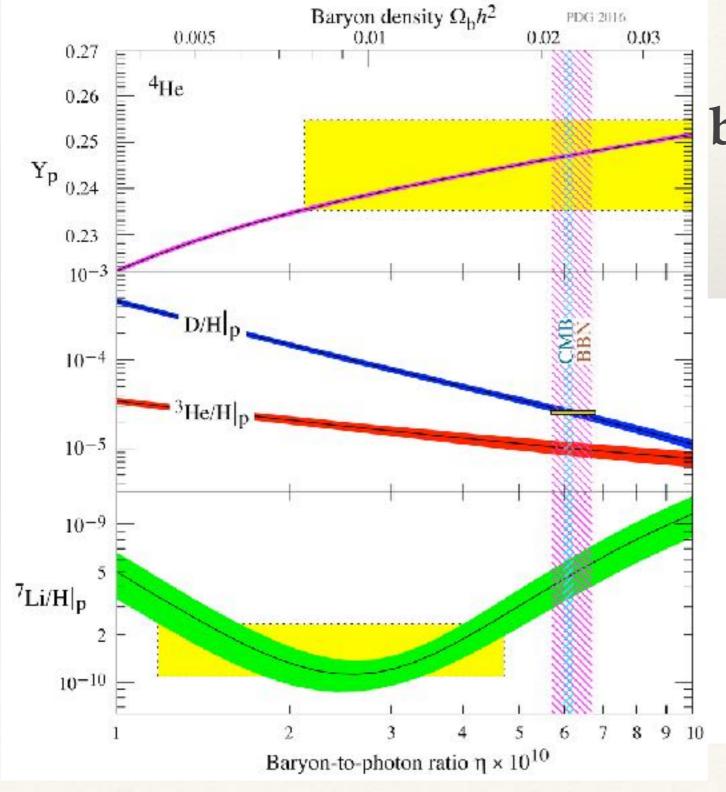


ordinary matter/radiation oscillates due to radiation pressure



gravitational potential well

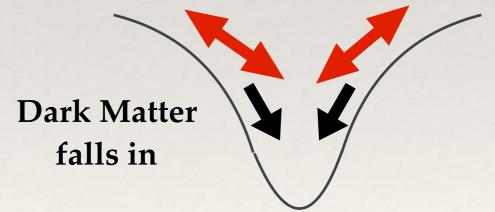
Cosmic Microwave Background



Constraint on baryon-to-photon ratio η:

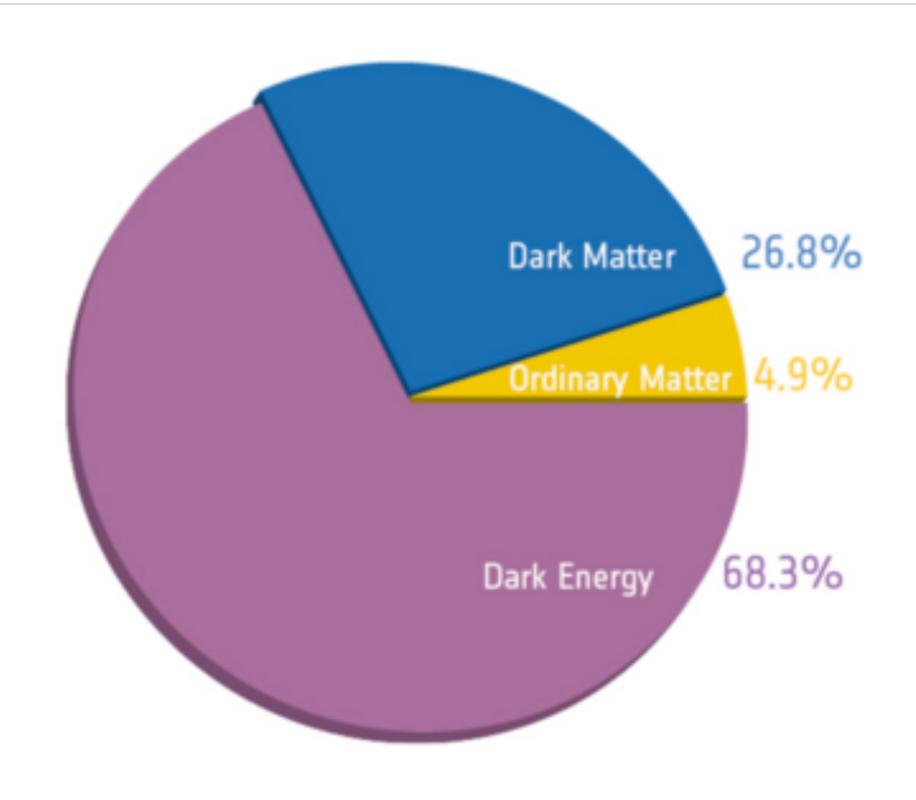
$$6.03 \times 10^{10} < \eta < 6.15 \times 10^{10}$$

ordinary matter/radiation oscillates due to radiation pressure



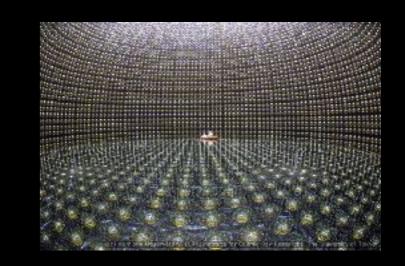
gravitational potential well

Energy Content of the Universe



What is the origin of neutrino mass?

Possible key to embed Standard Model in a more fundamental theory of Nature



known particles

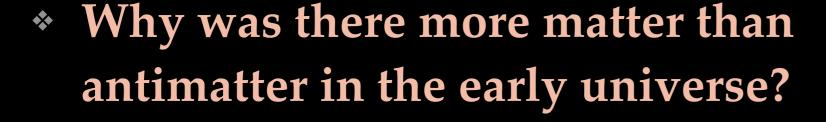
Dark Matter

82 %



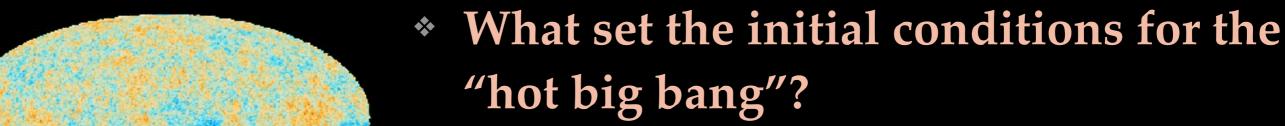
What is the Dark Matter made of?

It makes up most of the mass in the universe.

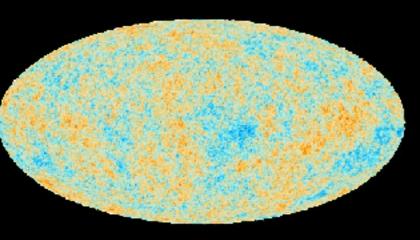


...so that some matter survived the mutual annihilation to form galaxies, stars etc.





Cosmic inflation? How did the transition to the radiation dominated epoch happen?



The Initial Conditions Problem

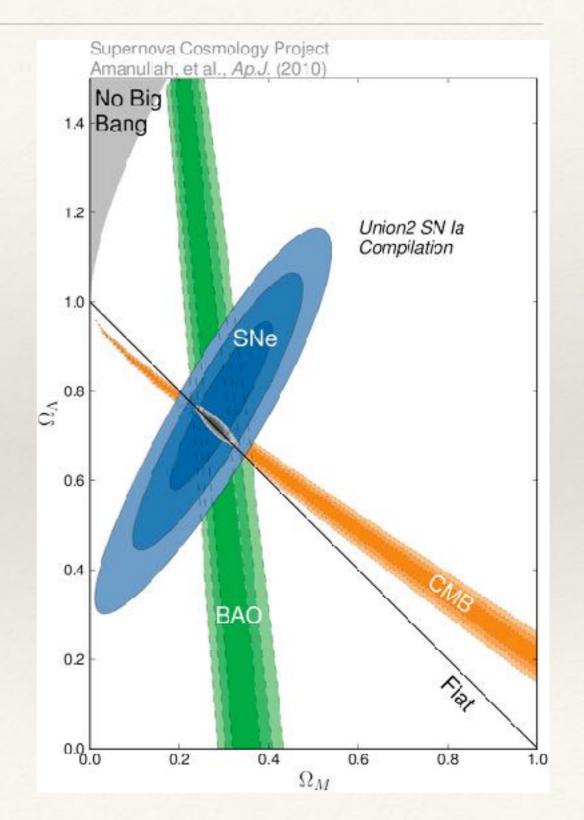
Flatness Problem

Flatness: Recall that the 3-dimensional curvature is

$$^{3}R = 6 k / a^{2}$$

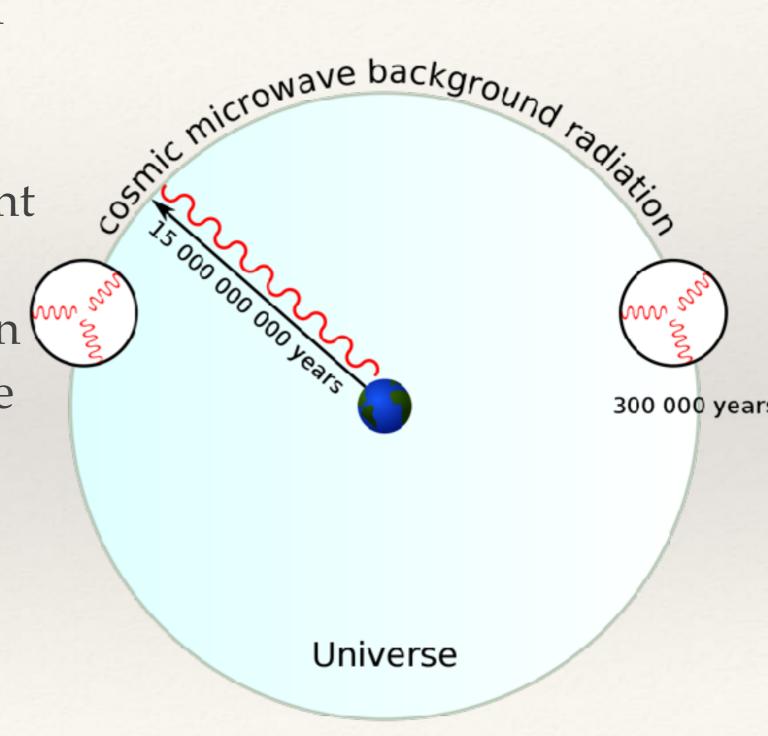
Observationally we know $(\Omega - 1) = k/(Ha)^2$ is small...

...curvature must have been super tiny in the past!!!



Horizon Problem

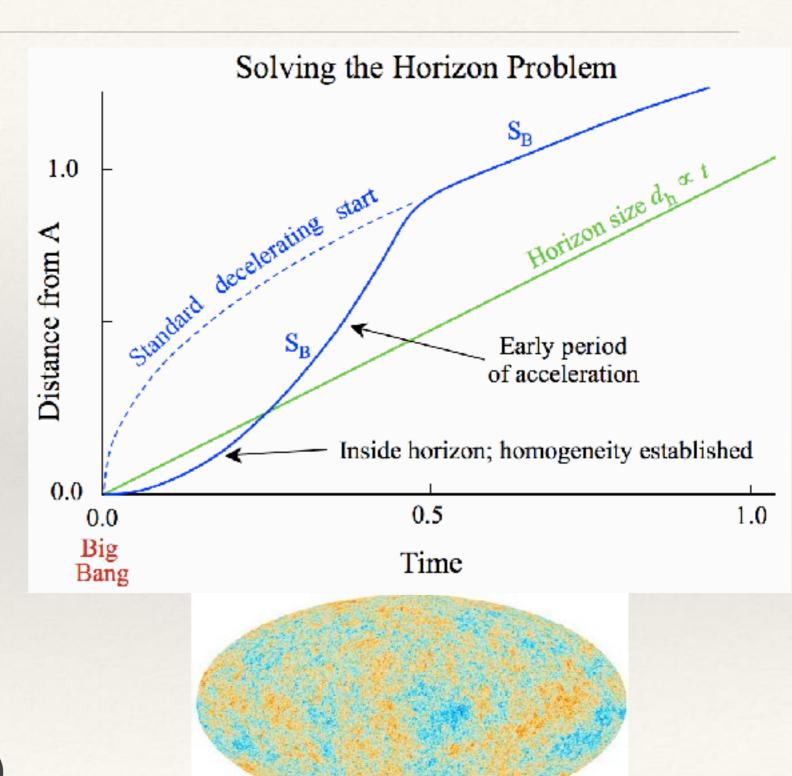
Comoving particle horizon at CMB decoupling was much smaller than the comoving distance that light has travelled to us (i.e., the patch that we know to be in equilibrium because we see the same temperature in opposite directions) Why do we see the same temperature in causally disconnected patches???



Inflation

Both problems (and some others) can be solved if the universe underwent a period of exponential growth... inflation!

This explains the CMB anisotropies as "blown up" quantum fluctuations and correctly predicted their properties (Gaussian, spectral index)



The Dark Matter Puzzle

There is compelling evidence that ~80% of the mass in the universe is

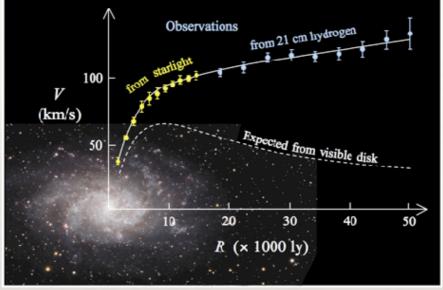
made of particles that are

neutral ("dark")

massive

non-baryonic

collisionless



There is compelling evidence that ~80% of the mass in the universe is

made of particles that are

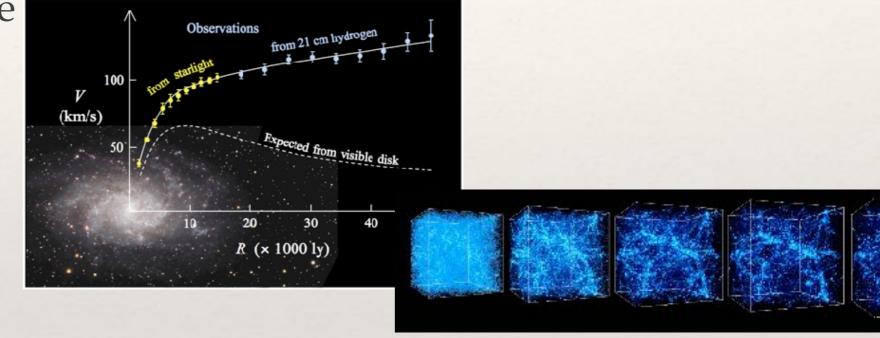
neutral ("dark")

massive

non-baryonic

collisionless

· cold



There is compelling evidence that ~80% of the mass in the universe is

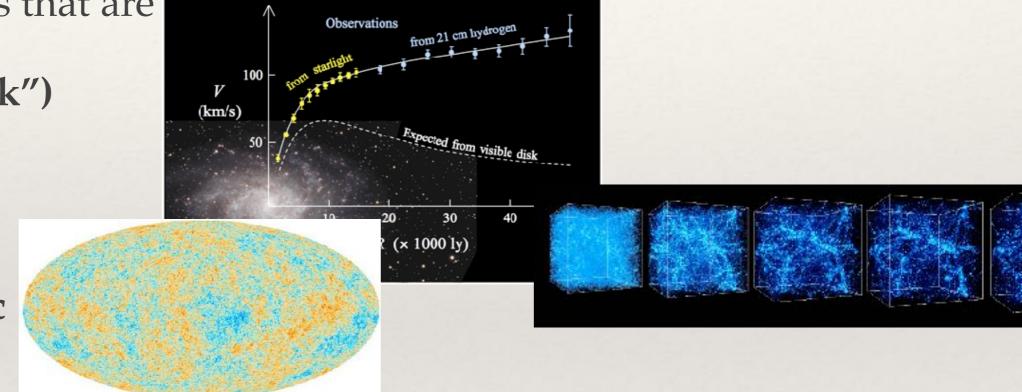
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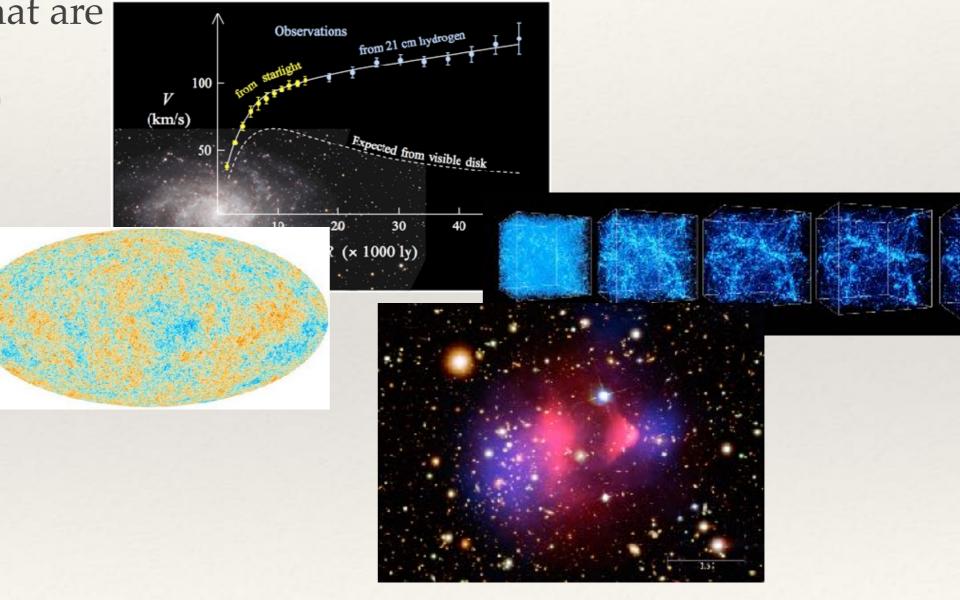
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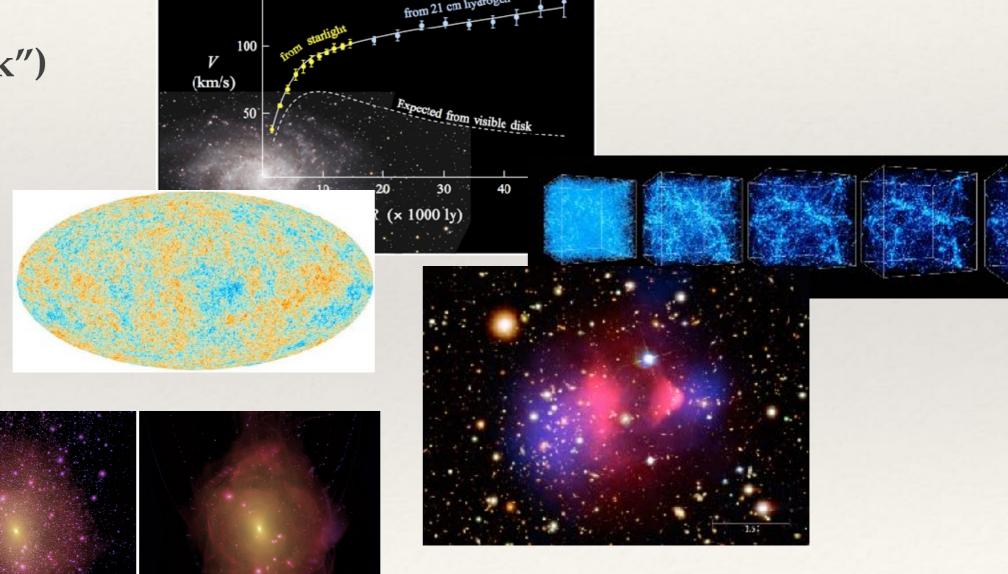
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There is compelling evidence that ~80% of the mass in the universe is

made of particles that are

neutral ("dark")

massive

non-baryonic

collisionless

cold

No known modification of gravity can explain all of this.

The Standard Model cannot, either.

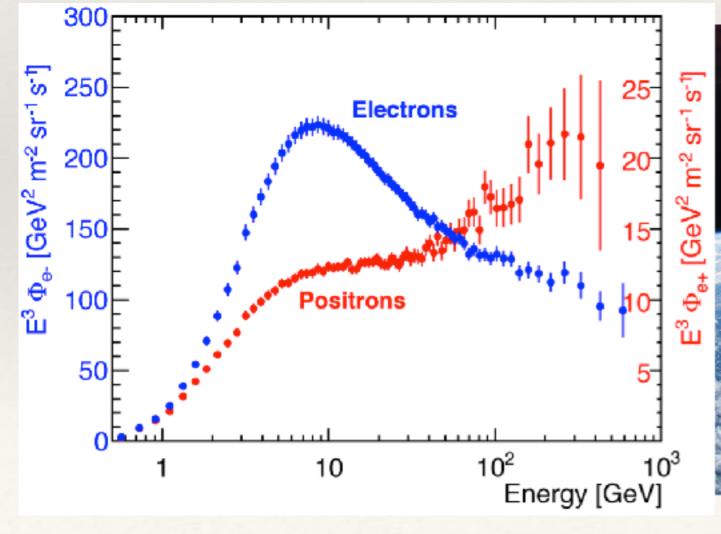
A simple extension by one (or several) new "Dark Matter" particles could do the job.

...but no such "Dark Matter" particle has been seen yet

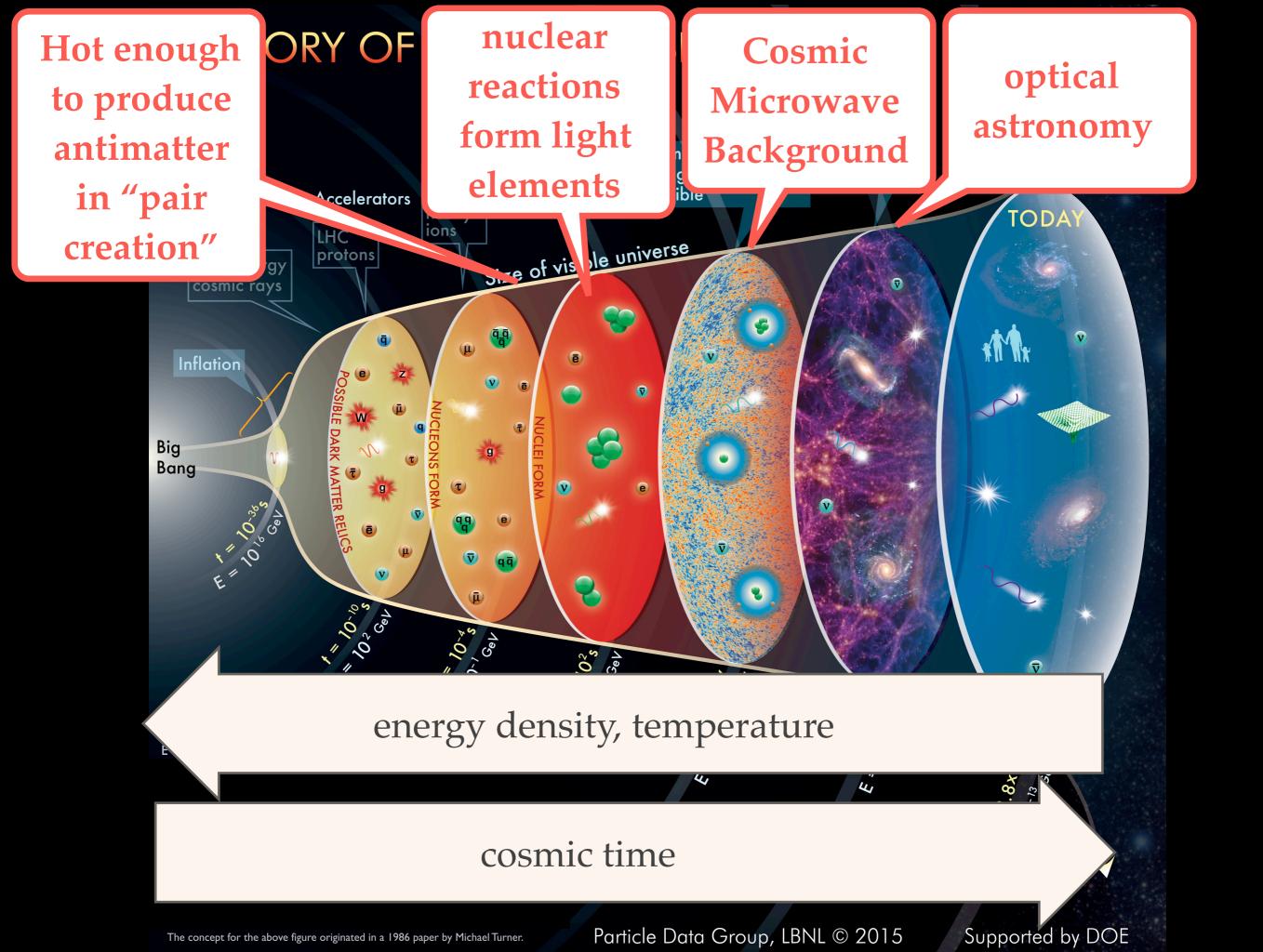
The Origin on (Baryonic) Matter

Antimatter in the Universe

* Is there any antimatter in the present universe? Yes. But little.







Baryon Asymmetry of the Universe

The observable universe contains almost no antimatter and a lot more photons than baryons. e.g. Canetti/MaD/Shaposhnikov arXiv:1204.4186

CMB constraint on baryon-to-photon ratio η :

6.03 x 10⁻¹⁰ < η < 6.15 x 10⁻¹⁰

(Planck Collaboration)

BBN constraint on baryon-tophoton ratio η : $5.8 \times 10^{-10} < \eta < 6.6 \times 10^{-10}$ (PDG)

Baryon Asymmetry of the Universe

The observable universe contains almost no antimatter and a lot more photons than baryons. e.g. Canetti/MaD/Shaposhnikov arXiv:1204.4186

 $rac{e^+}{rac}$ pair creation processes
freeze out $T > 2 \text{ mc}^2$ $T < 2 \text{ mc}^2$

CMB constraint on baryon-to-photon ratio η :

6.03 x 10⁻¹⁰ < η < 6.15 x 10⁻¹⁰

(Planck Collaboration)

BBN constraint on baryon-tophoton ratio η : $5.8 \times 10^{-10} < \eta < 6.6 \times 10^{-10}$ (PDG)

Baryon-to-Photon Ratio

- * When the temperature was very high, pair creation processes were in equilibrium.
- * When the temperature dropped below the positron mass, no new antiparticles could be produced, and all antiparticles were annihilated.
- * Obviously, some matter survived so there was more matter than antimatter in the early universe!
- * This "baryon asymmetry of the universe" was very small, it corresponds to today's "baryon-to-photon ratio" #nucleons/#photons: ~ 1 / 10.000.000.000

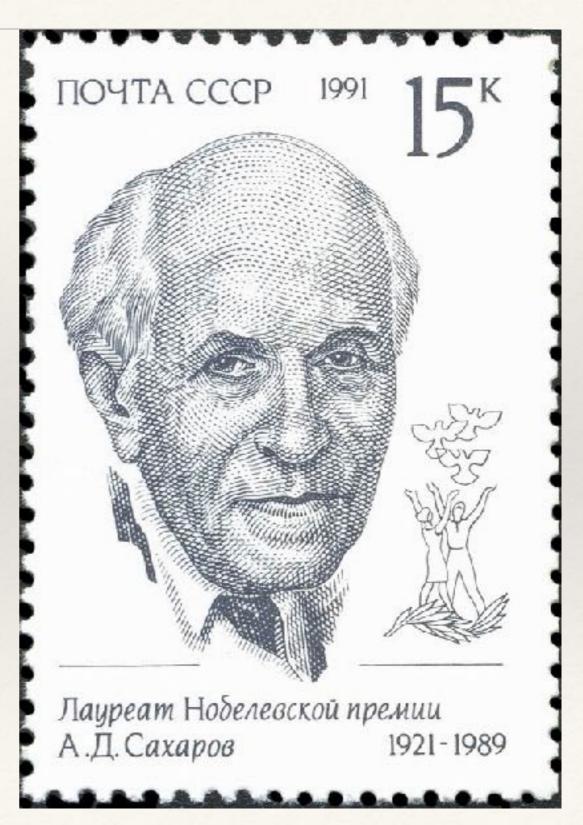
Where does the asymmetry come from?

Sakharov Conditions (1967)

* Baryon number violation

* C and CP violation

Deviation from thermal equilibrium



Where does the asymmetry come from?

Sakharov Conditions (1967)

* Baryon number violation

* C and CP violation

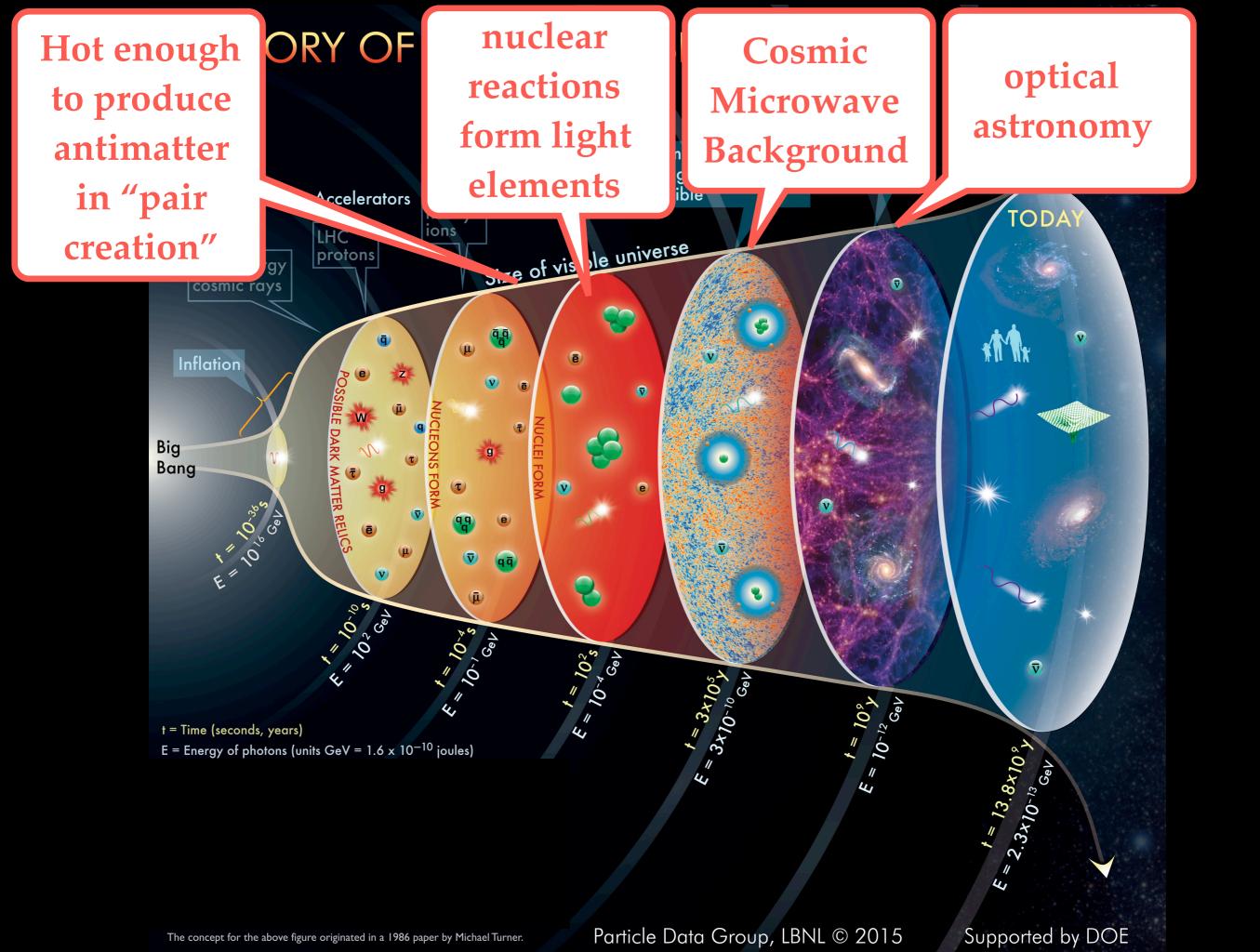
* Deviation from thermal equilibrium

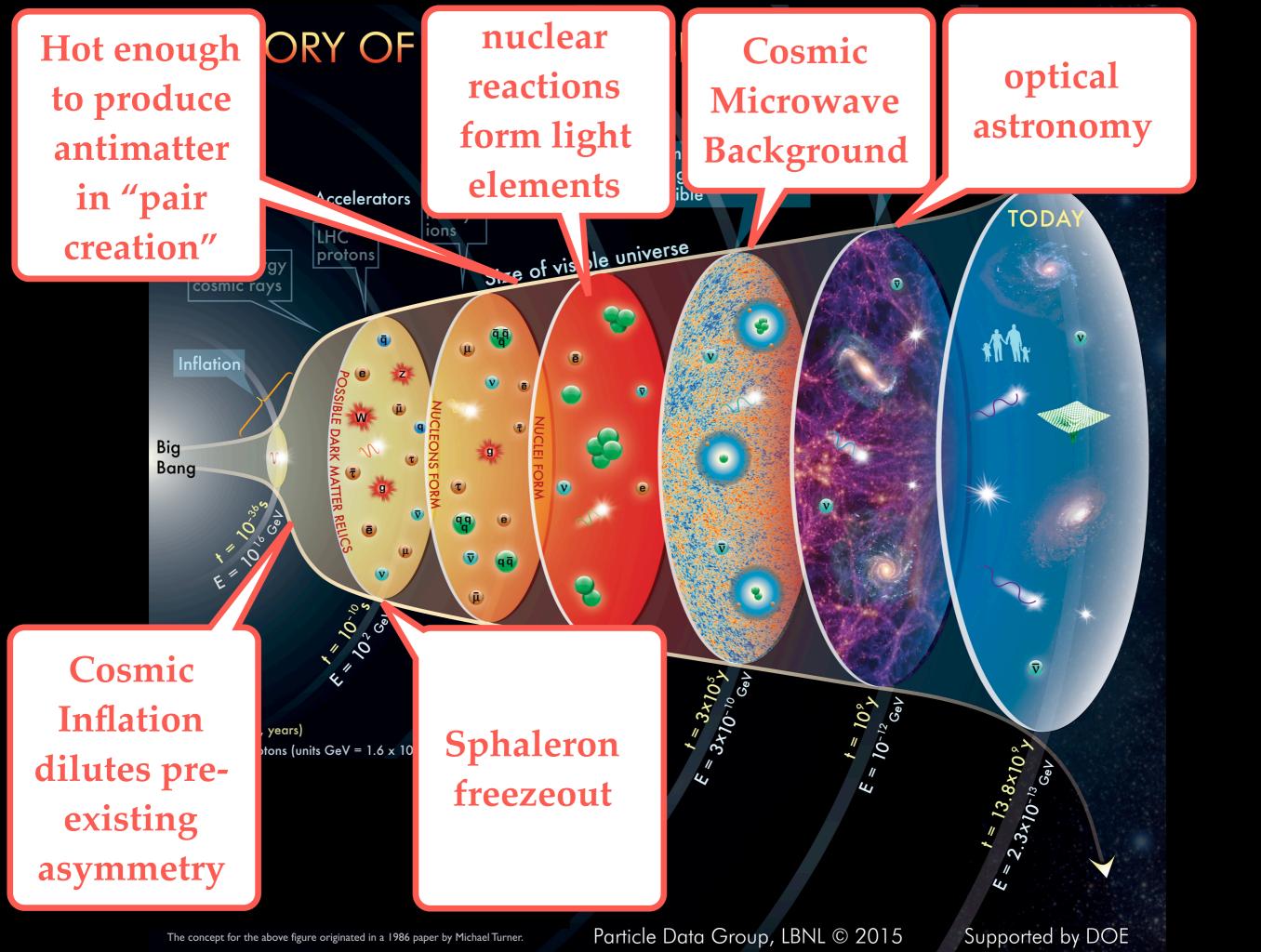
Exists in Standard Model at T > 130 GeV (sphaleron)

Exists in Standard Model (weak interaction, CKM phase)

...but Jarlskogg invariant too small!

Exists in Standard Model (Hubble expansion of the universe) ...but deviation too small!





Baryon
number diluted by inflation

Baryon number conserved

protons

Accelerat

POSSIBLE DARK MATTER RELICS

10°0°0 S

its GeV = 1.6

Baryon
number diluted by inflation

ccelerat

protons

10205

Cosmic
Inflation
dilutes preexisting
asymmetry

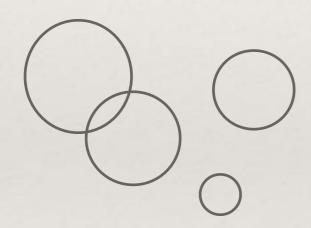
Baryon number conserved **Baryon asymmetry generated**

Baryon asymmetry generated ("Baryogenesis")

Where does the asymmetry come from?

Baryogenesis requires New Physics!

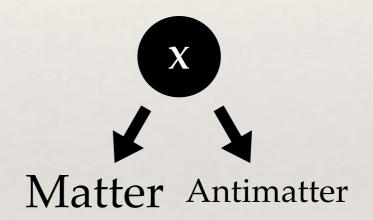
Cosmic phase transition?



Electroweak baryogengesis,

• • •

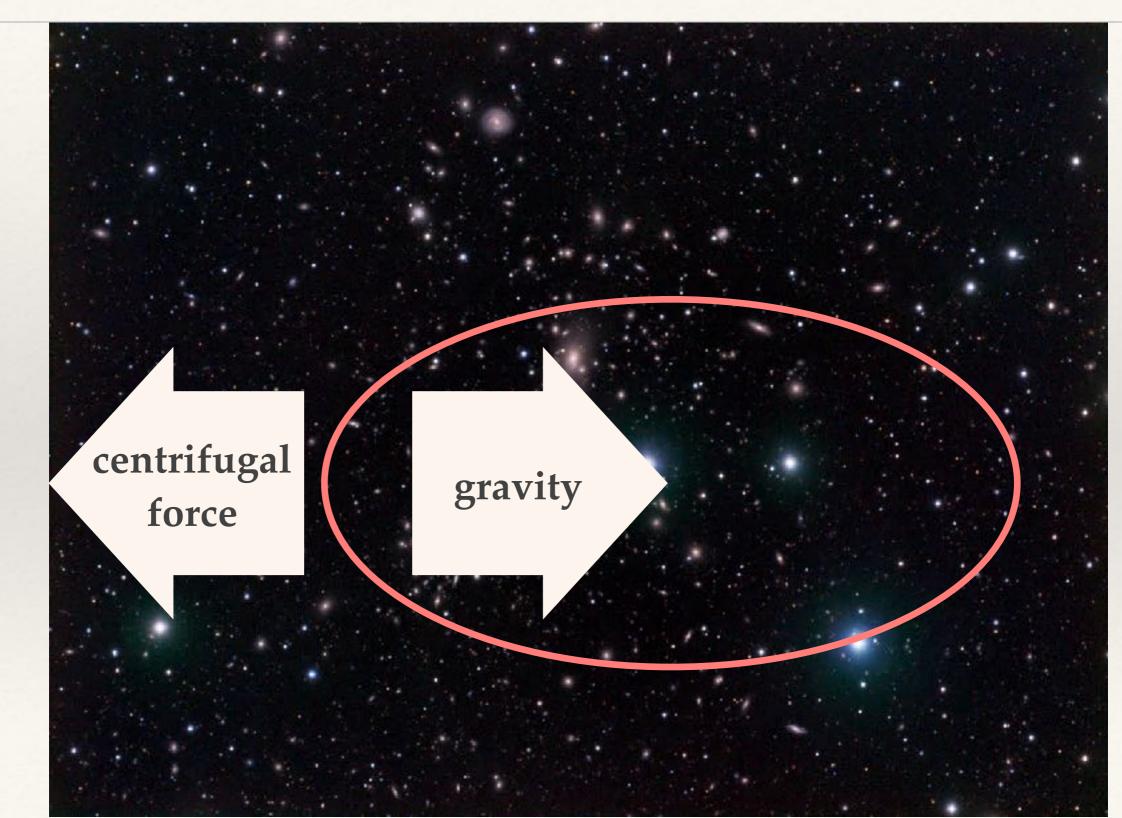
Decay of a heavy particle?

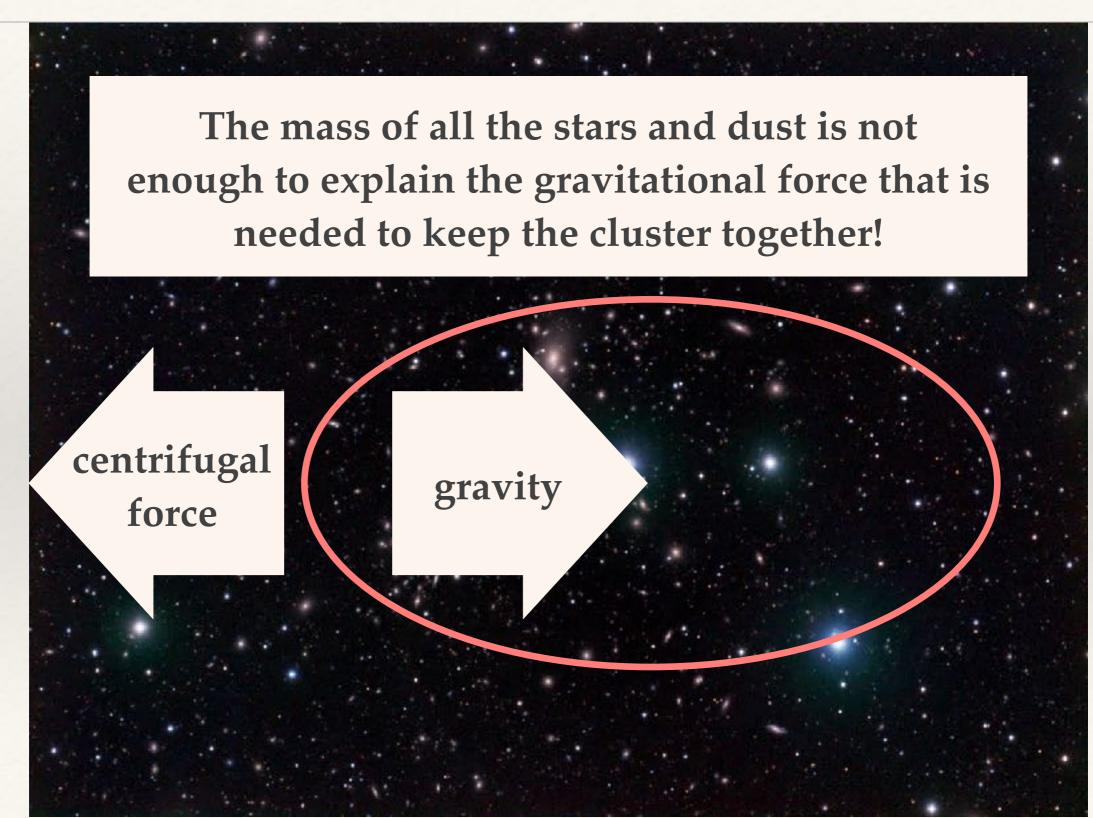


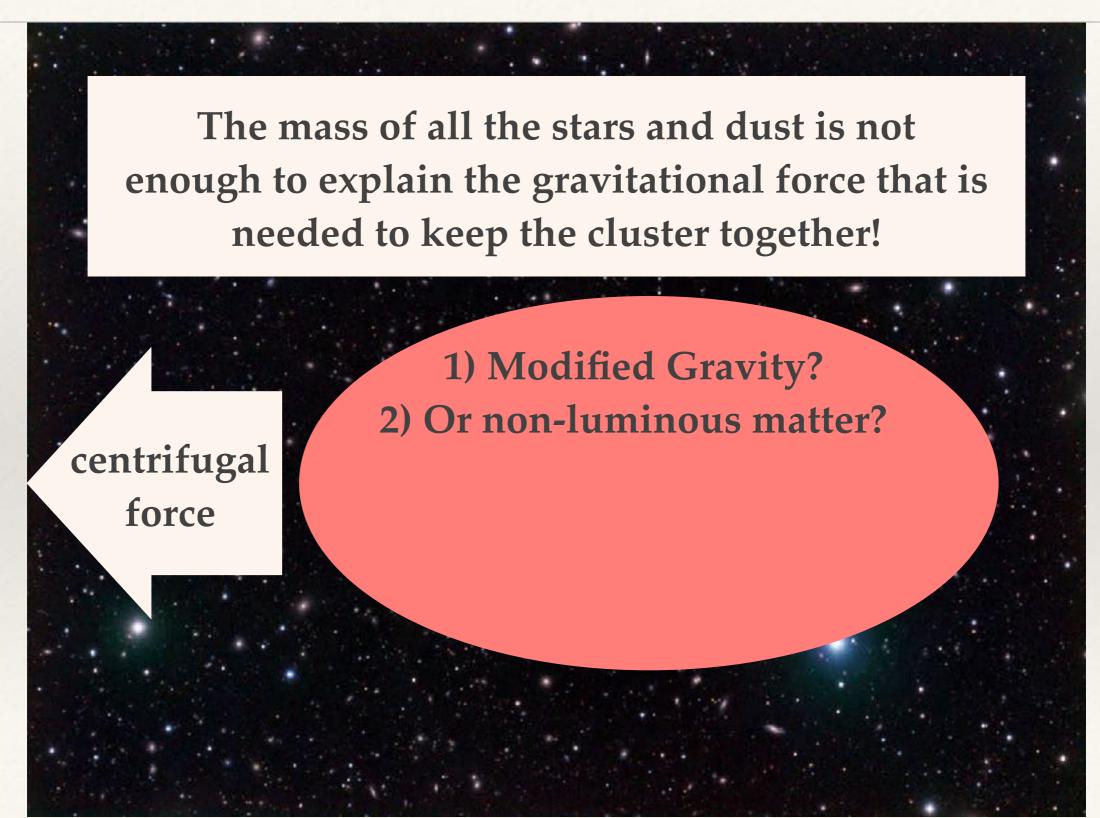
GUT baryogengesis, leptogenesis,

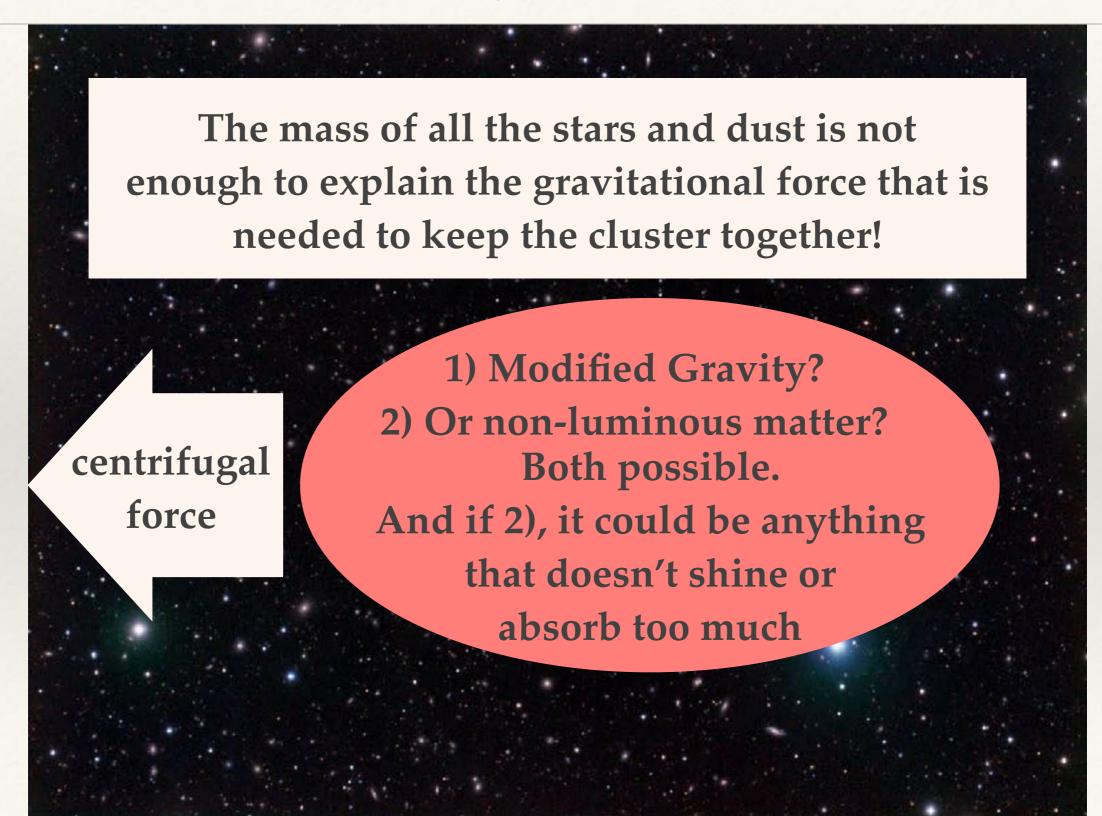
. . .

Appendix: Dark Matter vs Modified Gravity





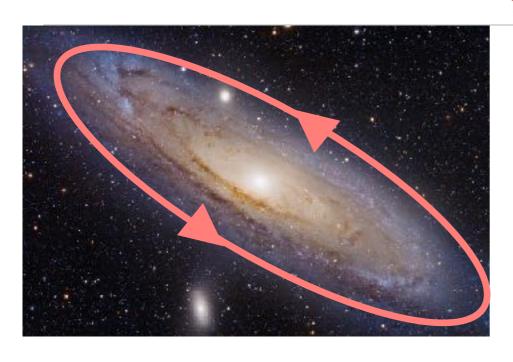




2) Galaxy Rotation Curves



2) Galaxy Rotation Curves

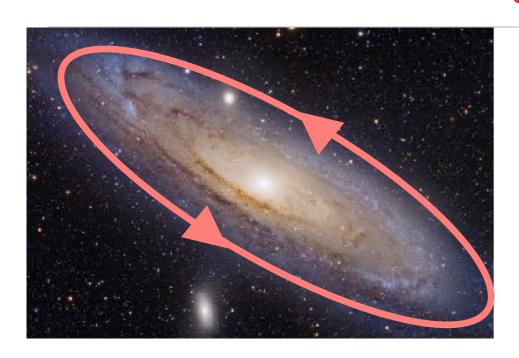


moves away from us

= redshifted

moves towards us

= blueshifted



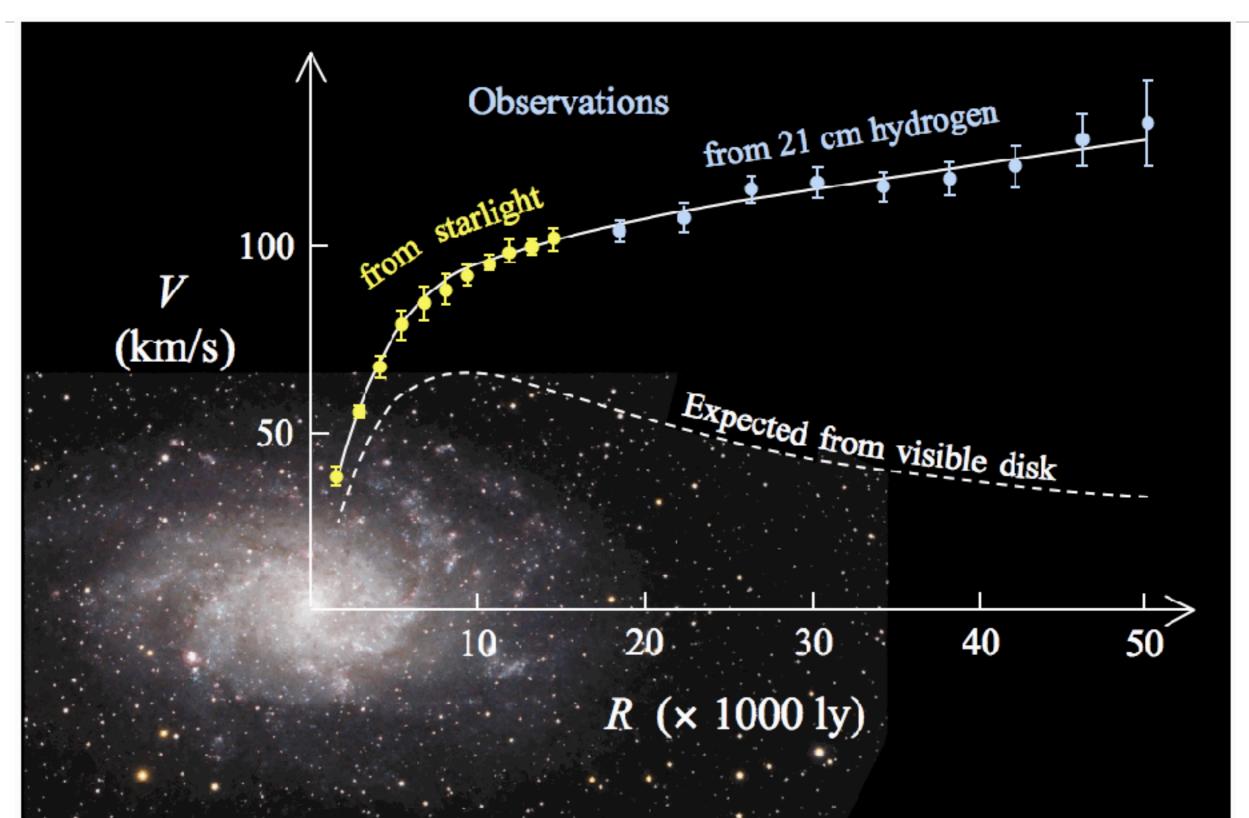
$$\frac{m\ M(r)}{r^2}G = \frac{m\ v^2(r)}{r}$$

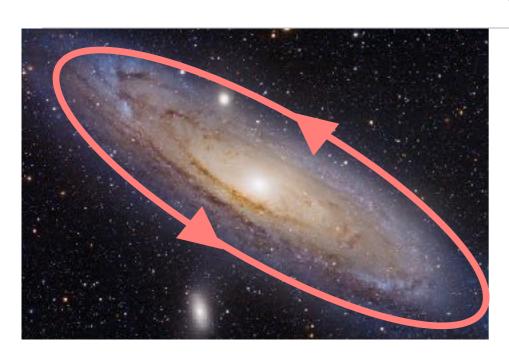
$$\Rightarrow v^2(r) = \frac{M(r)}{r}G$$

moves away from us

= redshifted

moves towards us = blueshifted



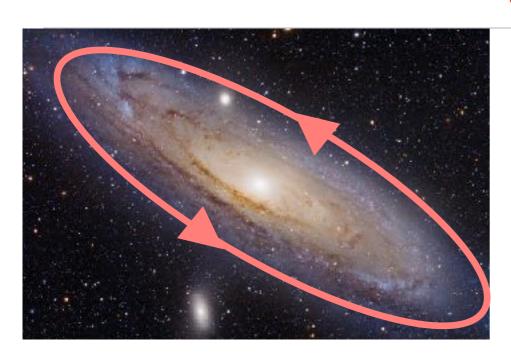


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- 1) Modified Gravity?
- 2) Or non-luminous matter?



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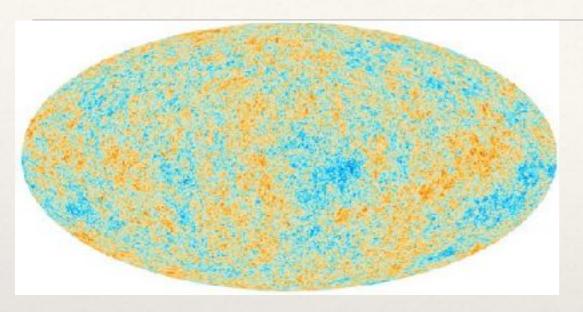
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1) Modified Gravity?

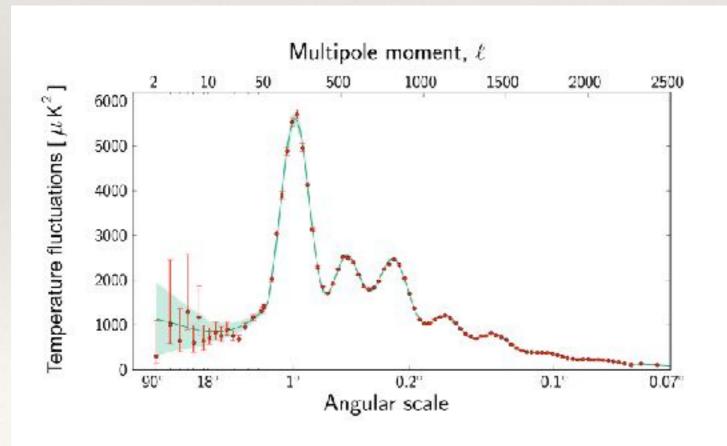
2) Or non-luminous matter? Both possible.

And if 2), it could be anything that doesn't shine or absorb too much

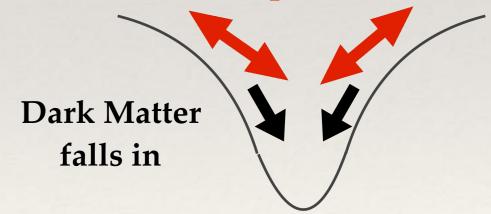
3) Cosmic Microwave Background



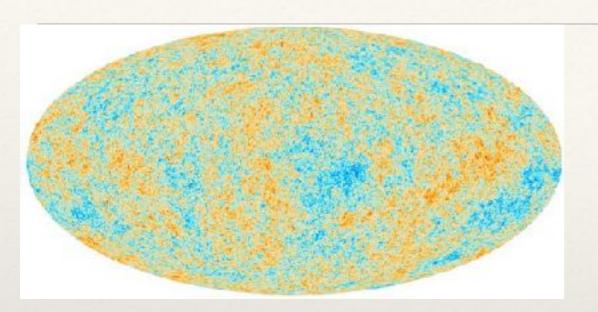
Radiation, baryons and Dark Matter affect the acoustic oscillations in the primordial plasma in different ways...



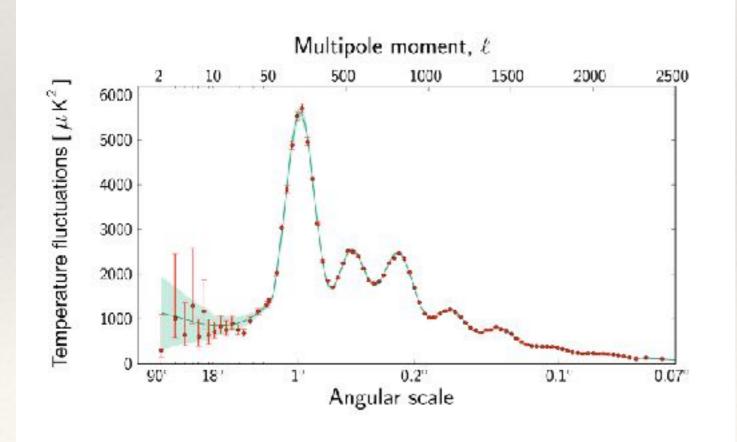
ordinary matter/radiation oscillates due to radiation pressure



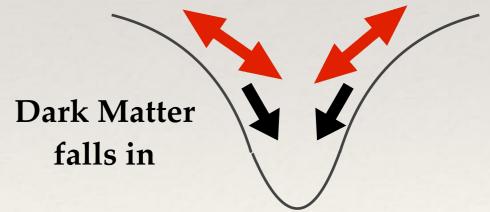
3) Cosmic Microwave Background



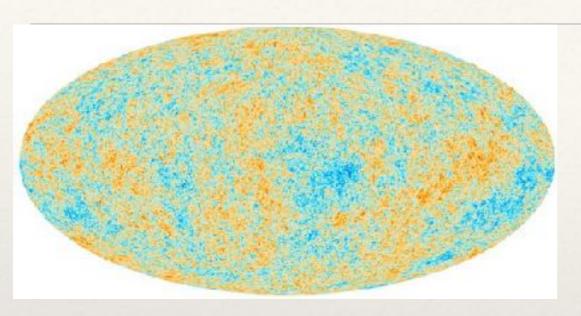
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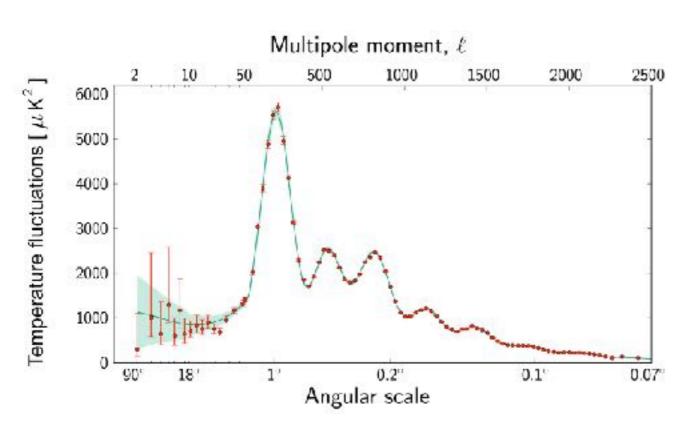
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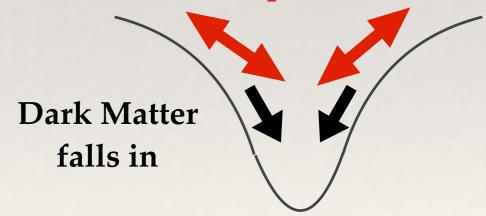
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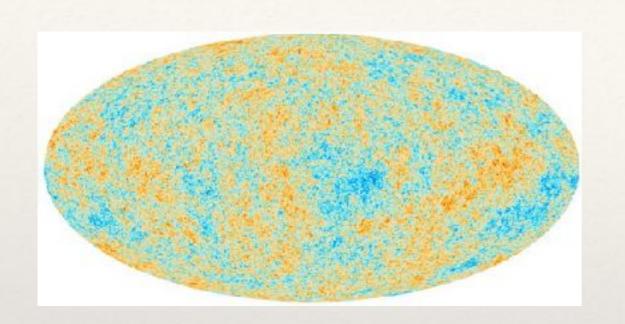


1) Modified Gravity?
2) Or non-luminous matter?
Very difficult to explain with modified gravity.
Shows that DM is not made of lumps of baryons.



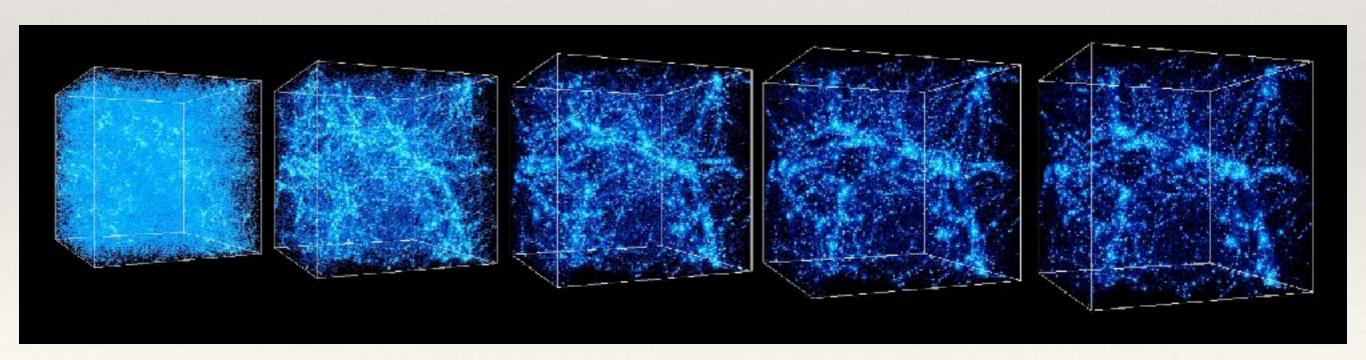
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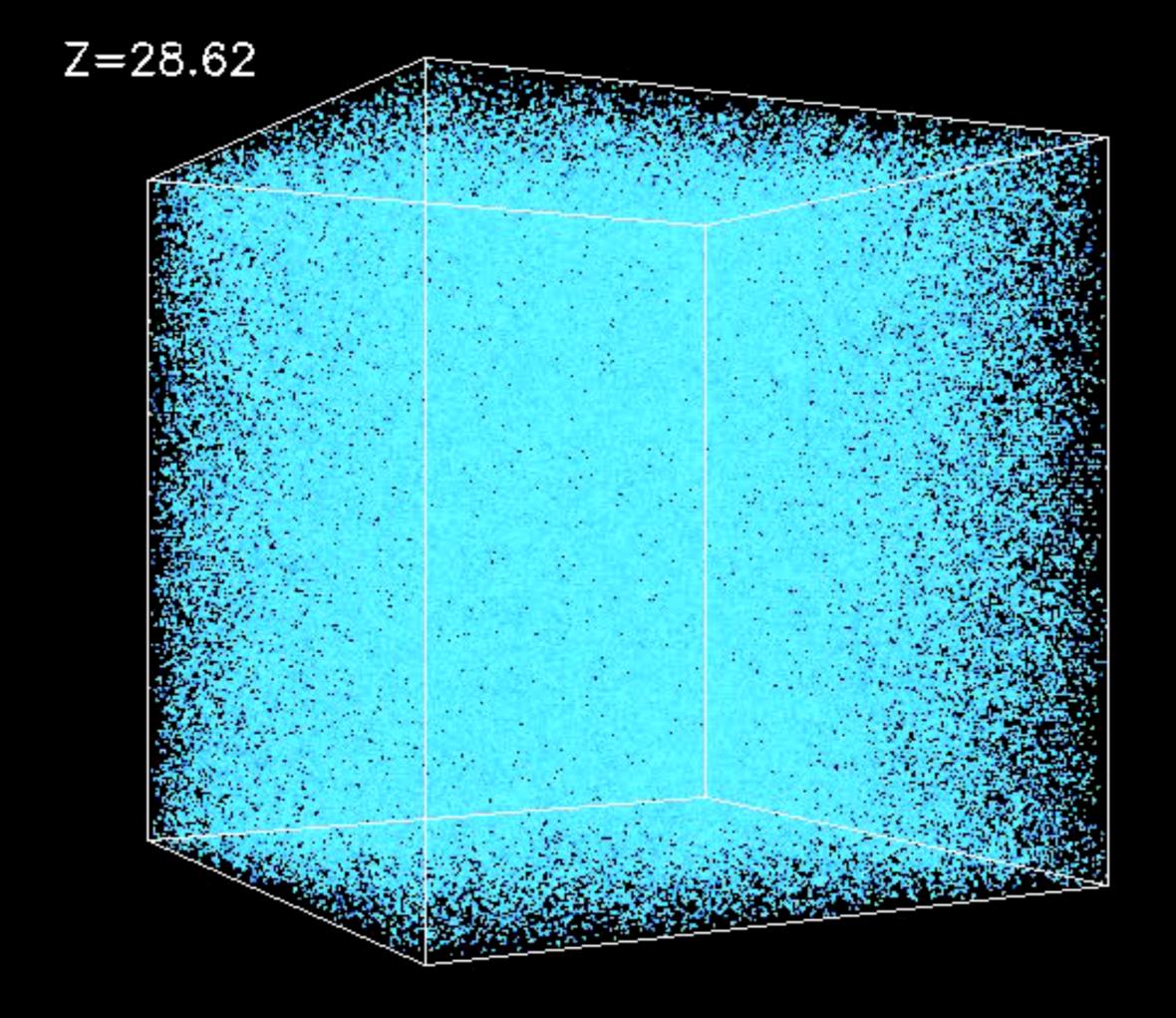


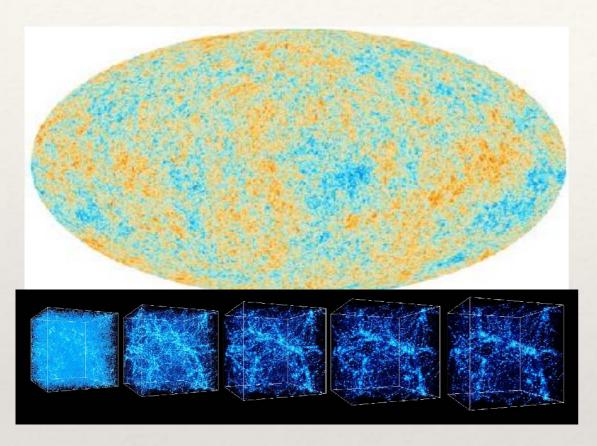


The early universe was very homogeneous and isotropic...

... structure formed by gravitational collapse of small inhomogeneities ~1/100.000







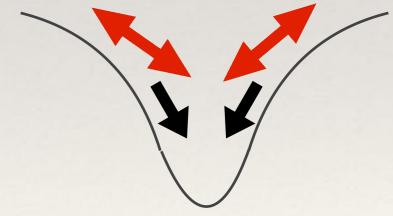
Simulations only agree with observation if this process starts before the CMB decoupled

Must be driven by particles that do not feel the radiation pressure!

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gravitational potential well

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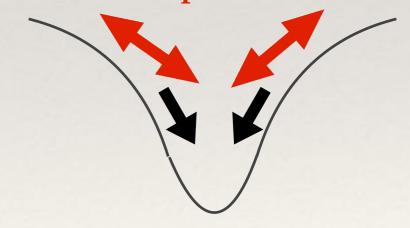
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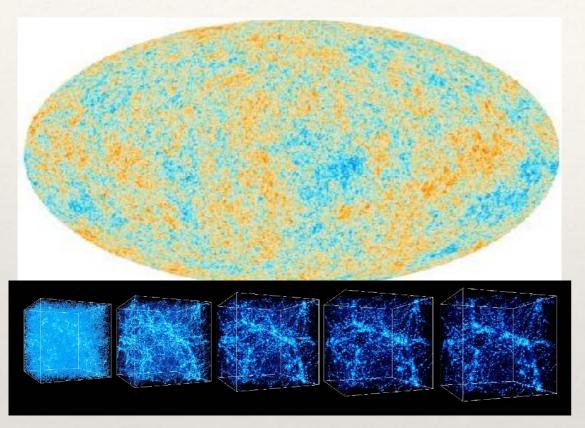
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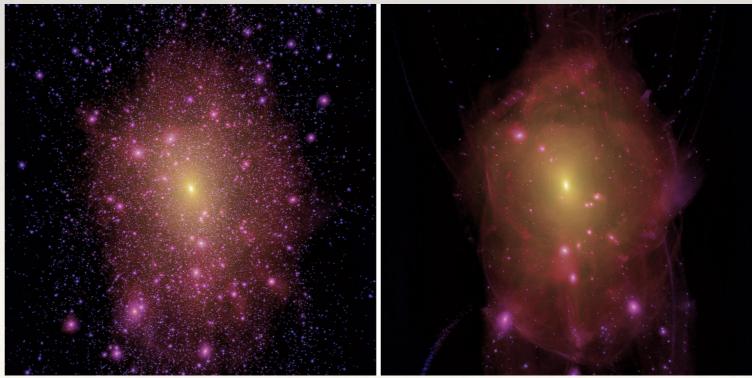


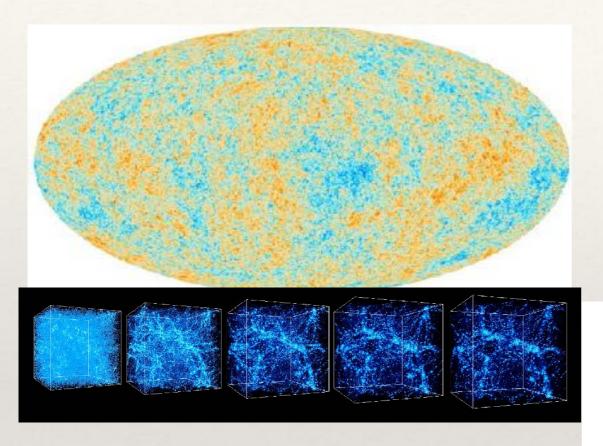
Movement of Dark Matter particles during structure formation "smears out" the inhomogeneities...

... which leads to a suppression of small scale structures in the universe!

"cold" DM

"warm" DM

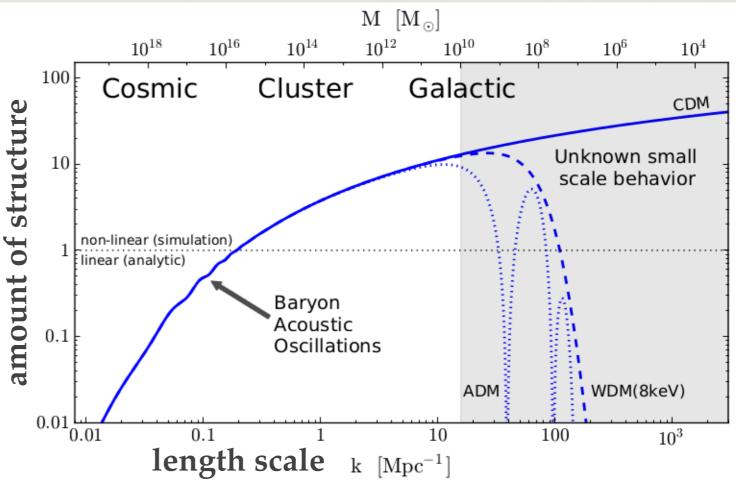




This can be quantified in terms of the matter power spectrum

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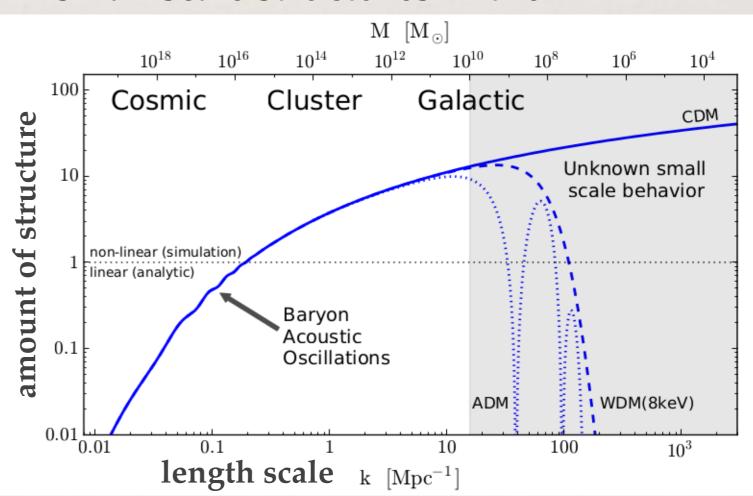


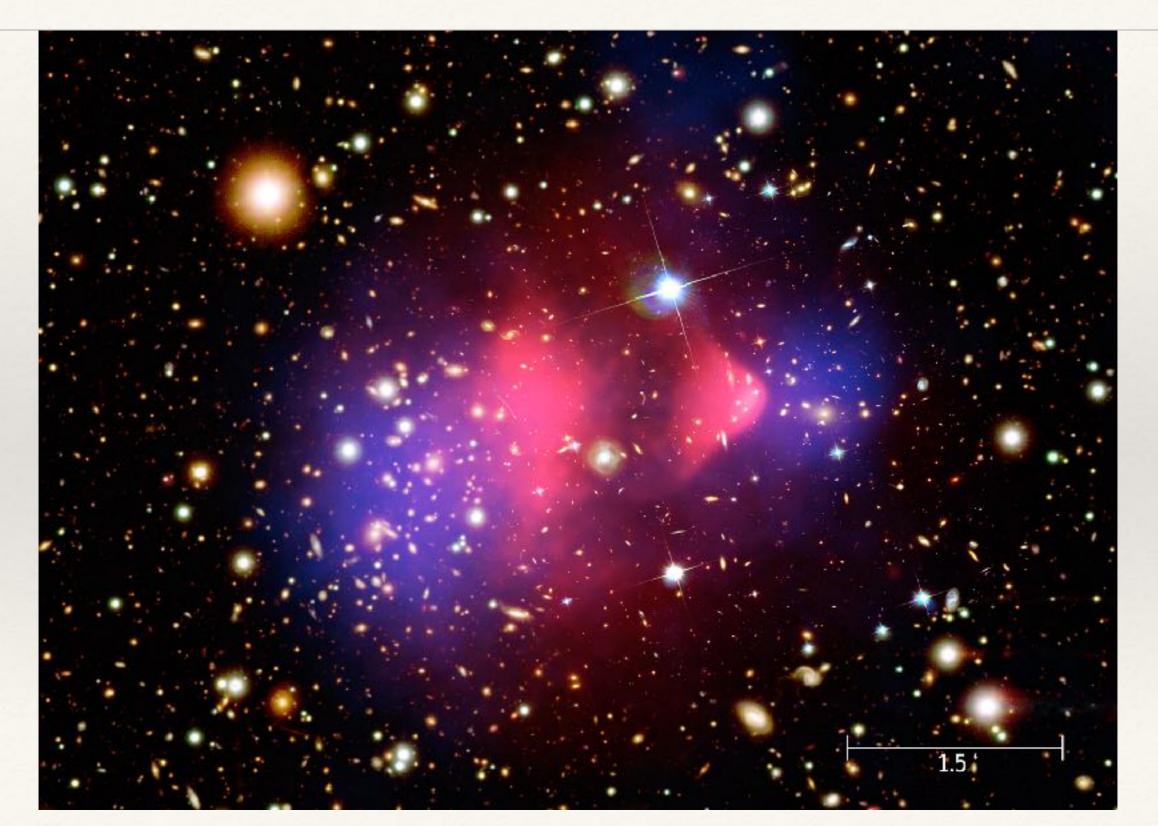
Dark Matter must be (relatively) "cold", i.e., non-relativistic at the time of structure formation

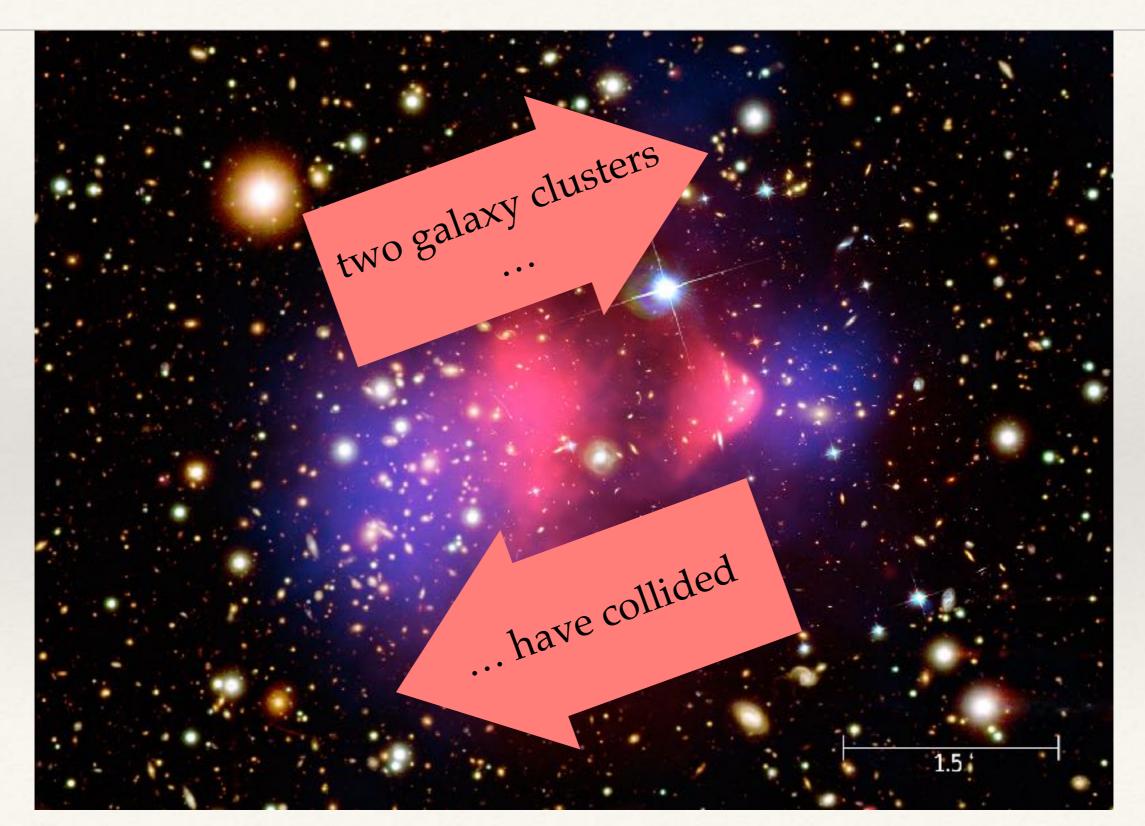
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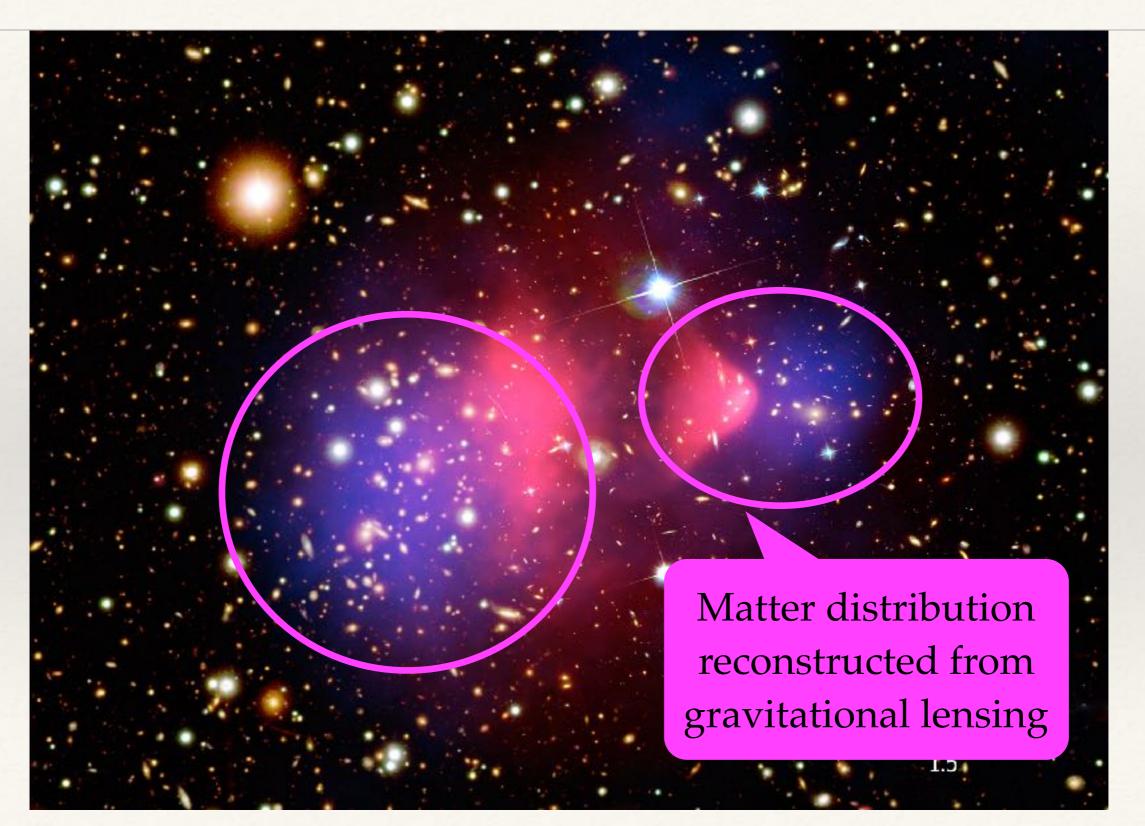
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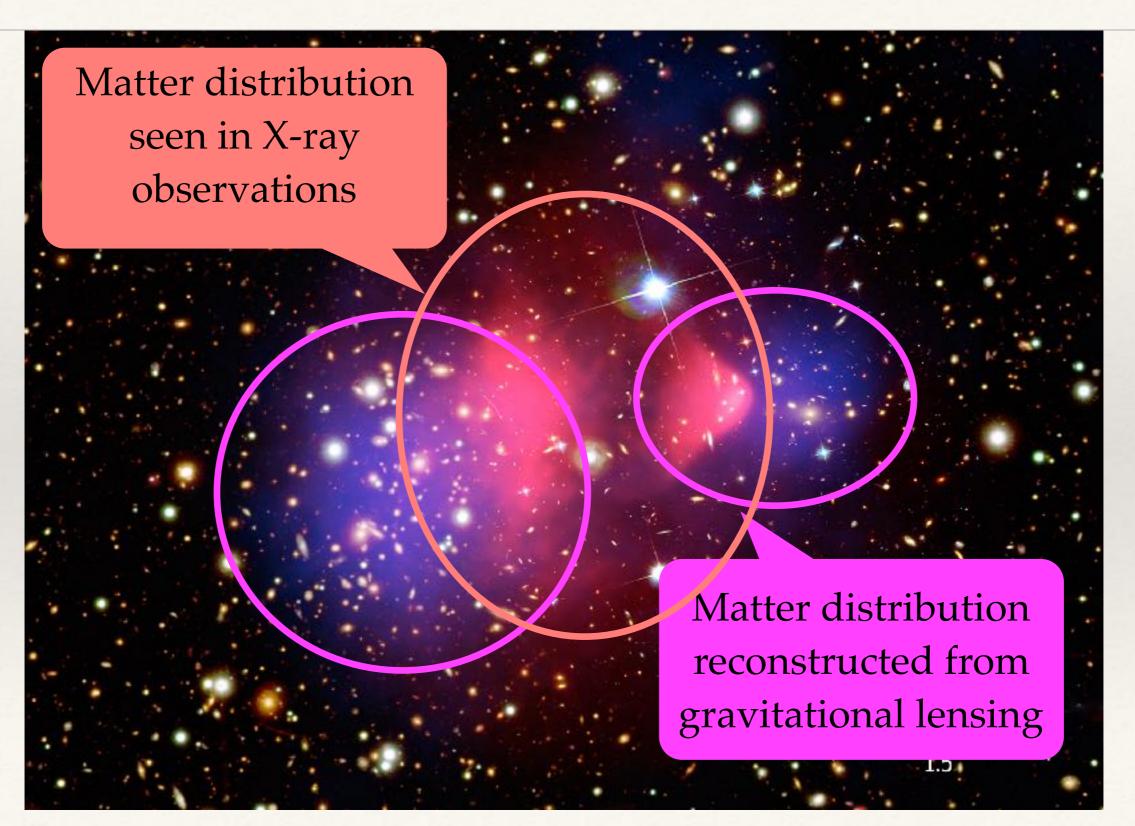


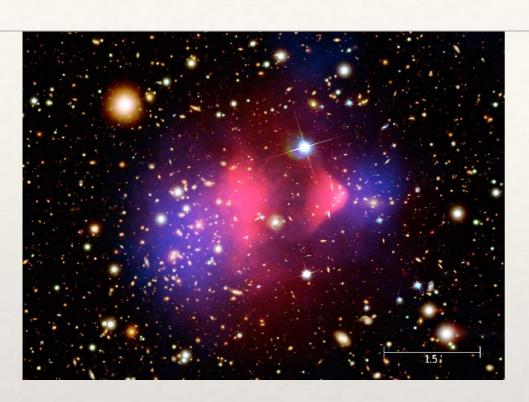






l friends) Matter distribution reconstructed from gravitational lensing

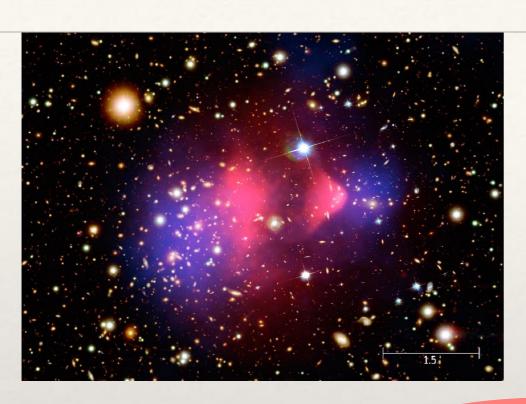




Interpretation:

The visible matter scatters and undergoes a merger.

The Dark Matter is collision free and passes.

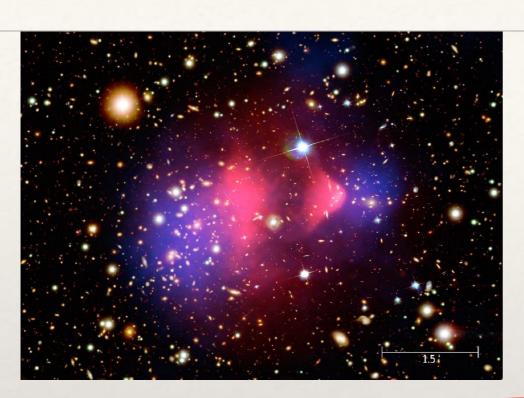


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