

# Simplified dark matter models with a spin-2 mediator at the LHC

Kimiko Yamashita (Ochanomizu Univ.)

Collaborators:

Sabine Kraml, Ursula Laa, Kentarou Mawatari  
(LPSC Grenoble)

Reference: arXiv:1701.07008

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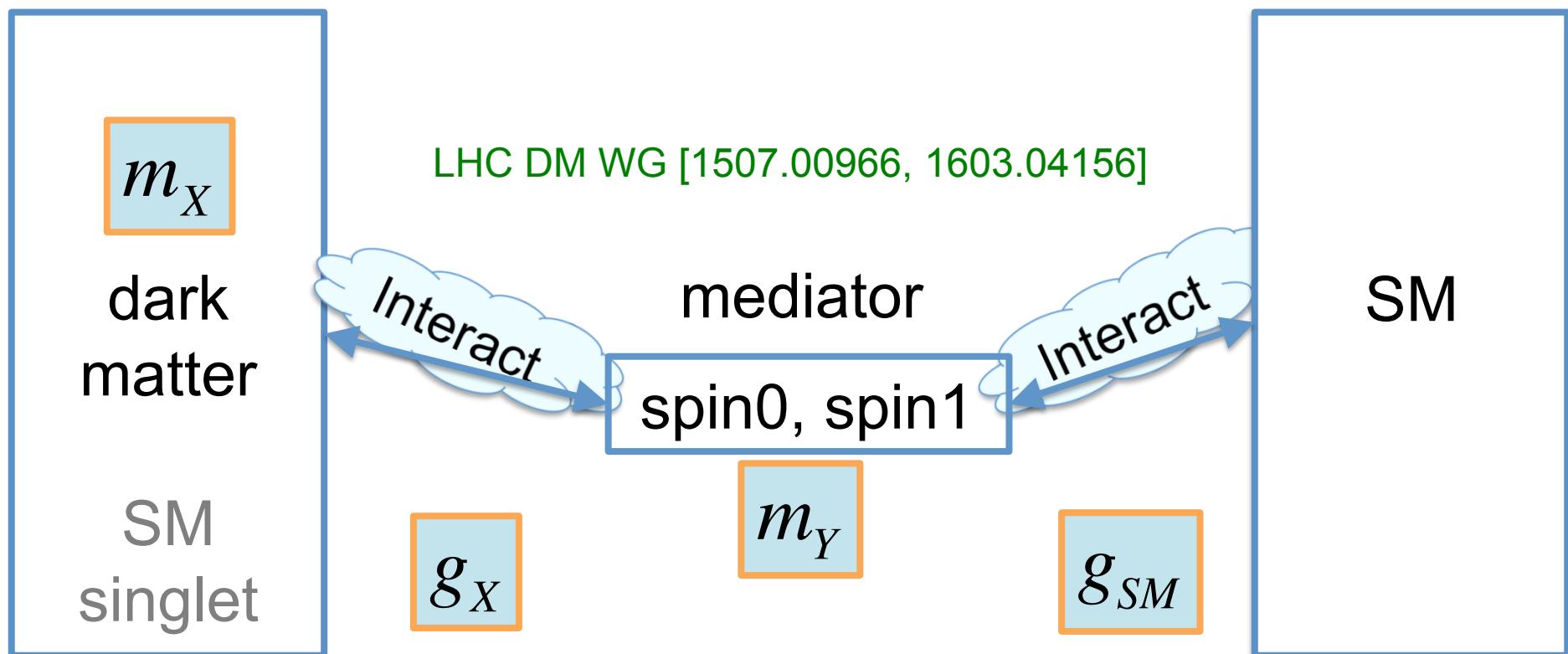
## 3. Result

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# s-channel simplified DM models

Model Parameter:  $m_X, m_Y, g_X, g_{SM}$



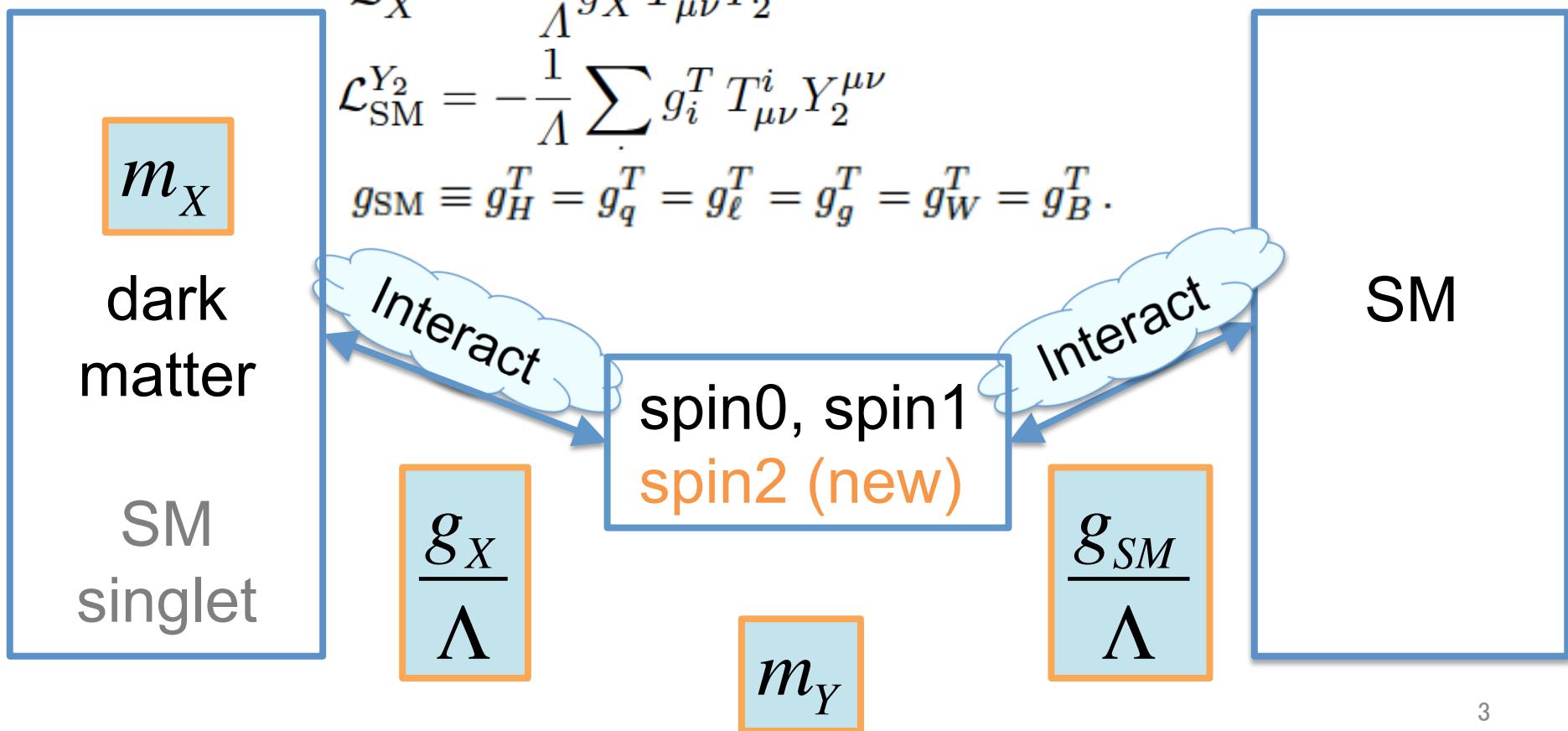
# spin2-mediated simplified DM model

Model Parameter:  $m_X, m_Y, \frac{g_X}{\Lambda}, \frac{g_{SM}}{\Lambda}$

$$\mathcal{L}_X^{Y_2} = -\frac{1}{\Lambda} g_X^T T_{\mu\nu}^X Y_2^{\mu\nu}$$

$$\mathcal{L}_{SM}^{Y_2} = -\frac{1}{\Lambda} \sum_i g_i^T T_{\mu\nu}^i Y_2^{\mu\nu}$$

$$g_{SM} \equiv g_H^T = g_q^T = g_\ell^T = g_g^T = g_W^T = g_B^T.$$



## Authors

# Simplified DM models (DMsimp)

- s-channel (spin-0 and spin-1)
  - Antony Martini (Université catholique de Louvain) & Kentarou Mawatari (LPSC Grenoble)
    - Emails: kentarou.mawatari @ ipsc.in2p3.fr
- s-channel (spin-0 and spin-1 electroweak)
  - Jian Wang (Johnnras Gutenberg University of Mainz) & Cen Zhang (Brookhaven National Laboratory)
    - Emails: cenzhang @ bnl.gov
- s-channel (spin-2)
  - Goutam Das (Saha Inst.), Celine Degrande (CERN) & Kentarou Mawatari (LPSC Grenoble)
    - Emails: goutam.das @ saha.ac.in, celine.degrande @ cern.ch, kentarou.mawatari @ ipsc.in2p3.fr

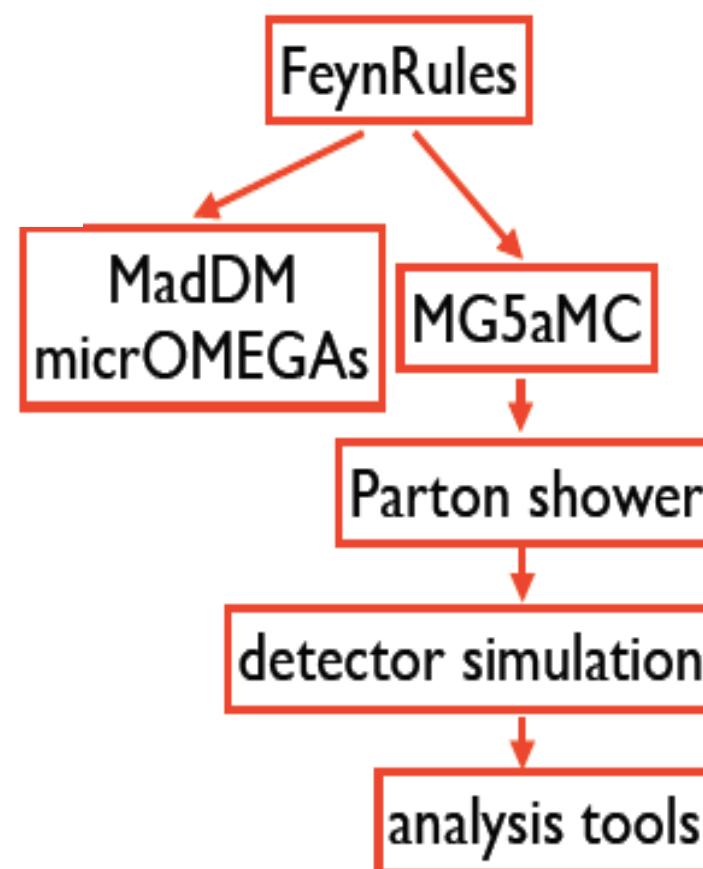
## Description of the model

This is simplified dark matter models for NLO.

- $X_r$  (real scalar DM)
- $X_c$  (complex scalar DM)
- $X_d$  (Dirac spinor DM)
- $X_m$  (Majorana spinor DM) [to be done.]
- $X_v$  (vector DM)
- ...

and different types of mediators:

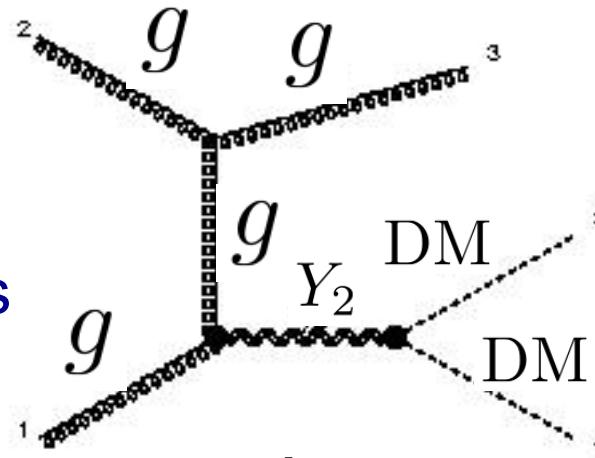
- s-channel
  - $Y_0$  (spin-0)
  - $Y_1$  (spin-1)
  - $Y_2$  (spin-2)



# LHC search

- monojet/multijet + missing  $E_T$  (DM search)

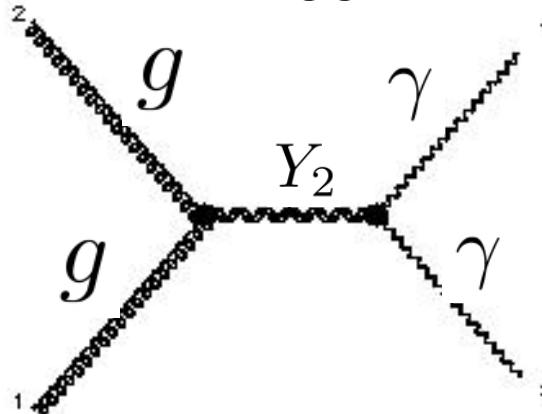
Y2 decay  
into DM/neutrinos



- resonance search

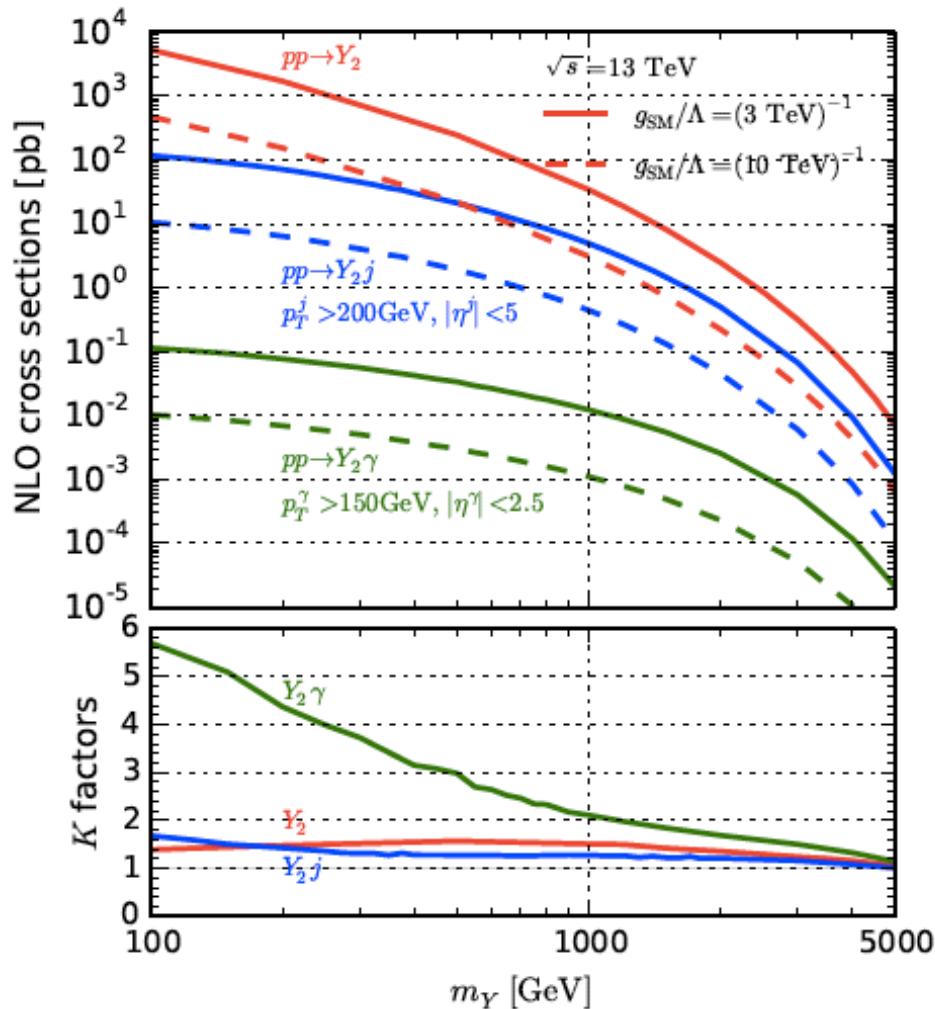
Resonance: 2 photons, 2 leptons, 2 jets, 2 jets(+j/gamma)  
2W, 2Z, 2 bottom, 2 higgs, 2 top

Y2 decay  
into SM



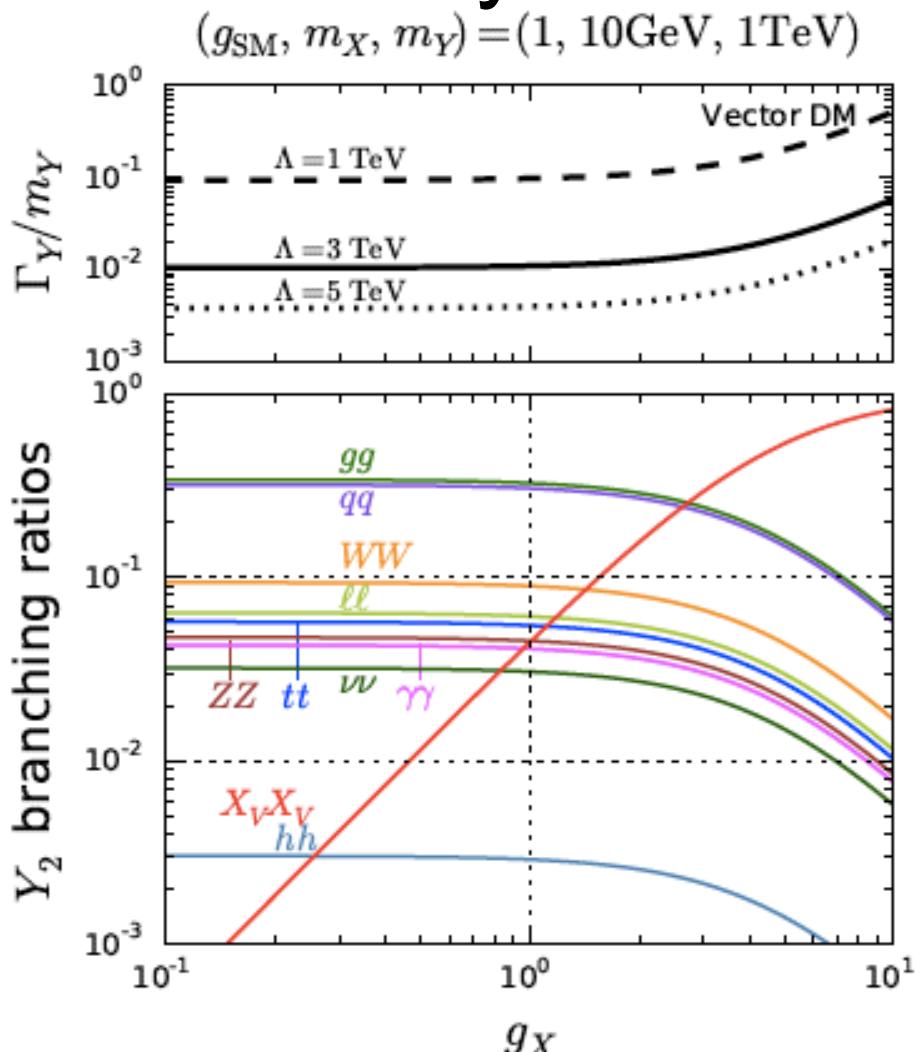
We used the NLO UFO model by Das, Degrande, Hirschi, Maltoni, Shao  
[1605.09359]

# $Y_2$ production and decay



$$\Gamma_S = \frac{g_S^2 m_Y^3}{960\pi\Lambda^2} \beta_S^5, \quad \Gamma_F = \frac{g_F^2 N_\nu N_C^F m_Y^3}{160\pi\Lambda^2} \beta_F^3 \left(1 + \frac{8}{3}r_F\right), \quad \Gamma_V = \frac{g_V^2 N_s N_C^V m_Y^3}{40\pi\Lambda^2} \beta_V f(r_V)$$

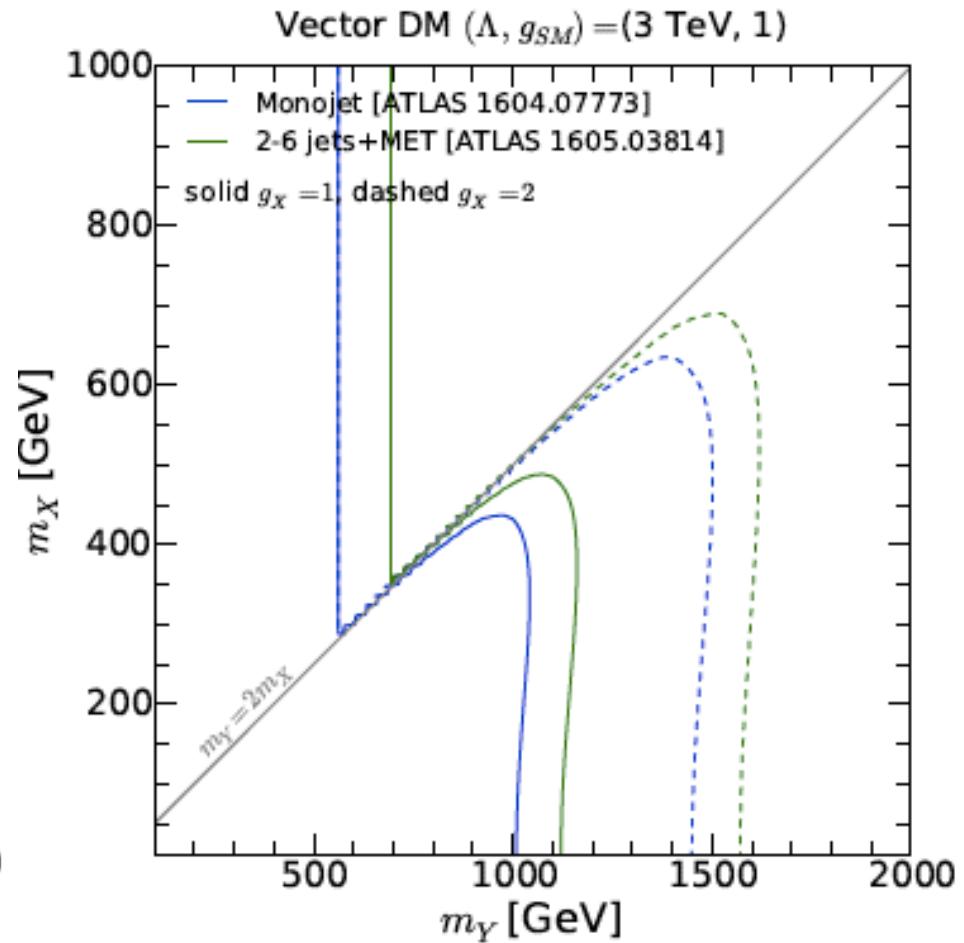
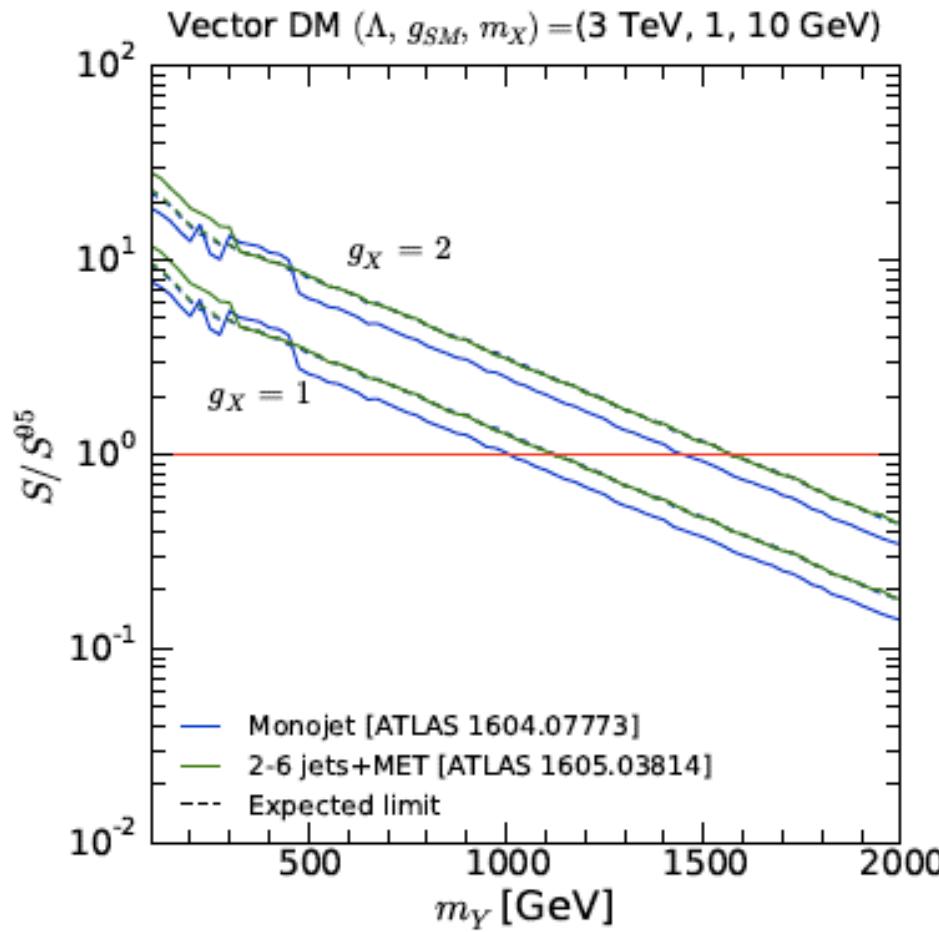
where  $\beta_i = \sqrt{1 - 4r_i}$  with  $r_i = m_i^2/m_Y^2$



by CheckMATE (validation: MadAnalysis5 and Rivet)

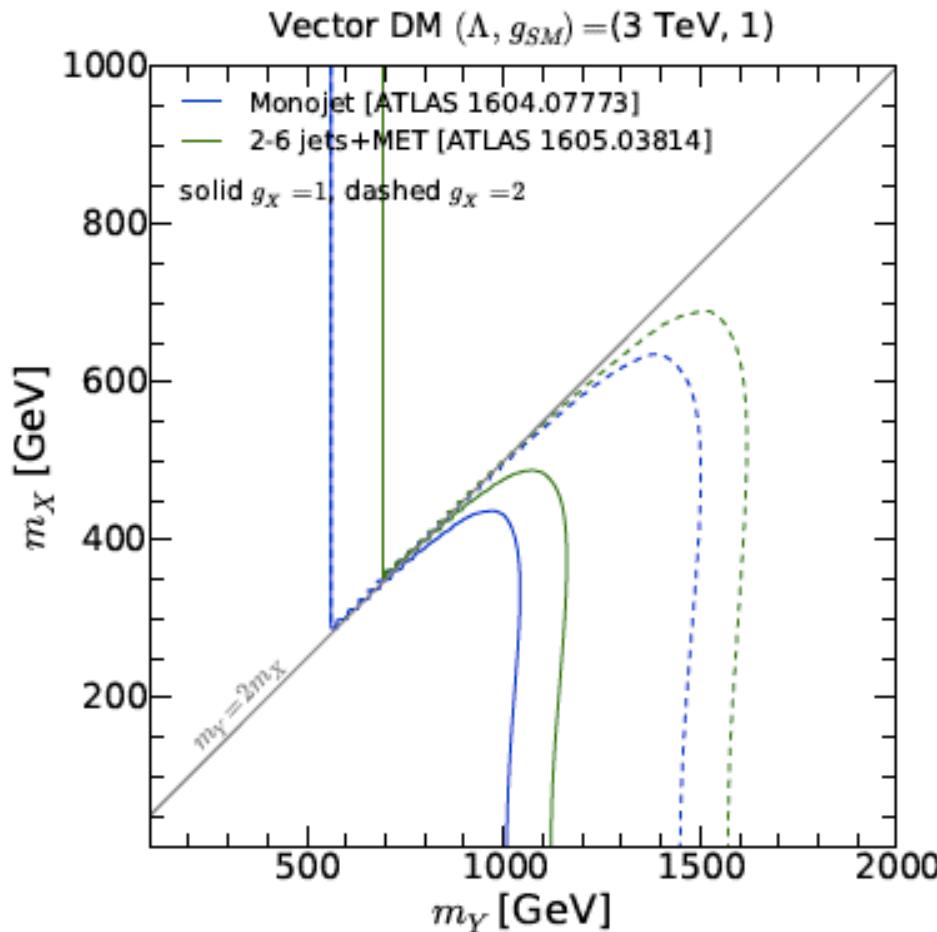
# Constraints from missing $E_T$ searches (1/2)

13 TeV LHC with 3.2/fb



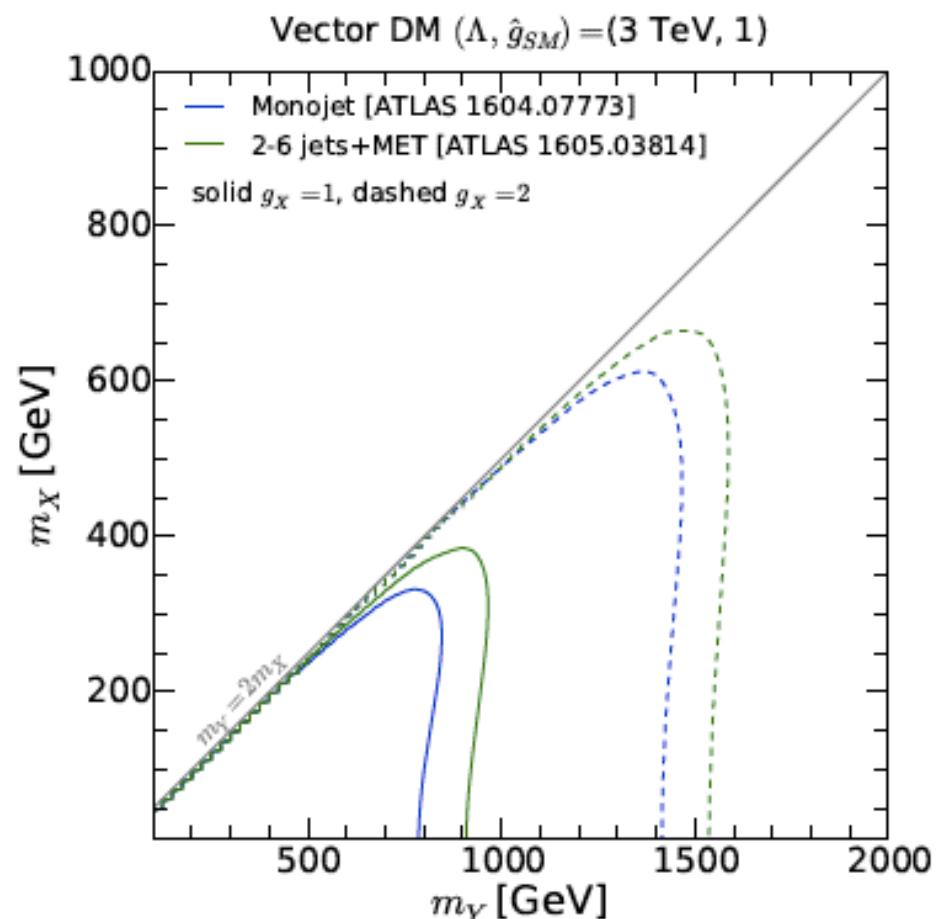
universal  $g_{SM}$        $g_{SM} \equiv g_H^T = g_q^T = g_\ell^T = g_g^T = g_W^T = g_B^T$ .

# Constraints from missing $E_T$ searches (2/2)



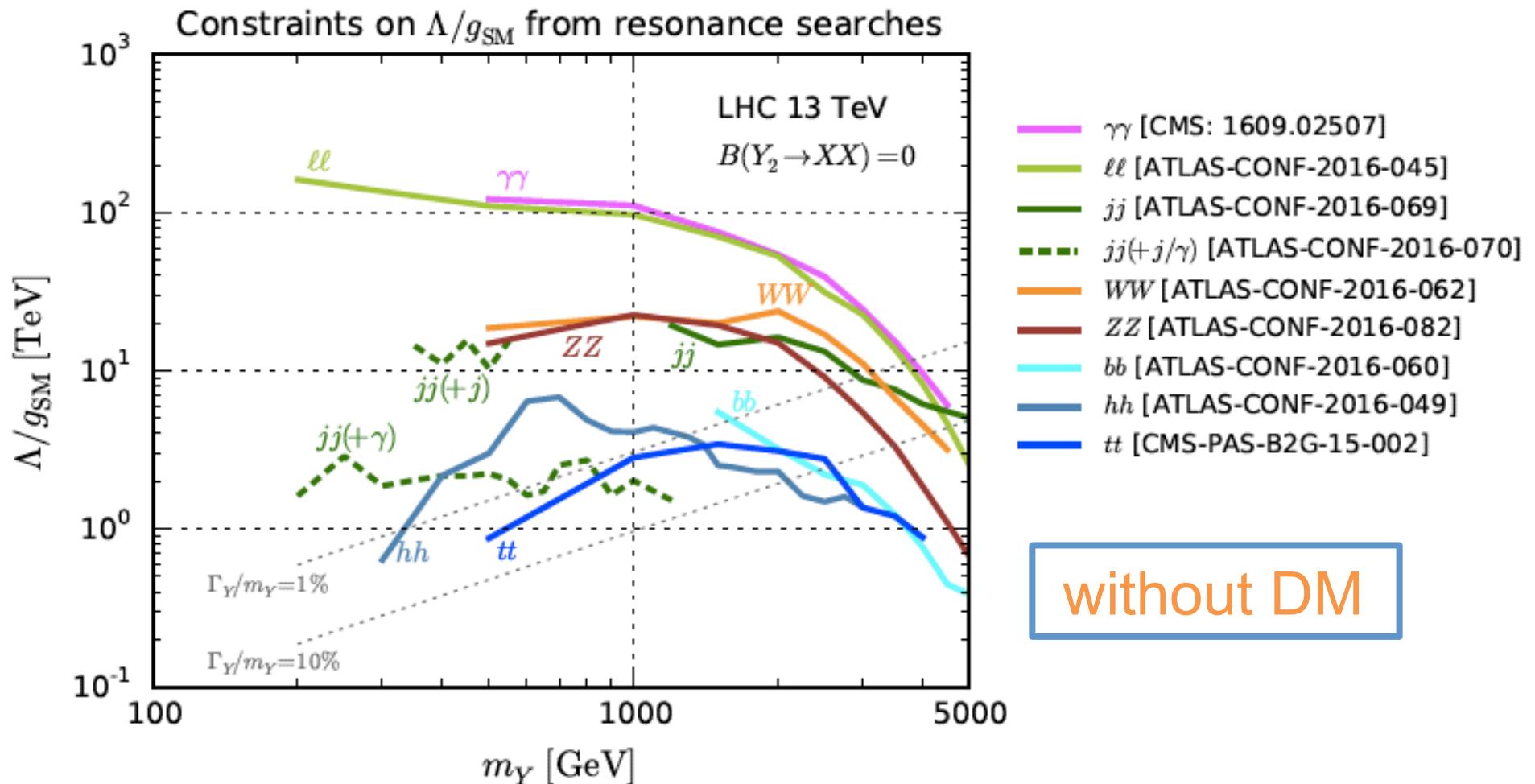
universal  $g_{SM}$

$$g_{SM} \equiv g_H^T = g_q^T = g_\ell^T = g_g^T = g_W^T = g_B^T. \quad g_\ell^T \ll \hat{g}_{SM} \equiv g_H^T = g_q^T = g_g^T = g_W^T = g_B^T$$

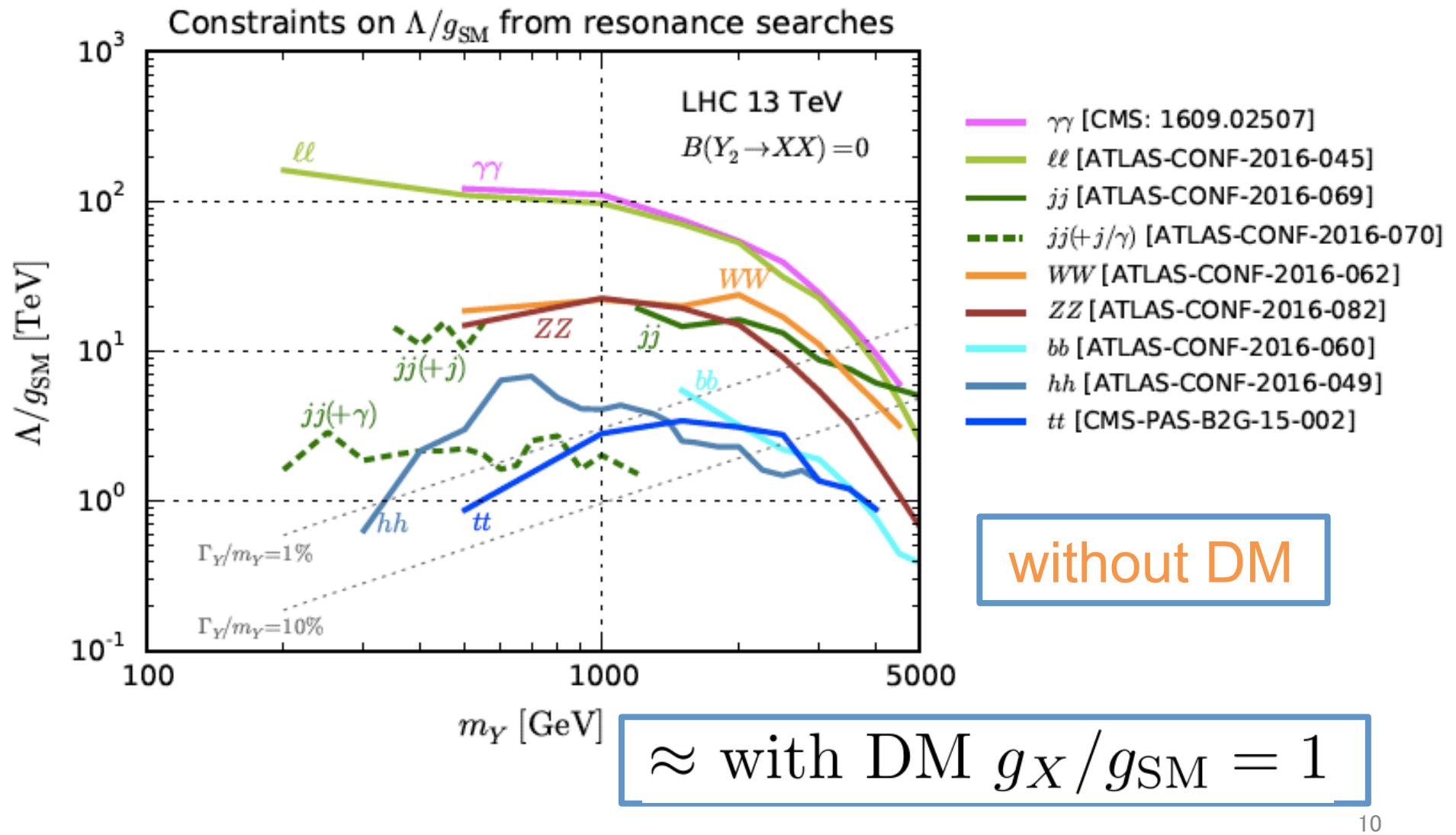


leptophobic

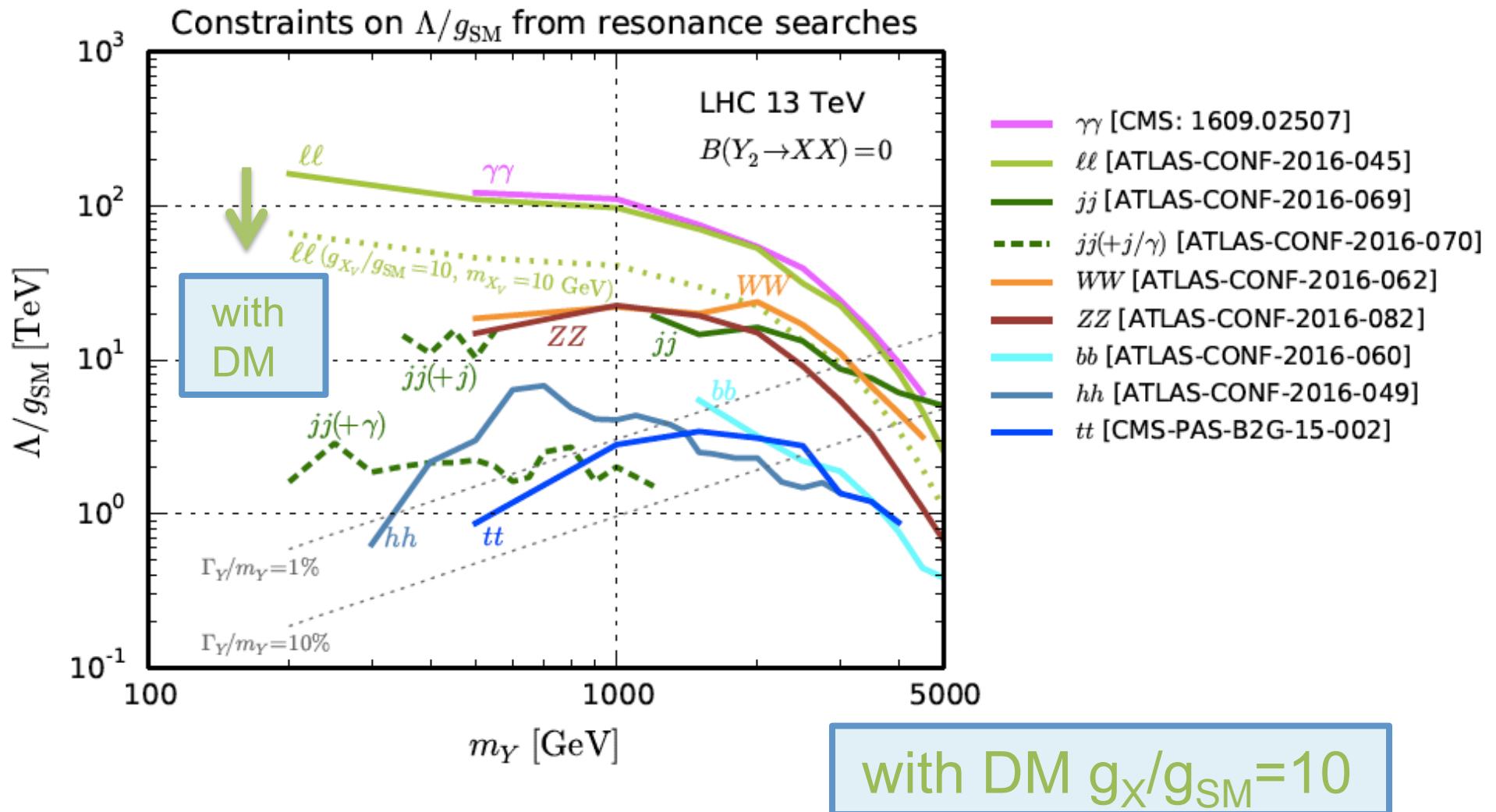
# Constraints from resonance searches (1/2)



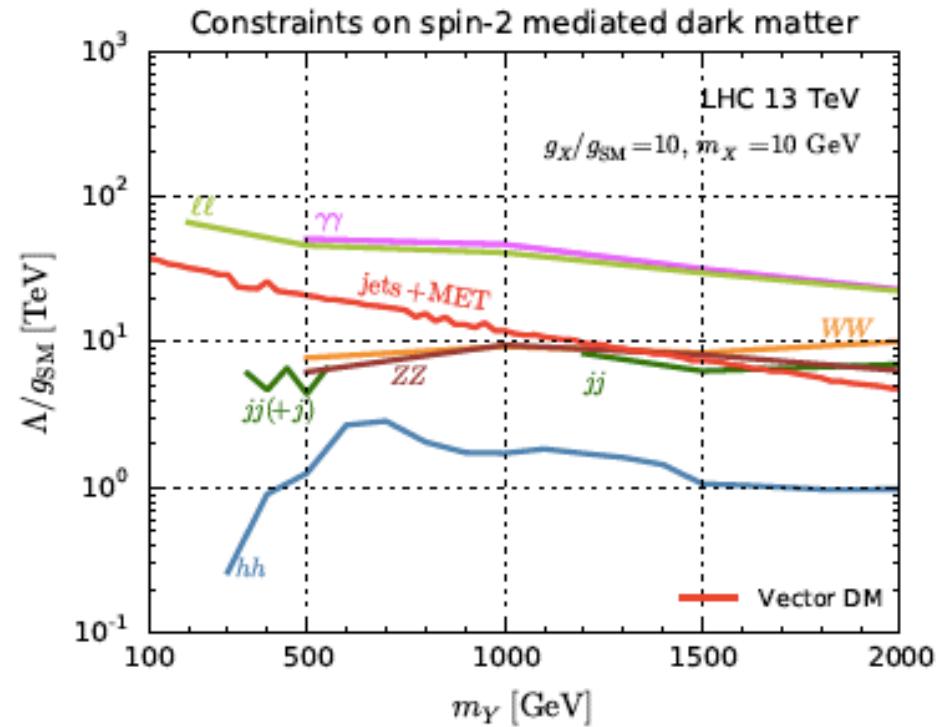
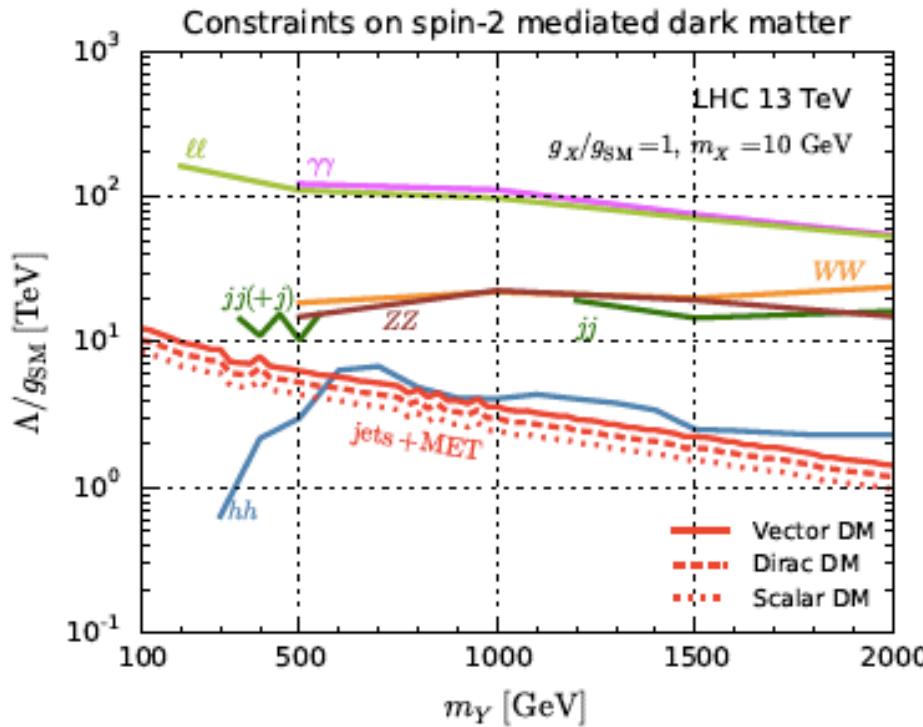
# Constraints from resonance searches (1/2)



# Constraints from resonance searches (2/2)



# Comparison between MET & resonance searches



$$g_X = g_{\text{SM}} = 1$$

$$\begin{aligned} g_X &= 10 \\ g_{\text{SM}} &= 1 \end{aligned}$$

# Summary (1/2)

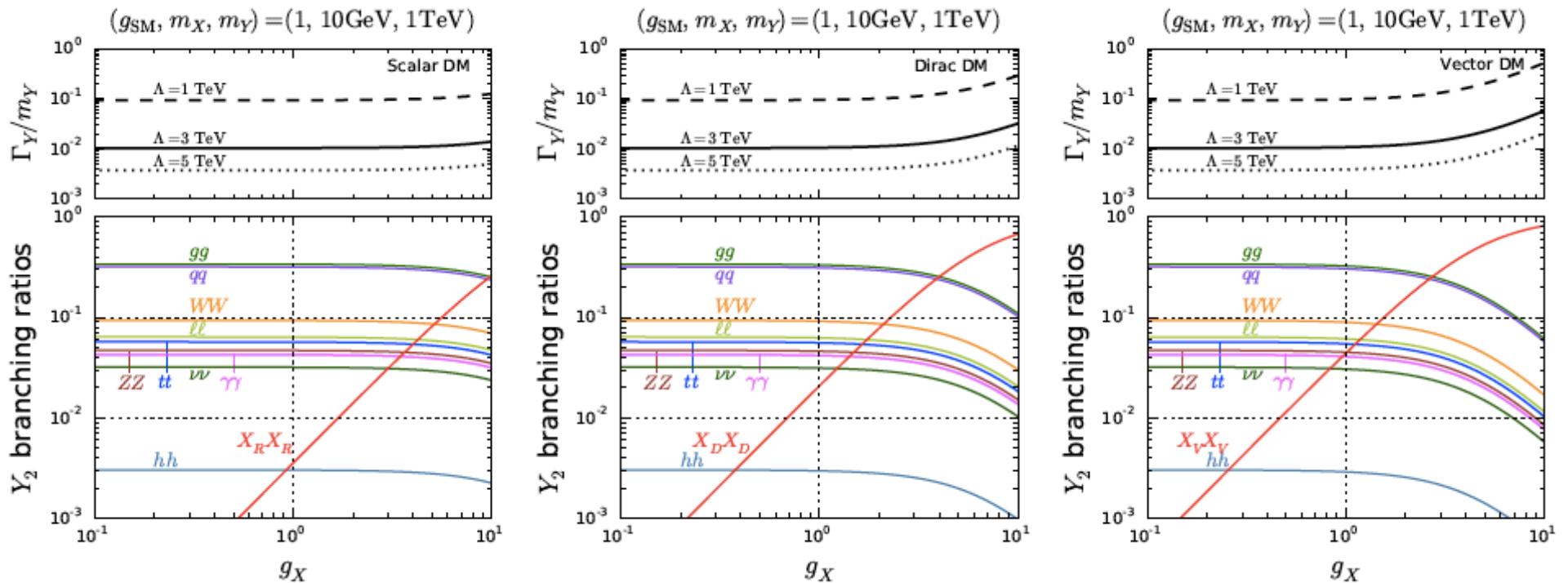
1. We considered spin2-mediated simplified DM Model, and studied the current LHC constraints:
  - monojet/multijet + missing  $E_T$
  - resonance searches
2. We show the complementarity among different searches
  - dilepton/diphoton resonance searches
    - > strongest constraint for  $m_{Y_2} > 200/500$  GeV
  - Missing  $E_T$  searches
    - > competitive only in the low-mass region
    - > leptophobic scenario/larger  $g_X$  (DM coupling) give important role for this searches

## Summary (2/2)

3. DMsimp s-channel spin- 0,1,2 NLO UFO models are publicly available :  
<http://feynrules.irmp.ucl.ac.be/wiki/DMsimp>
  
4. Efficiencies and MadGraph5 aMC@NLO & Pythia6 run cards for monojet and 2-6jets + MET searches, Digitized limits from resonance searches are publicly available :  
<http://ipsc.in2p3.fr/projects-th/recasting/spin2mediator/>

# Backup

# $Y_2$ decay: Scalar, Dirac, Vector DM

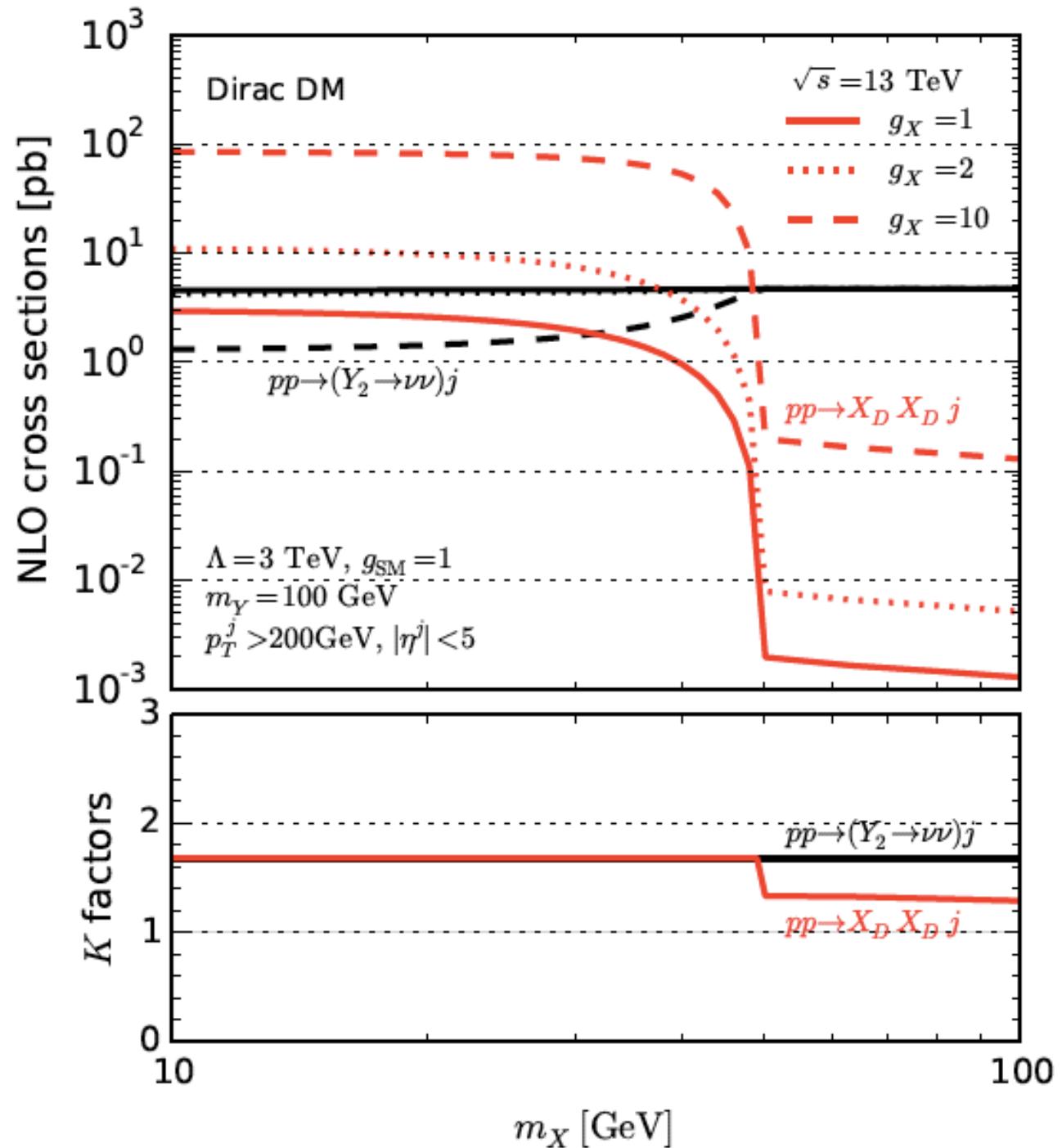


$$\Gamma_S = \frac{g_S^2 m_Y^3}{960\pi\Lambda^2} \beta_S^5,$$

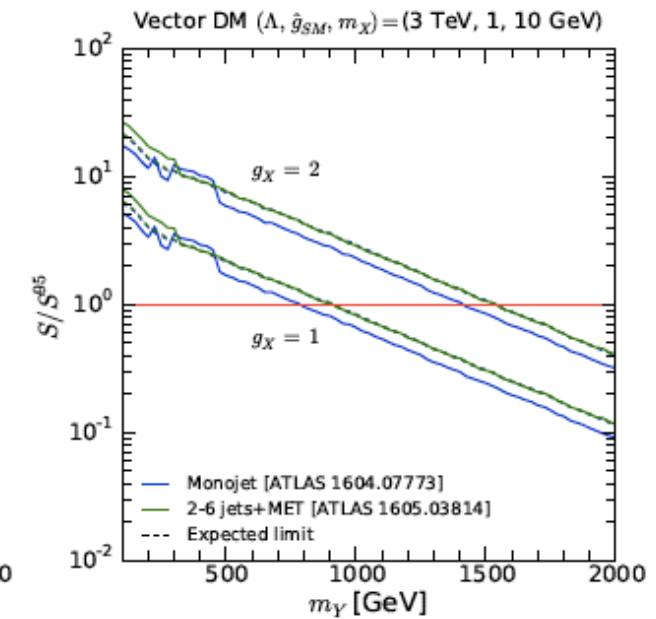
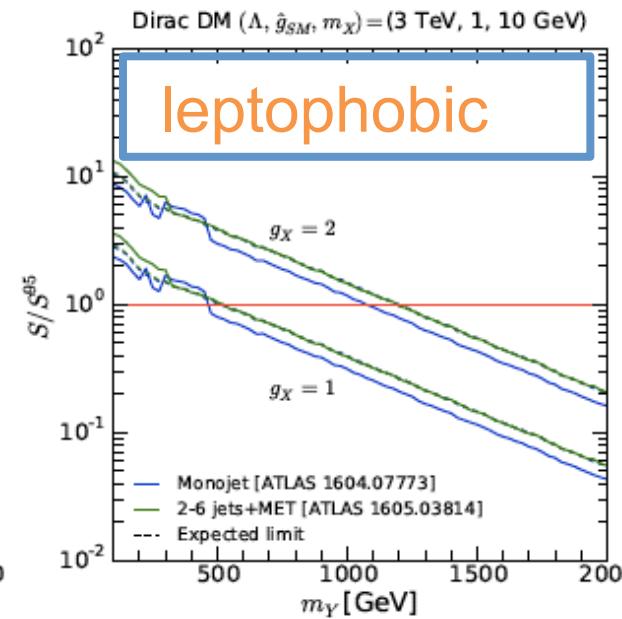
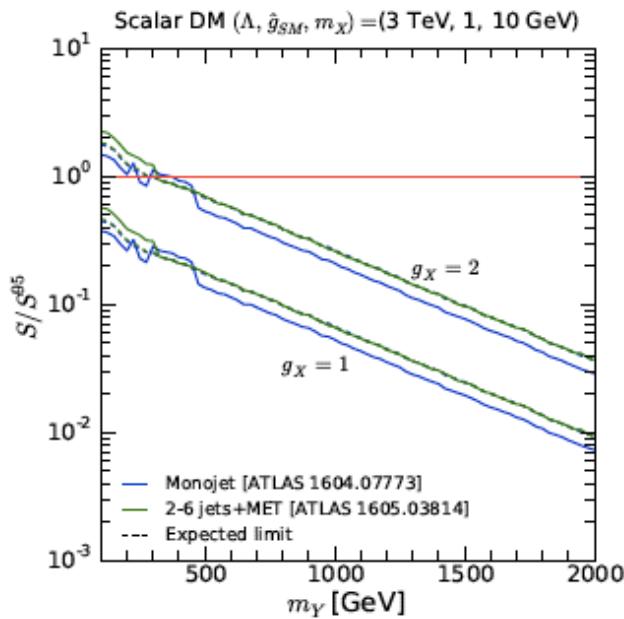
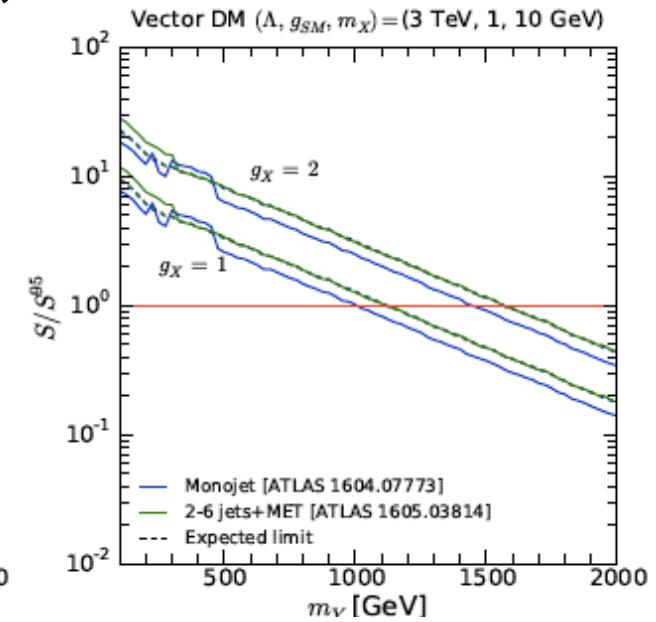
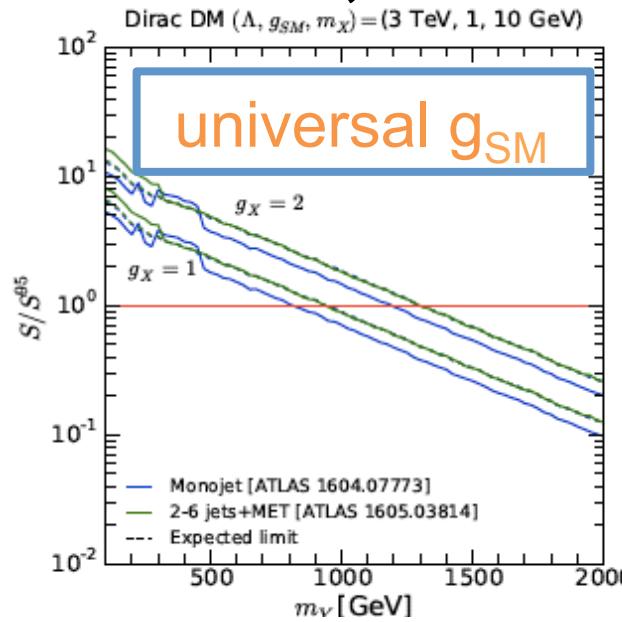
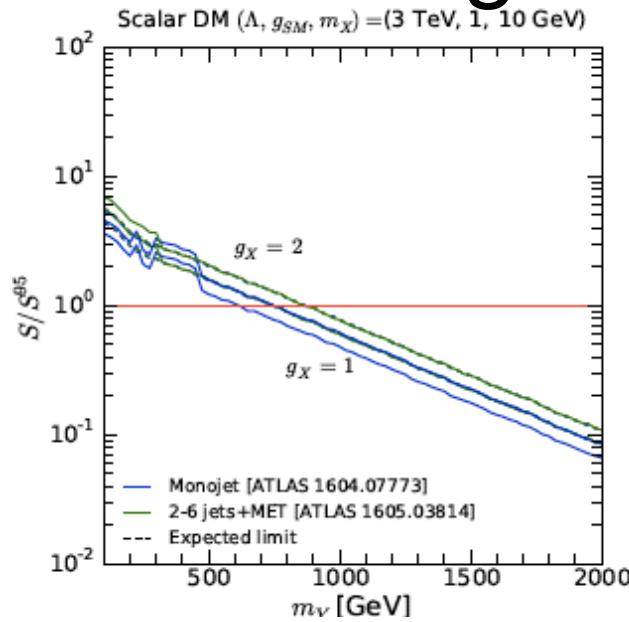
$$\Gamma_F = \frac{g_F^2 N_\nu N_C^F m_Y^3}{160\pi\Lambda^2} \beta_F^3 \left(1 + \frac{8}{3} r_F\right),$$

$$\Gamma_V = \frac{g_V^2 N_s N_C^V m_Y^3}{40\pi\Lambda^2} \beta_V f(r_V),$$

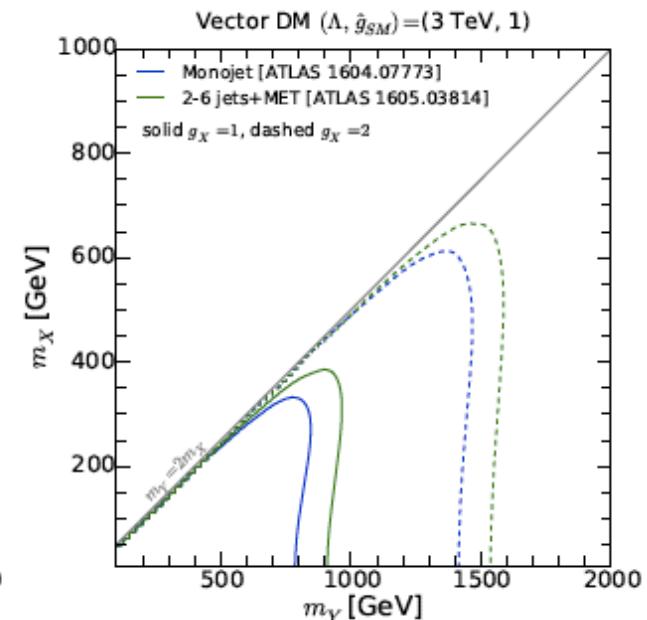
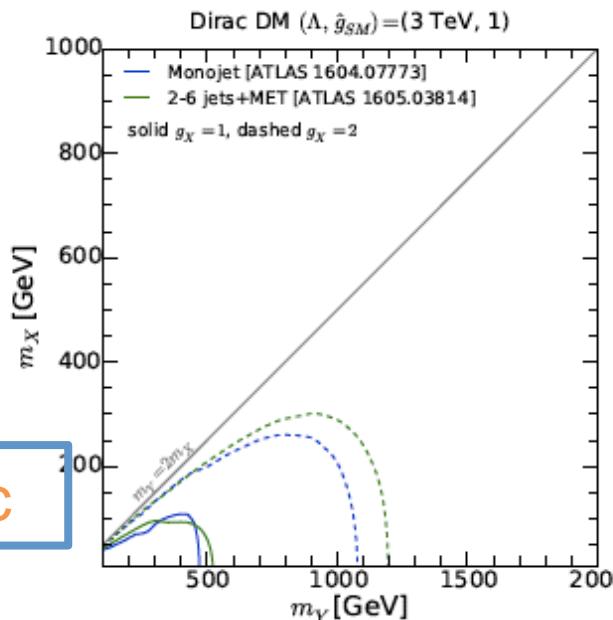
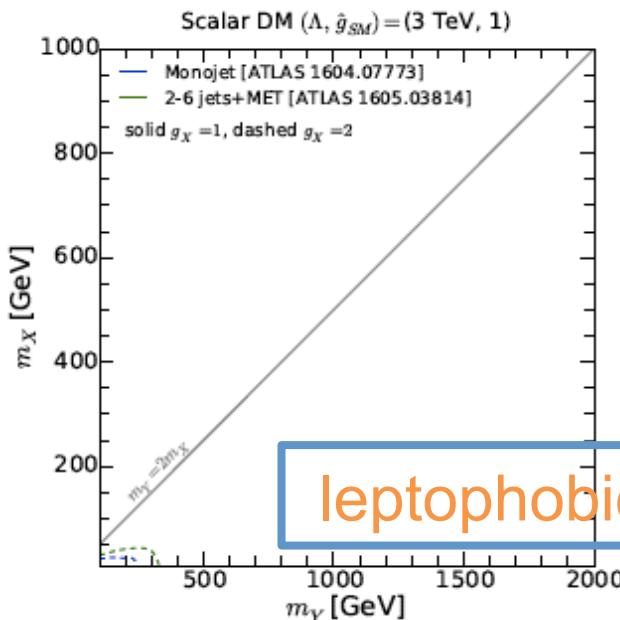
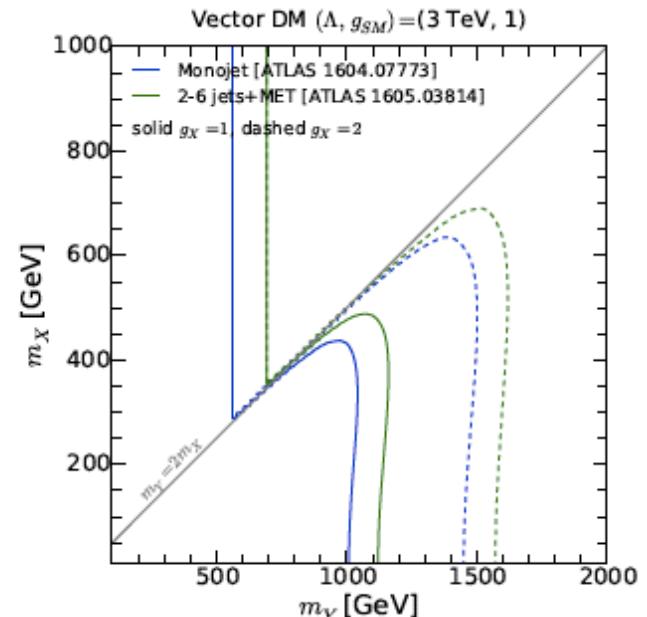
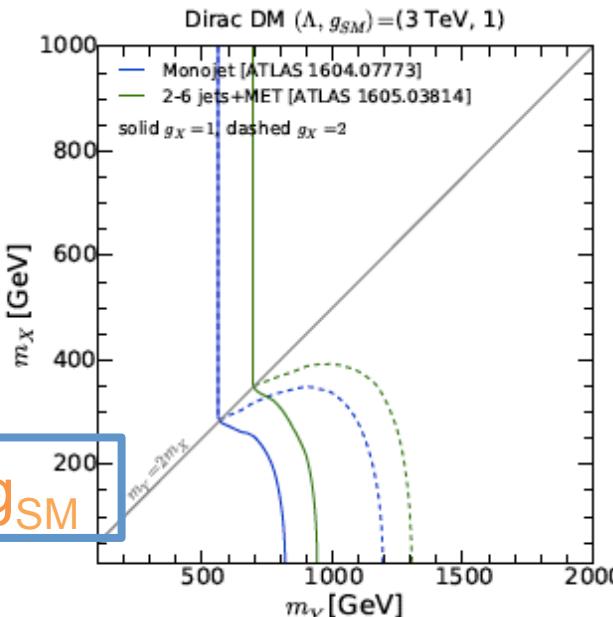
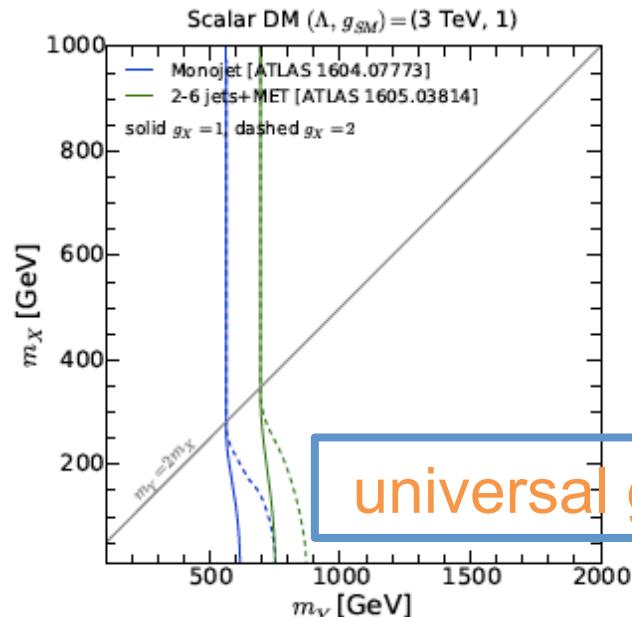
where  $\beta_i = \sqrt{1 - 4r_i}$  with  $r_i = m_i^2/m_Y^2$



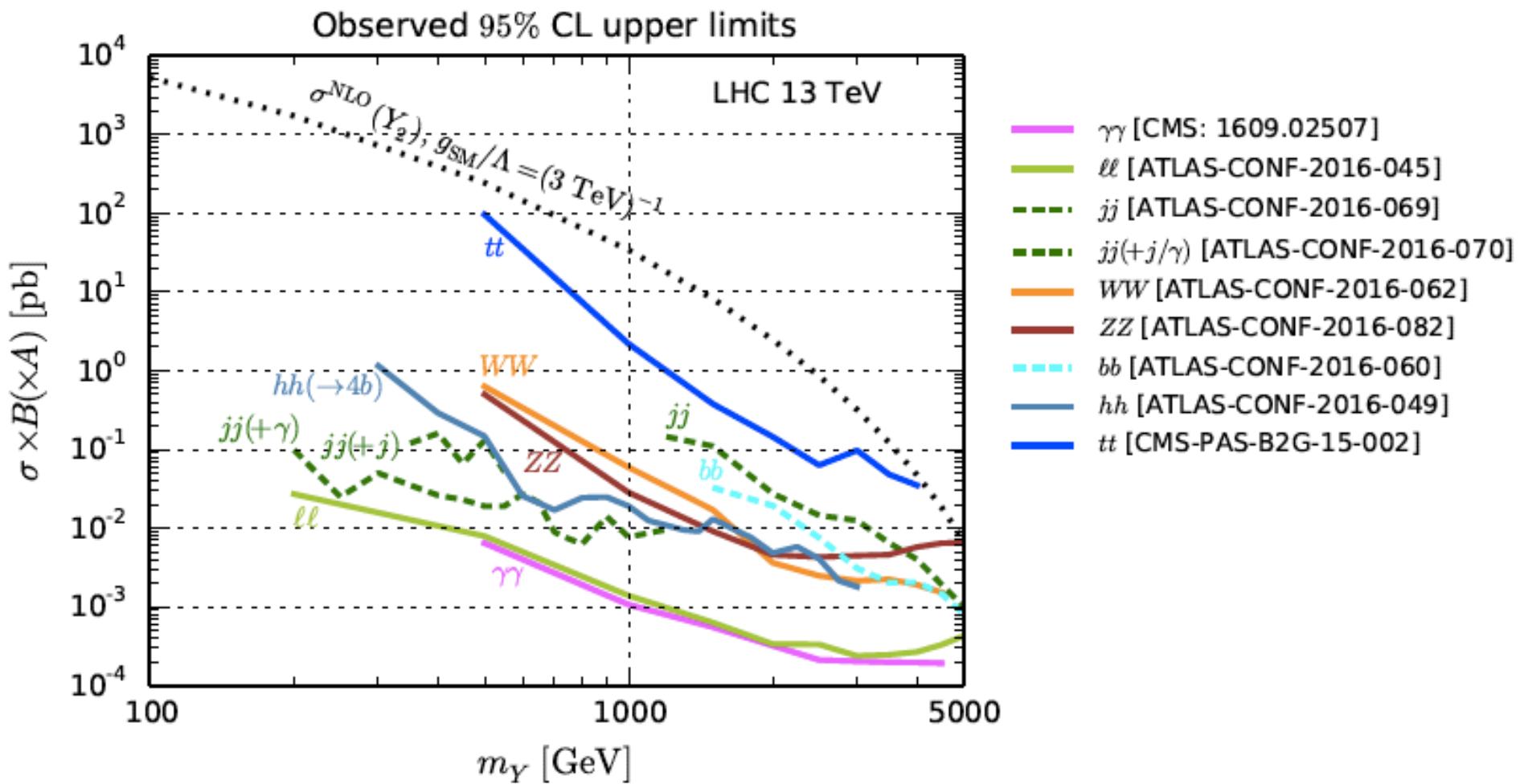
# missing ET: Scalar, Dirac, Vector DM



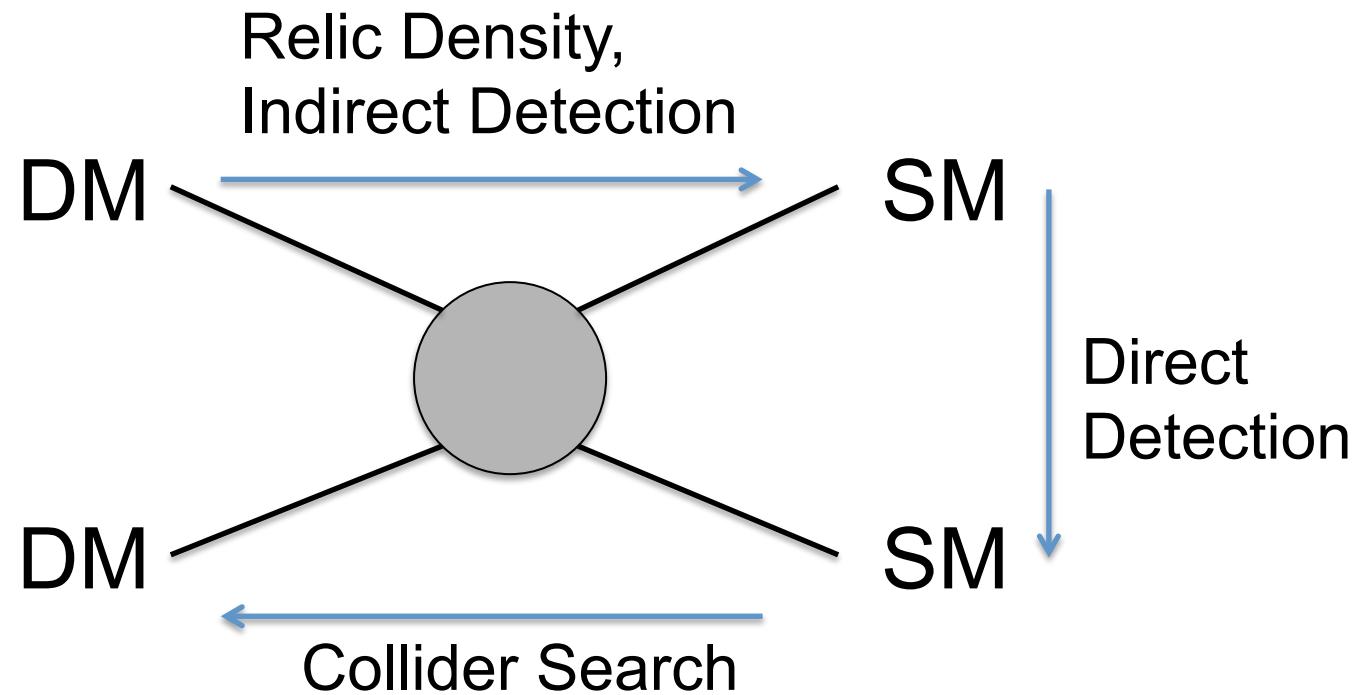
# missing ET: Scalar, Dirac, Vector DM



# Observed 95% CL upper limits



# Dark Matter: LHC search



In this talk: LHC search

# RS model 5-dimensional model

## 1. Metric

$$ds^2 = e^{-2k|y|} \eta_{\mu\nu} dx^\mu dx^\nu - dy^2, \text{ where}$$

$k$  : Ads curvature  $\sim M_{pl}$ ,  $kr_c \sim 12$

## 2. Interaction of Kaluza-Klein(KK) gravitons

$$-\frac{1}{\Lambda_\pi} T^{\alpha\beta}(x) \sum_{n=1}^{\infty} h_{\alpha\beta}^{(n)}(x), \text{ where } \Lambda_\pi = e^{-kr_c\pi} \bar{M}_{pl} \sim \text{TeV}$$

$$\bar{M}_{pl} = \frac{M_{pl}}{\sqrt{8\pi}},$$

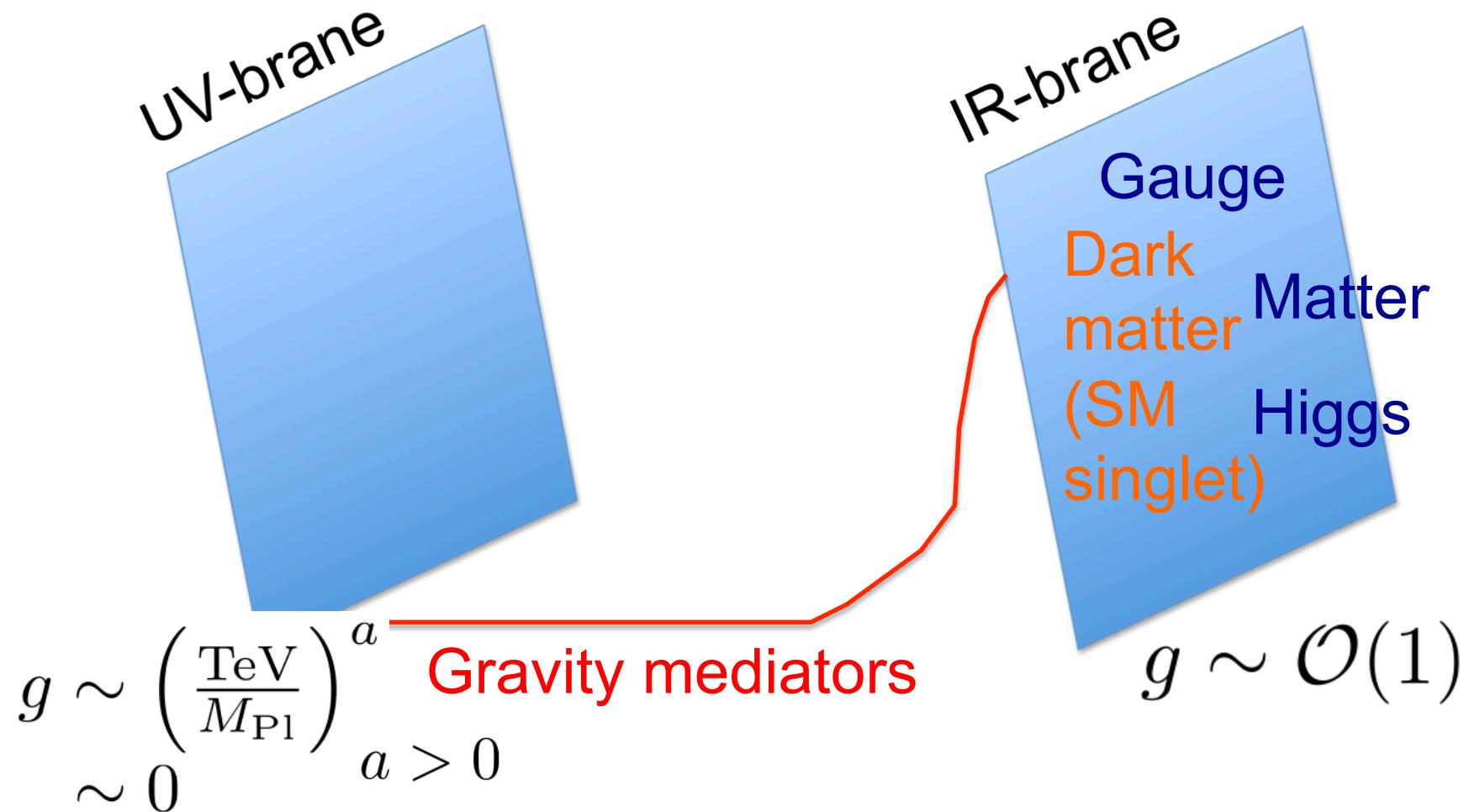
All the massive KK gravitons are suppressed by  $\Lambda_\pi \sim \text{TeV}$

## 3. Model parameters

$$(m_G, \frac{k}{\bar{M}_{pl}}), \text{ where } m_G = kx_1 e^{-kr_c\pi} (x_1 \cong 3.83)$$

$m_G$  : mass of the 1<sup>st</sup> KK graviton excitation state

# Gravity-mediated Dark Matter



SM boson  $\Leftrightarrow$  KK graviton  $\Leftrightarrow$  Dark matter

# RS model & gauge hierarchy problem

$$m_H = \tilde{m} \cdot e^{-\pi \cdot k \cdot R}$$

$$m_G = x_1 \cdot k \cdot e^{-\pi \cdot k \cdot R}$$

Reference:

Arun, *etal*, arXiv:1410.5591

$$\frac{m_G}{m_H} = x_1 \cdot \frac{k}{\tilde{m}}$$

126 GeV      3.83      planck scale:  $\bar{M}_{pl}$

$$0.01 \leq \frac{k}{\bar{M}_{pl}} \leq 1$$

$$m_G \leq 126 \text{GeV} \cdot 3.83 \cdot 1 = 480 \text{ GeV}$$

$$(m_G, \frac{k}{\bar{M}_{pl}}) = (1.25 \text{ TeV}, 0.01) \rightarrow m_H \Rightarrow 33 \text{ TeV}$$