Heavy lons in non-LHC experiments - an overview



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Workshop on "Heavy lons and Hidden Sectors" 4-5 December 2018, Louvain La Neuve, Belgium

photo: Ukraine

Sonia Kabana, Heavy lons in non-LHC experiments, 4-5 December 2018, Louvain La Neuve, Belgium

Outline

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I Introduction

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Relativistic Heavy Ion Collisions and Quark Gluon

The QCD phase transition between hadronic and partonic phase

QCD on the lattice predicts a cross over at zero net baryon density with critical temperature Tc~154+-9 MeV (2014), critical energy density ~0.6 GeV/fm^3

(Nuclear Density: rho=0.15 GeV/fm³ Density inside Nucleon: rho=0.5 GeV/fm³)



Zero net baryon density

F. Karsch, Lect. Notes Phys. 583 (2002) 209, hep-lat/0106019



The order of the transition depends on the parton masses. A cross over is expected by Lattice QCD for the physical point (for the physical u,d,s masses).

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The transition from quarks and gluons to hadrons is believed that took place few 10-6 sec after the Big Bang.

The QCD phase transition is the only phase transition of the early universe that can be reproduced in the Lab today.

Why is this possible ? Because T(critical) is expected to be ~200 MeV, and this is in principle reachable with todays accelerators

The expected QCD phase diagram



Phases of QCD Matter

Areas of different net baryon densities and temperatures can be probed using different collision energies and nuclei.

The order of the transition is expected to change with the net baryon density.

Goal: explore experimentally the QCD phase diagram (order of transition, critical point, properties of the QGP).

The QCD Phase Diagram and the Path of Early Universe through this diagram



Calculation and plot of the Path of Early Universe in (T, mu-B) plane was done by Uli Heinz and later from S.K., P. Minkowski, J Phys G 28 (2002) 2063-2067, hepph/0204103 (fig.1).

S.K. P. Minkowski, Space Sci.Rev. 100 (2002) 175-192



Later calculation: M Fromerth, J. Rafelski, astro-ph/0211346



Accelerators

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Relativistic Heavy Ion Collider

at the Brookhaven Lab, Long Island, New York, USA





RHIC has been exploring nuclear matter at extreme conditions over the last 15 years 2000-2015

4 experiments initially: STAR PHENIX BRAHMS PHOBOS

Still runing: STAR and PHENIX

Colliding systems:

p+p, d+Au, Cu+Cu, Au+Au Cu+Au, U+U Energies A+A :

√**s**_{NN} = 62, 130, 200 GeV and low energy scan 7.7, 11.5, 19.6, 22.4, 27, 39 GeV + Fixed target

RHIC

Luminosity evolution of hadron colliders



Luminosity per nucleon pair $L_{NN} = A_1 A_2 L_1$

Peak Luminosity AuAu 200 GeV 2016 : 1.55x10²⁸ cm-2 s-1

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RHIC integrated Luminosity



Current non-LHC Experiments with Heavy Ion program

STAR at RHIC



PHENIX at RHIC (data analysis only)



Stopped data taking but data analysis is in progress

HADES at GSI

NA61/SHINE at SPS



Current non-LHC Experiments with Heavy Ion program PHENIX at RHIC (data analysis only)



Hadron, electron, photon identification Muon identification

Stopped data taking but data analysis is in progress



The STAR detector





- Tracking and PID (full 2π) TPC: $|\eta| < 1$ TOF: $|\eta| < 1$ BEMC: $|\eta| < 1$ EEMC: $1 < \eta < 2$ HFT (2014-2016): $|\eta| < 1$ MTD (2014+): $|\eta| < 0.5$
- MB trigger and event plane reconstruction BBC: $3.3 < |\eta| < 5$ EPD (2018+): $2.1 < |\eta| < 5.1$ FMS: $2.5 < \eta < 4$ VPD: $4.2 < |\eta| < 5$ ZDC: $6.5 < |\eta| < 7.5$
- On-going/future upgrades iTPC (2019+): $|\eta| < 1.5$ eTOF (2019+): $-1.6 < \eta < -1$ FCS (2021+): $2.5 < \eta < 4$ FTS (2021+): $2.5 < \eta < 4$

Previous and present experiments

BNL AGS (E866, E917 ...)

CERN SPS (NA35, NA36, NA49, NA44, WA80, WA93, WA98 WA85, WA94, WA97, NA57, NA50, NA52. NA61/SHINE...)

BNL RHIC (STAR PHENIX PHOBOS BRAHMS)

CERN LHC (ALICE, CMS. ATLAS,, LHCb)

Reach of accelerators in terms of initial Temperature



Signatures of the Quark Gluon Plasma

Direct photons from QGP \rightarrow T(QGP) Strangeness enhancement (Mueller, Rafelski 1981) \rightarrow K/pi U,d,s yields for T(freeze out) or pT slopes (Van Hove, H Stoecker et al) \rightarrow plateau vs energy at Tc \rightarrow e_init(crit), sqrt(s)("crit") Multiquark states from QGP (Greiner et al) \rightarrow 'small QGPlumps' Critical fluctuations near the critical point, Tc \rightarrow K/ pi, <pT>, etc

Hadronic mass/width changes (Pisarski 1982) → rho etc

Charmonia suppression (Satz, Matsui 1987) \rightarrow T(dissociation) of ccbar, bbbar

Jet quenching (J D Bjorken 1982) → medium density

--> Goal is to achieve a combination of many signatures

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Onset of QCD phase transition at mu_B=0 around ~ 0.8 GeV/fm³

S.K., Peter Minkowski, 2001 New J. Phys. 3 4



systems exhibit universality, versus the Bjorken energy density.

The onset of saturation, approx 0.8 GeV/fm³ reveals the onset of the QCD phase transition (vgl Van Hove's signature)

Sonia Kabana, Lecture on Strangeness in HI Collisions, Dubna, Russia, 20-31 August 2018

RHIC PHENIX: Direct photon excess in min bias Au+Au at 200



Confirmed also with other measurement method : PHENIX 1405.3940, published in PRC 91 (2015) 064904

Direct photons in p+p described by NLO

Direct photon excess in min. bias Au+Au at 200 GeV over p+p at 200 GeV below pT ~2.5 GeV

Exponential spectrum in Au+Au - consistent with thermal below pT ~2.5 GeV with inverse slope 220 ± 20 MeV --> T(init) from hydrodynamic models : 300-600 MeV, depending on thermalization time

Critical d+Au check : No exponential excess in d+Au

RHIC Theory on direct photons

C. Gale et al, 1308.2440



The 3rd dimension in these plots is cross section of photons

 dN^{γ} / dydTd au dN^{γ} / dy

T(init) ~350 MeV -> T(observed)=260 MeV

Number of Constituent Quark scaling

Number-of-Constituent-Quarks scaling Au+Au at sqrt(s)=200 GeV 1701.06060, STAR

Mass ordering

NCQ scaling



Strange hadrons as well as D0 follow the Number-of-Constituent-Quarks scaling

- -> suggest hadronization out of partonic matter
- -> Thermalization of charmed mesons?

-> Production of hadrons mainly via

Tecomologication of Dartons Sonia Kabana, Heavy Ions in non-LHC experiments, 4-5 December 2018, Louvain La Neuve, Belgium 23

Number of constituent quark scaling seen also in small systems: 3He+Au



The familiar behavior of number of quark scaling observed in Au+Au collisions is also seen in the small ³He+Au system

Heavy Ion Collisions: dominant hadron production mechanism via quark coalescence

- The "Number of Constituent Quark Scaling" suggests dominal hadron production via quark coalescence out of a hadronizing QGP.
- Other data like particle yields and ratios agree also with production via coalescence

- As a consequence the hot and dense partonic matter produced in heavy ion collisions is expected to be a good source of multipartonic states like : pentaquarks, tetraquarks, dibaryons, strangelets, glueballs and hybrids

since the coalescence production mechanism is expected to be enhanced in the QGP and can dominate the overall cross section

Antihypernuclei

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RHIC antihypertriton

Hypertriton: consists of a Lambda, a proton and a neutron, was discovered in 1952. However no antihypernuclei were observed, until STAR.



- ★ Anti-hypertriton: anti-proton, anti-neutron & anti-Λ the first antinucleus with strangeness, and the heaviest antinucleus so far.
- After searching >100 million AuAu collisions, found 70 antihypertritons.
- Published in Science in March 2010; News stories in Nature, Scientific American, National Geographic, many news outlets worldwide.

Sonia Kabana, Lecture on Strangeness in hi Comsions, Dubna, Kussia, 20-51 August 2016

STAR antihyernucleus



Sonia Kabana, Lecture on Strangeness in HI Collisions, Dubna, Russia, 20-31 August 2018

Hypertriton and Lambda lifetime STAR, 2010 paper



STAR Hypertriton lifetime=

 $182 \pm {}^{89}_{45} \pm 27 \text{ ps}$

Lambda->protor lifetime PDG=

 $\tau = 263 \pm 2 \text{ ps}$

2010: lifetime within uncertainties consistent with Lambda to proton + pi- lifetime

Sonia Kabana, Lecture on Strangeness in HI Collisions, Dubna, Russia, 20-31 August 2018

Hypertriton and Lambda lifetime

Phys.Rev. C97 (2018) no.5, 054909



STAR Hypertriton and Lambda lifetime

Phys.Rev. C97 (2018) no.5, 054909



STAR Antihypernucleus



$$\frac{{}^{3}_{\Lambda}H}{{}^{3}\text{He}\times\frac{\Lambda}{p}}$$



Ratio is close to unity, and is larger than at lower beam energy, suggesting an equilibrium in coordinate and momentum space populations of up down and strange quarks and antiquarks

3-D chart of the Nuclides



3-D chart of the Nuclides



3-D chart of the Nuclides



Phase diagram with Strangelets and hypernuclei



Int J of Mod Phys 1996, W. Greiner

Sonia Kabana, Lecture on Strangeness in HI Collisions, Dubna, Russia, 20-31 August 2018
STAR: Observation of the antimatter helium-4 nucleus Nature 473 (2011) 353



STAR: Observation of the antimatter helium-4 nucleus



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NA52 Nuclei and antinuclei PbPb at sqrt(s)=17 GeV NA52, New J.Phys. 5 (2003) 150



Data consistent with nuclei and antinuclei COales C. B. R. C. B. L. C. S. C. B. L. B. C. Belgium 39

Dark Photon Searches

PHENIX dark photon

PHENIX Dark Photon Search PHENIX 1409.0851

A number of deviations between experimental data and the Standard Model like

- * the 3.6 sigma deviation of (g-2) of the muon
- * the positron excess in cosmic rays observed in ATIC, PAMELA, and AMS-2

can be attributed to the existence of a "dark photon", which is weakly coupled to the QED photon.

For m(DP)=m(DarkPhoton)>2m(electron) and < 2m(muon) decay into e+e- can be considered PHENIX experiment at RHIC searhed for the Dark Photon in the channel

$$\pi^0, \eta \to \gamma U, U \to e^+ e^-$$

PHENIX 1409.085 PHENIX Dark Photon Search



PHENIX e+e- invariant mass of data with fit describing known sources of e+e-

PHENIX 1409.085



2-3 10⁻⁶ Upper limit derived on the square of the coupling of dark photon to Standard Model photon epsilon²

PHENIX Dark Photon Search



Comparison to ALICE and other data Taku Gunji et al, ALICE collaboration



ALICE: Similar ε² at 90% CL compared to other experiments

Worse than NA48-2 and BaBar

Comparison. HI data: ALICE, PHENIX, HADES Dark photon to e+e-, is excluded as source of the muon g-2 anomaly but it may be source of positron excess in space or not have e+e- as final state

ALICE simulated future feasibility for dark photon

Expected improvement by an order of magnitude



Better statistics and resolution due to ALICE upgrades

Mass range 20-90 MeV

Taku Gunji et al, ALICE

HADES (GSI) dark photon

Search for dark photon in p (3.5 GeV) +p, Nb and in Ar (1.756 GeV)+KCI collisions in e+e- final state

$$\pi^0 \to \gamma U, \eta \to \gamma U, \text{ and } \Delta \to NU$$

 $U \to e^+ e^-.$

They set upper limit at 90% CL in the kinetic mixing parameter squared, for the mass range for dark photon M(U) =0.02 - 0.55 GeV



http://inspirehep.net/record/1333078/files/ epjconf_meson2014_03006.pdf

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First measurement of the Vorticity of QGP

First Vorticity measurement in AuAu 200 GeV 20-50% centrality STAR, Nature, 2017, 1701.06657



Average vorticity points towards the direction of the angular momentum J(sys) of the collision.

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First Vorticity measurement in AuAu 200 GeV 20-50% centrality STAR, Nature, 2017, 1701.06657



$$\frac{dN}{d\cos\theta^*} = \frac{1}{2} \left(1 + \alpha_{\rm H} |\vec{\mathcal{P}}_{\rm H}| \cos\theta^* \right).$$

H: Lambda/Anti-Lambda

P_{H :} Lambda/AntiL polarizatin vector in the hyperon rest frame

decay parameter $\alpha_{\Lambda} = -\alpha_{\overline{\Lambda}} = 0.642 \pm 0.013$

Average projection of the Polarization on J(sys) is extracted:

$$\overline{\mathcal{P}}_{\mathrm{H}} \equiv \langle \vec{\mathcal{P}}_{\mathrm{H}} \cdot \hat{J}_{\mathrm{sys}} \rangle = \frac{8}{\pi \alpha_{\mathrm{H}}} \frac{\left\langle \cos\left(\phi_{p}^{*} - \phi_{\hat{J}_{\mathrm{sys}}}\right) \right\rangle}{R_{\mathrm{EP}}^{(1)}},$$

noted as "global polarization"

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sQGP vorticity measured

P_H: average polarization with H: Lambda or Antilambda



STAR, Nature, 2017, 1701.06657

Measurement of vorticity in Au+Au collisions with 20-50% centrality via the average polarization of Lambda and Antilambda.

Fluid vorticity can be calculated using the hydrodynamic relation (Becatini et al 1610.02506.)

$$\boldsymbol{\omega} = k_B T \left(\overline{\mathcal{P}}_{\Lambda'} + \overline{\mathcal{P}}_{\overline{\Lambda}'} \right) / \hbar,$$

With T the temperature. The vorticity found is omega = (9+-1) 10²¹ s-1 with an additional systematic error of a factor of 2 which by far surpasses the vorticity of all known fluids

For example solar subsurface flow has omega= 10-7 s-1, and superfluid nanodroplets omega=10⁷ s-1

- * The Quark Gluon Plasma produced in heavy ion collisions is
- hotter
- least viscous
- and has larger vorticity,

from all fluids ever produced in the laboratory !

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New STAR results on global polarization of Lambda, Antilambda in Au+Au at 200 GeV 1805.04400



High precision measurement of a finite Lambda and Antilambda global polarization of the level of 0.1-0.5% (depending on centrality) in Au+Au at 200 GeV

Global polarization increases with decreasing collision energy

Exotic states

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Strange Quark Matter

Strange Quark Matter (SQM) is a hypothetical state of matter consisting of roughly equal numbers of u, d and s quarks.

Bodmer A R 1971 Phys. Rev. D 4 1601 Witten E 1984 Phys. Rev. D 30 272 Terazawa H 1979 INS Report 338 University of Tokyo, Terazawa H 1989 J. Phys. Soc. Japan 58 3555, Terazawa H 1989 J. Phys. Soc. Japan 58 4388, Terazawa H 1990 J. Phys. Soc. Japan 59 1199

Absolutely stable Strange Quark Matter ... is possible E Witten, 1984

For more than few hundreds u,d,s quarks, the energy per baryon (E/ A) of quark matter can be below the energy of the most stable atomic nucleus 56Fe, which has Energy per baryon number Mass(56Fe)c²/56 = 930.4 MeV

With E/A(strange quark matter)= 4Bpi²/mu³, the E/A of Strange Quark Matter can be

E/A = 829 MeV for bag constant B=57.5 MeV fm-3 (or B^(1/4)=145 MeV) and

E/A = 915 MeV for bag constant B=85,3 MeV fm-3 (B^{(1/4)=160} MeV).

In this cases Strange Quark Matter from u,d,s quarks would be the ground state of matter.

From Strangelets to Strange Stars

One may expect novel states to exist like strange nuggets or strangelets (for small A like A~10-100) to strange stars

Experimental searches for strangelets with small A In accelerators (AGS BNL, NA52 at CERN SPS, etc) In cosmic rays with detectors on earth (centauro events)

In Space (AMS)

Upper Limits at 90% CL on strangelet production from SPS NA52 and BNL heavy ion experiments E858, E878, E864



Soma Kabana, Heavy ions in non-LHC experiments, 4-5 December 2016, Louvain La Neuve, Beigium

STAR Strangelet Search at the BNL Relativistic Heavy Ion STAR Phys. Rev. C76 (2007) 011901

Strangelets in forward region are expected to have small charge to mass ratio and large rigidity (p/Z). They are expected to deposit energy in the forward calorimeter STAR is sensible to strangelets with lifetime > **0.1 nsec**



Upper limits of a few 10–6 to 10–7 per central Au +Au collision are set for strangelets with mass > 30 GeV.

STAR Strangelet Search at the BNL Relativistic Heavy Ion STAR Phys. Rev. C76 (2007) 011901

Search done in 61 million central Au+Au events

strangelet mass > 30 GeV.

lifetime > 0.1 nsec



Pentaquark searches

Search for the anti-Theta- ---> K- anti-n with PHENIX

J.Phys. G30 (2004) no.8, S1201-S1206



Figure 4. $K^+ \overline{n}$ Invariant mass distribution. No peak is expected in the mass range between 1.5 GeV/c^2 and 1.6 GeV/c^2 . It serves as a crosscheck for technical problems and particle misidentifications.

Search for the anti-Theta- ---> K- anti-n with PHENIX J.Phys. G30 (2004) no.8, S1201-S1206

Conf proceeding



Charged kaons can be identified using the standard tracking and time of flight up to a momentum of 1.5 GeV/ Anti–neutron candidates are detected via their annihilation signal in the highly segmented electromagnetic calorimeter (EMCal).

No anti-Theta-signal observed (no upper limits extracted).

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STAR

Search for theta+ ->K0s p

S. Salur et al STAR coll., poster QM2004







Theta++ searches in STAR $_{d+Au \ 200 \ GeV} \theta^{++} \rightarrow p+K^+$

STAR, Huan Huang et al nucl-ex/0509037



Fig. 1. Invariant mass distributions of $pK^+ + \overline{p}K^-$ (left) and $pK^- + \overline{p}K^+$ (right) from d+Au collisions at 200 GeV. The histograms are combinatorial backgrounds from event-mixing technique.



Left plot theta++ candidate peak at mass=1528+-2 MeV and significance of 4.2 sigma Width is corresonding to detector resilution Peak at 1460 MeV is due to missidentified pion as Kaon from Delta decays.

Theta++ searches in STAR, Huan Huang et al nucl-ex/0509037

 $\theta^{++} \to p + K^+$

dE/dx in the TPC: pi/kaon separation possible up to 0.6 GeV/c and pi/proton separation up to 1.0 GeV/c Au+Au collisions 200 GeV



Fig. 3. Invariant mass distributions of $pK^+ + \overline{p}K^-$ from Au+Au collisions at 62.4 GeV (left) and at 200 GeV (right) after subtraction of combinatorial background. The curves are a fit of Gaussian and polynomial function.

HADES search for a doubly charged dibaryon 1410.8004

Sigma+=uus y++ (uusuud)

$$pp \to \mathcal{Y}^{++} K^0 \to \Sigma^+ p K^0 \to n \pi^+ p K^0$$

Measurement: proton, pi +,pi-,pi+

Namely pi+pi- from K0s pi+ from the Sigma+ decay into neutron + pi+ and proton associated with Sigma

No pi+enhancement seen in the missing mass spectrum



QGP: a Glueball production Fabrik?

Glueballs from QGP



Peter Minkowski, Wolfgang Ochs, Eur.Phys.J. C9 (1999) 283-312

"Red Dragon" decay into pipi, KK, eta eta



name	PDG	mass (MeV)	${ m mass}^2 ({ m GeV})^2$	width (MeV)
$gb \;(\; 0^{++}\;)$	$f_0(400 - 1200)$	~ 1000	$\sim 1.$	500 - 1000
	$f_0(1300)$			
$gb~(~0^{-+}~)$	$\eta(1440)$	1400 - 1470	2.07	50 - 80
$gb~(~2^{++}~)$	$f_{J}(1710)$	$1712~\pm~5$	2.93	$133~\pm~14$

Table 15: Properties of the basic triplet of binary glueballs gb.

Glueballs from QGP



"Red Dragon" decay into pi+pi-

Peter Minkowski, Wolfgang Ochs Eur.Phys.J. C9 (1999) 283-312
Glueballs from QGP SK P. Minkowski, Phys.Lett. B472 (2000) 155-160



CERES data e+e-



Glueballs from QGP

SK P. Minkowski, Phys.Lett. B472 (2000) 155-160

This analysis could explain CERES excess data by a " sigma " -assumed as a 0++ glueball- enhancement

We proposed an enhancement of Glueball as a signature o QGP and we suggested search for glueballs in Heavy lon collisions

Conclusions and perspectives

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sQGP

- * The Quark Gluon Plasma produced in heavy ion collisions is an extraordinary state of matter :
- 100,000 times hotter than the core of the sun (T ~ 200 MeV, 2.3x10^{{12}} K).
- least viscous than any observed fluid -> the Perfect Fluid
- with larger vorticity by many orders of magnitude than all fluids produced in the laboratory
- * Very high density of partons -> hadrons produced
- * Very large magnetic field

That may be relevant for new particle and new phenomena searches. in Heavy Ion Collisions

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STAR future plans



Beam Energy Scan (BES) II 2019-2020 Will continue the BES I program "Hot" QCD, search for a possible critical point and discontinuities in the energy dependence of QGP signatures -> FAIR and NICA

STAR forward rapidity program (2.5-eta-4): Hcal, Ecal, tracki (Silicon and sTGCs) "Cold" QCD, Proton TMDs, gluon saturation Test Electron Ion Colider (EIC) detector technologies Milestone: 2021 p+p run and sPHENIX data taking 2022+ -> EIC





STAR forward rapidity program





3 Silicon discs

4 Small-strip Thin Gap Chambers

ECal: use upgraded PHENIX PbSc calorimeter

HCal: Iron-scintillator



STAR upgrades





iTPC: inner sector of TPC. Extends peudorapidity acceptance from 1 to 1.5. Improves dE/dx

Endcap TOF: particle identification 0.9-eta-1.5

Event Plane Detector: will provide better and independent determination of centrality and event plane

Sonia Kabana, Selected Highlights from the STAR experiment at RHIC, ICNFP 2018, Crete, Greece 79

STAR goals

BES-II

Beam Energy	$\sqrt{s_{NN}}$ (GeV)	$\mu_{\rm B} ({\rm MeV})$	Run Time	Number Events
(GeV/nucleon)				
9.8	19.6	205	4.5 weeks	400M
7.3	14.5	260	5.5 weeks	$300\mathrm{M}$
5.75	11.5	315	5 weeks	$230\mathrm{M}$
4.55	9.1	370	9.5 weeks	$160\mathrm{M}$
3.85	7.7	420	12 weeks	$100\mathrm{M}$
31.2	7.7 (FXT)	420	2 days	100M
19.5	6.2 (FXT)	487	2 days	$100\mathrm{M}$
13.5	5.2 (FXT)	541	2 days	$100\mathrm{M}$
9.8	4.5 (FXT)	589	2 days	$100\mathrm{M}$
7.3	3.9 (FXT)	633	2 days	$100\mathrm{M}$
5.75	3.5 (FXT)	666	$2 \mathrm{days}$	$100\mathrm{M}$
4.55	3.2 (FXT)	699	2 days	$100\mathrm{M}$
3.85 SK fo	or the STAR C	ollaboratio	n. PCNFP2	D18 ^{100M}

STAR BES-II goals

Table 8: Event statistics (in millions) needed in BES-II for various observables. This table update estimates originally documented in Ref. [45].

Collision Energy (GeV)	7.7	9.1	11.5	14.5	19.6
μ_B (MeV) in 0-5% central collisions		370	315	260	205
Observables					
R_{CP} up to $p_T=5~{ m GeV}/c$	-		160	125	92
Elliptic Flow (ϕ mesons)	80	120	160	160	320
Chiral Magnetic Effect	50	50	50	50	50
Directed Flow (protons)	20	30	35	45	50
Azimuthal Femtoscopy (protons)	35	40	50	65	80
Net-Proton Kurtosis	70	85	100	170	340
Dileptons	100	160	230	300	400
$>5\sigma$ Magnetic Field Significance	50	80	110	150	200
Required Number of Events		160	230	300	400

+100M for each FXT energy

Typically factor 20 more than for BES-I

* detector project at RHIC: sPHENIX

sPHENIX: start data taking 2022

Extended Calorimetry precision vertexing and tracking for jet quenching, charm, beauty

Will be able to search for dark photon with higher mass than PHENIX



M. Connors, Nucl.Phys. A967 (2017) 548-551

Energy scans with Heavy lons Future: BESII, NICA, FAIR, J-PARC



T. Sakaguchi, QM2017

Center of mass energy (sqrt(s)NN) of facilities for future heavy ion runs: FAIR: 2-6 (10) GeV, NICA: 4-11 GeV, RHIC: 7 (2.5) -200 GeV LHC: 2.76, 5 TeV, J-PARC: 1-10 GeV FCC (100 km circular ring, p+p at sqrt(s)=100 TeV, Pb+Pb at sqrt(s)=39 TeV) Sonia Kabana, Heavy lons in non-LHC experiments, 4-5 December 2018, Louvain La Neuve, Belgium

Conclusions and perspectives

Heavy lon collisions offer special conditions of hot and dense partonic matter with high multiplicity, high magnetic field and large vorticity

* promising for discovering new states that can benefit from these conditions

 Heavy ion data exhibit the "Number of Constituent Quark Scaling" suggesting dominant hadron production via quark coalescence out of a hadronizing QGP.

Major advantage of Heavy lons for multipartonic states is the production via colaescence

Conclusions and perspectives

- Absolutely stable Strange Quark Matter may be possible (E Witten).
- Strange Quark Matter can exist in the core of neutron stars or quark stars
- Small droplets of Strange Quark Matter can be produced in particle collisions in accelerators and are searched by experiments as well as by space experiments like AMS.
- Exotic hadrons with strangeness like antihypernuclei and the anti-helium-4 have been observed in heavy ion collisions.
- Search for a Dark Photon in p+p,p+A and A+A collisions have set upper limits

Perspectives non-LHC STAR - RHIC BESII (2019-2020) and fixed target

- sPHENIX (2020+)
- NA61 (SPS)
- New dedicated accelerator facilities and corresponding new experiments
- NICA in Dubna, Russia
- FAIR in GSI, Germany
- J-PARC in Japan,

- will allow to progress in significant way in the next decades.

Thank you very much

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NICA



Flow coefficients v_n, n=1,2,3..



n = 2

n = 3



Matter in the overlapp area of two colliding nuclei gets compressed and heated Initial anisotropy gets transfered into the momentum space via pressure gradients

$$\frac{dN}{d\phi} \propto 1 + 2\sum_{n=1}^{\infty} v_n \cos[n(\phi - \Phi_n)]$$
$$v_n = < \cos[n(\phi - \Phi_n)] >$$

n = 4

n = 5

v : flow coefficients(v1: directed flow,v2: elliptic flow, ...)

Higher harmonics

n = 6

LHC experimental upgrades ALICE upgrades for run-3

GEM-TPC

ITS

MFT



MFT: will provide secondary vertex reconstruction in forward rapidity ITS : low pT reach and improved accuracy High rate The Upgrades will allow high statistics and high resolution strangeness measurements

Sonia Kabana, Heavy lons in non-LHC experiments, 4-5 December 2018, Louvain La Neuve, Belgium 91

A view into the far Future : FCC



FCC: The Vision

~100 km tunnel, 16 T magnets sqrt(s)= 100 TeV pp collisions

FCC-hh FCC-ee FCC-he

Possible first steps *FCC-ee, E_CM=90-400 GeV **HE-LHC* 16T 28 TeV in LEP/LHC tunnel

FCC-AA : sqrt(s)NN=40 TeV

Strangeness related possible Highlight of FCC ? : Strangelet production may be possible in FCC !



range of photon	fraction of tot	ion of total photon yield			
emission	AuAu@RHIC	PbPb@LHC			
	$0\mathchar`-20\%$ centr.	0-40% centr.			
$T=120\text{-}165\mathrm{MeV}$	17%	15%			
$T=165\text{-}250\mathrm{MeV}$	62%	53%			
$T>250{\rm MeV}$	21%	32%			
$\tau=0.6-2.0\mathrm{fm/c}$	28.5%	26%			
$\tau > 2.0{\rm fm/c}$	71.5%	74%			

C. Gale et al, 1308.2440

* Most photons at RHIC and LHC are emitted from time near Tc

* Their effective temperature is enhanced by strong radial flow (effective temperature of hadrons decaying into photons are above Tc due to mass dependence of radial flow).

* However a very high temperature early initial collision stage is required to generate this radial flow

Conclusions:

* Photons can be used as a thermometer

* T>Tc is reached

* More model calculations needed to fit the data and extract the T(init)