

Cosmology: Planck 2015 results and what's next

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- ♦ Sky maps at various
 Ofrequencies
- ♦ The CMB anisotropies in 2015
- ♦ What is new since the 2013 release?
- ♦ CMB polarisation
- ❖ CMB polarisation with Planck 2015
- Polarisation power spectra
- First measurement of polarised foregrounds
- ❖ Planck 2015 bounds on Early Universe models
- ❖ Galactic clusters as watermarks
- ❖ Gravitation lensing
- Some other results
- ❖ After Planck?
- ❖ Cosmology after the CMB?

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The Planck satellite was:

Microwave telescope (ESA)

▶ PLM (Payload Module)

- Telescope (1.5m)
- Focal plane instruments
- Cryo-structure (cooler, baffles, Vgrooves)
- SVM (Service Module)
 - Avionics (Control (data, attitude), power management, Telecom + Instruments electronics
- Satellite:
 - 2000kg, 1600W,
 4.2m x 4.2m
 - Lifetime 21 months
 - Satellite in rotation 1 turn/min



- Two instruments on board: LFI / HFI
 - ◆ 9 frequencies 30, 44, 70, 100, 143, 217, 353, 545, 857 GHz
 - ◆ Decommissioned since October 2013 and three data releases: 2013, 2015 and final in 2016
 - lacktriangle Objectives were measurements of T down to astrophysical limitation (polarization was not originally planned!)

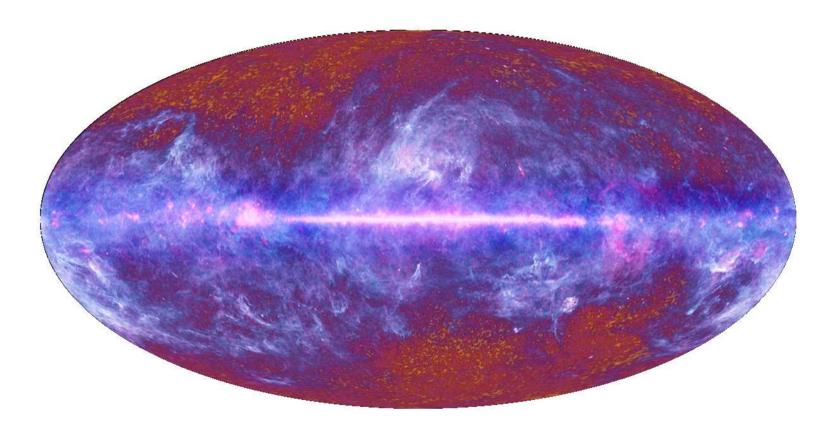


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First light (2010)

• Resolution: 50 000 000 pixels



The CMB lies behing various astrophysical foregrounds



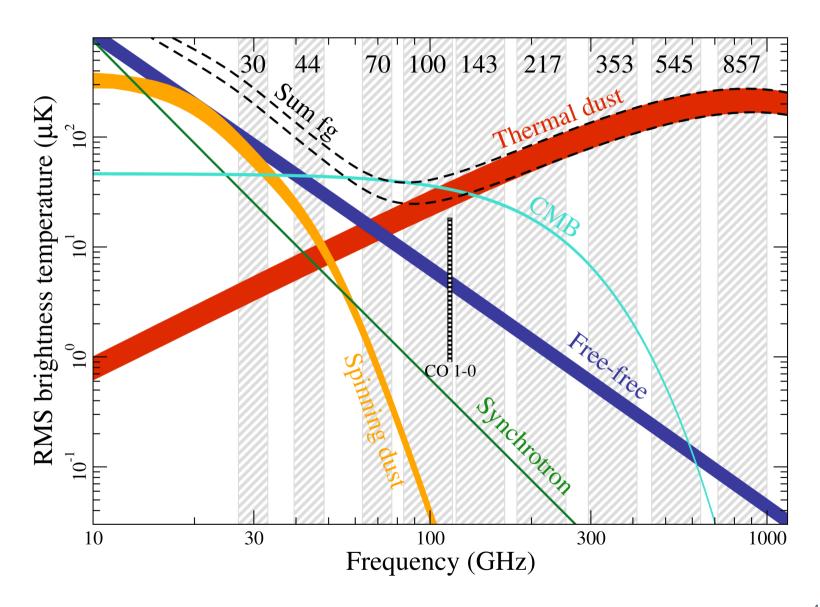
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Sky maps at various frequencies

 \bullet 9 frequencies = combination of foregrounds + CMB

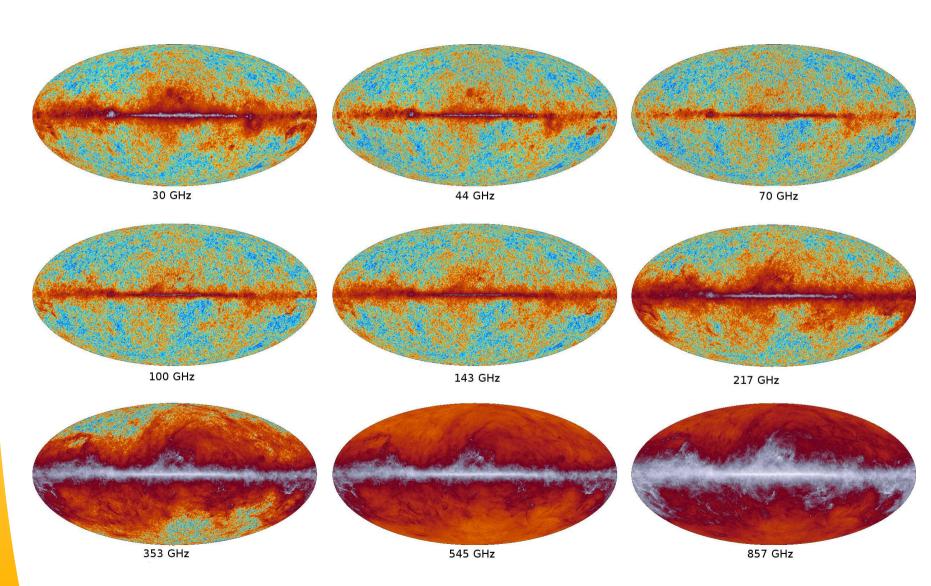




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Sky maps at various frequencies

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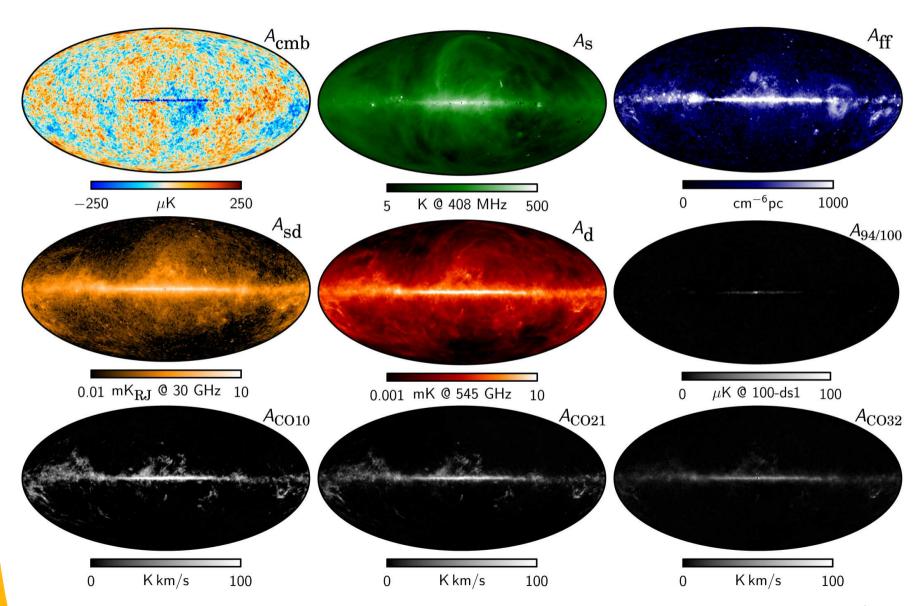


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Sky maps at various frequencies

• After diffuse component separations





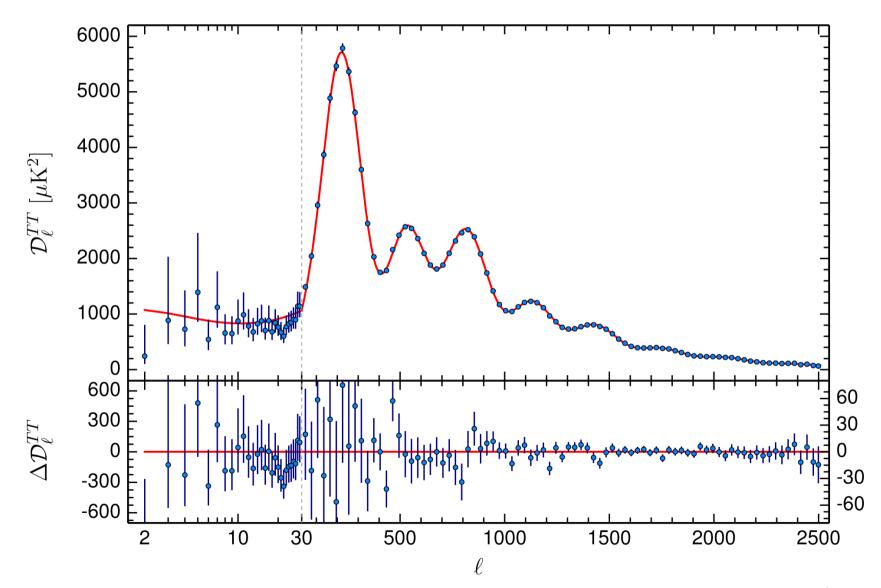
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The CMB anisotropies in 2015

Angular power spectrum





- Sky maps at various frequencies
- ◆ The CMB anisotropies in 2015

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What is new since the 2013 release?

- ullet 4 years of measurements included: twice the amount of data on T
 - Actually not used in that way: new data analysis consists of cross-correlating the first 2 years with the last 2 years to kill systematics
- Inclusion of external measurements: WMAP9 @ $23\,\mathrm{GHz}$ - $94\,\mathrm{GHz}$ + Radiotelescope survey @ $0.408\,\mathrm{GHz}$
- ullet Polarization measurements over 7 frequencies: two new observables E (and B) in addition to T
- A lot of improvements in the understanding of noise, calibration, satellite pointing, ADC, and astrophysics
- What's new for Cosmology?
 - \bullet The picture is the same: $\Lambda {\rm CDM}$ works very well with 6 parameters (over almost $2\,000\,000$ of modes): error bars on all cosmological parameters have been reduced by 1 to 1.5σ
 - ♦ More constraints on early universe acceptable models (in progress)
 - ♦ A lot of new results for Astrophysics and a lot of new data still unexploited for Cosmology and other fields



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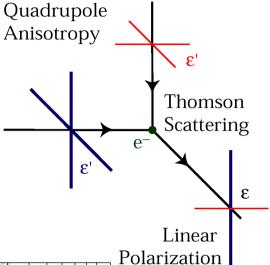
CMB polarisation

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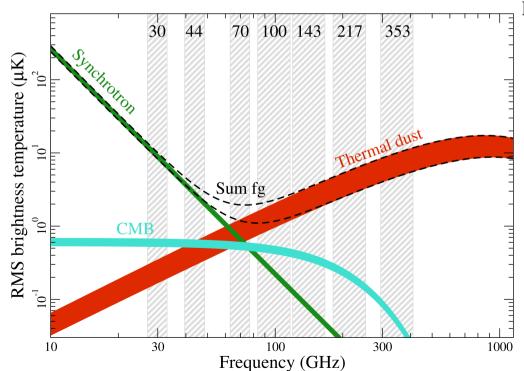
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CMB polarisation

- Physical origin: photons-electrons interactions
 - ◆ AND anisotropies
 - lacktriangle Two new observables E and B (or Stockes parameters Q and U)



ullet Main issue: Signal/Foregounds < 1





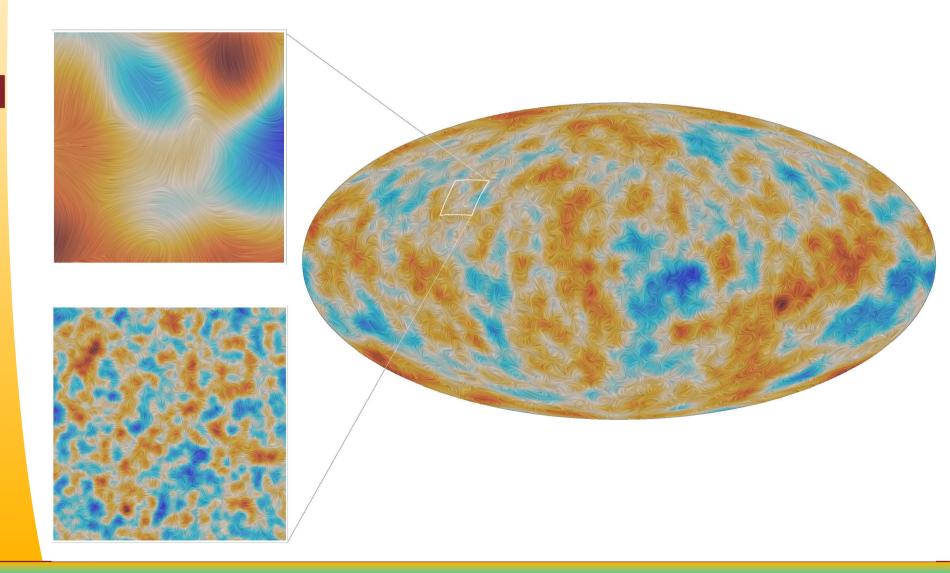
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CMB polarisation with Planck 2015

After component separations: intensity (color) and direction (integral lines)





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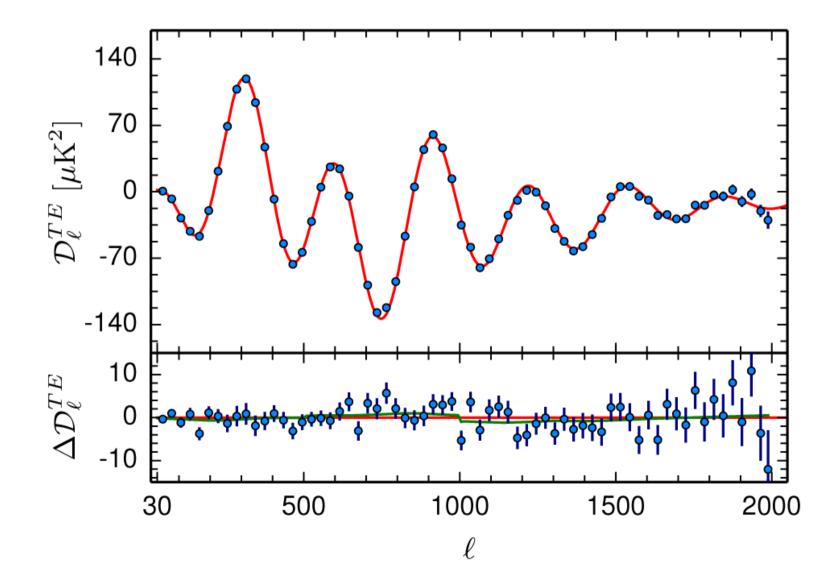
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Polarisation power spectra

ullet Correlation intensity-direction TE





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Polarisation power spectra

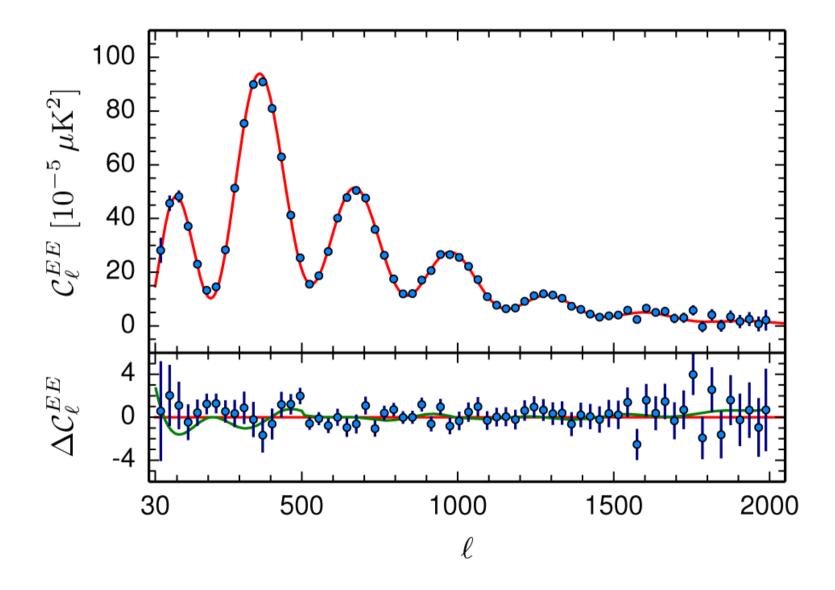
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Polarisation power spectra

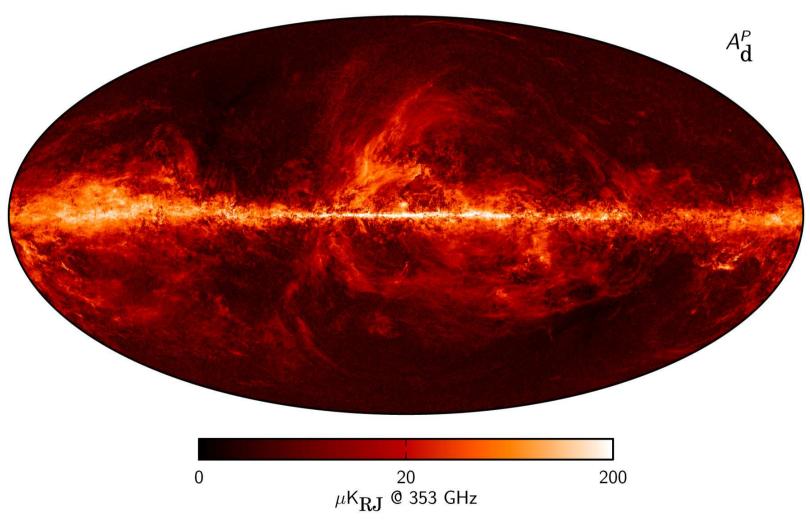
Autocorrelation EE





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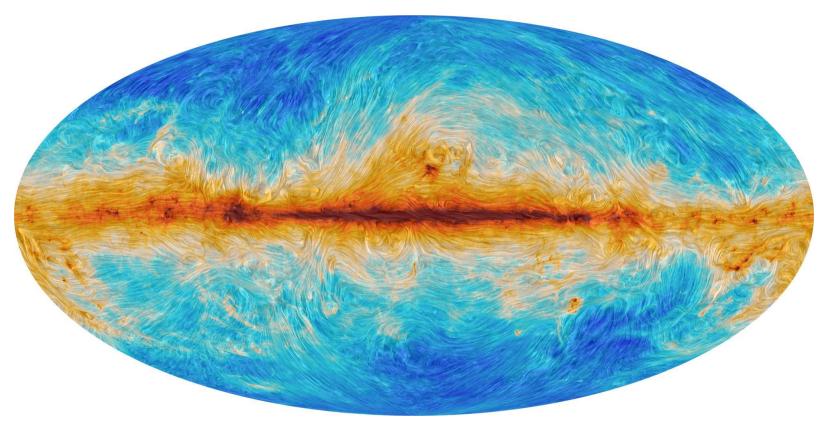
- Non-spherical grains aligned along galactic magnetic field
 - lacktriangle Polarized intensity I_p





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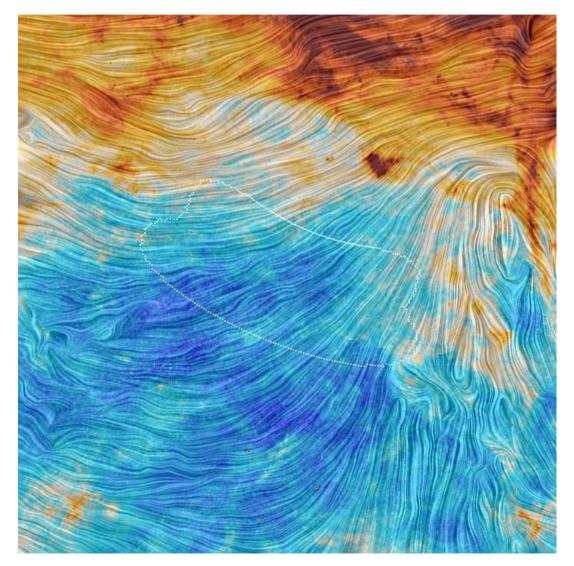
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 - ◆ Direction ⇔ magnetic field direction





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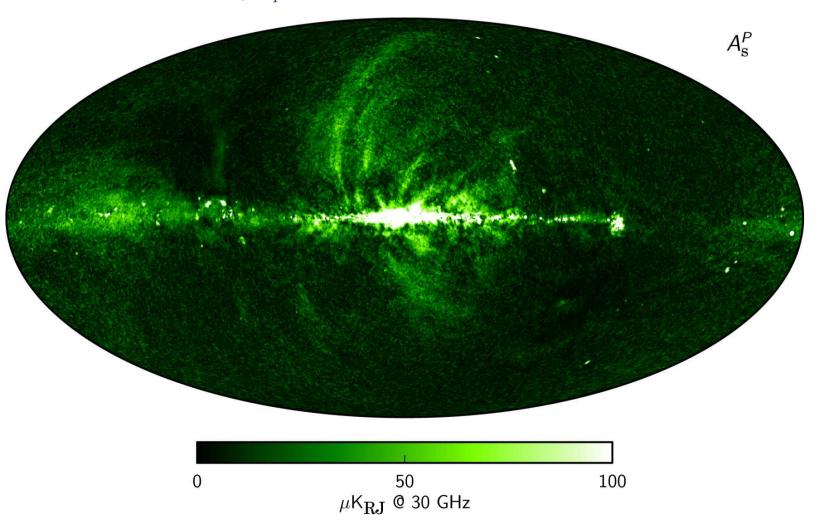
- Non-spherical grains aligned along galactic magnetic field
 - ◆ zoom within the BICEP2 region





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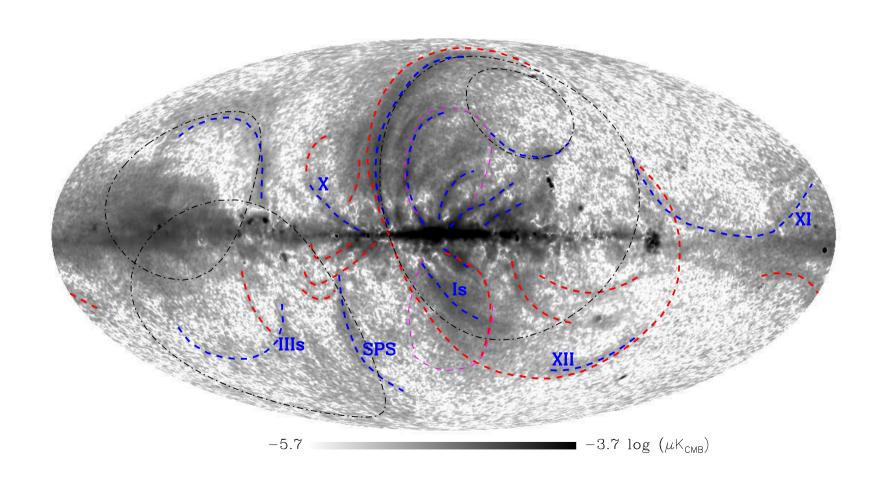
- Polarised synchrotron emission from electrons spiraling aroung the galactic magnetic field
 - lacktriangle Polarized intensity I_p





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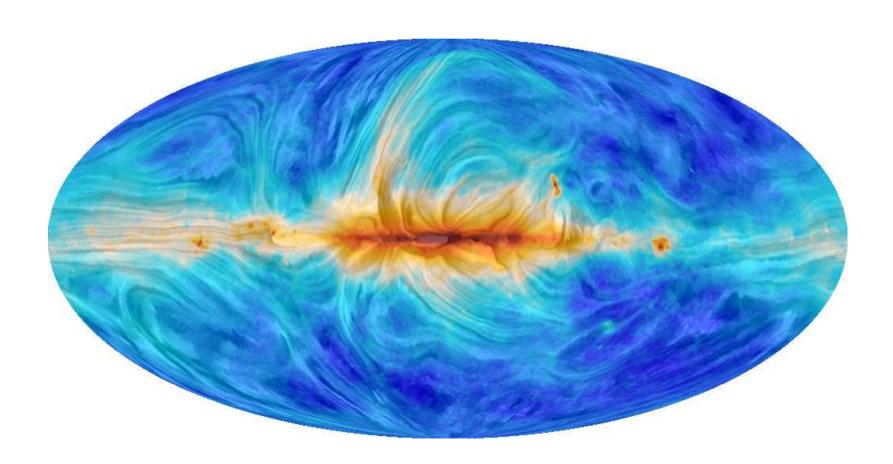
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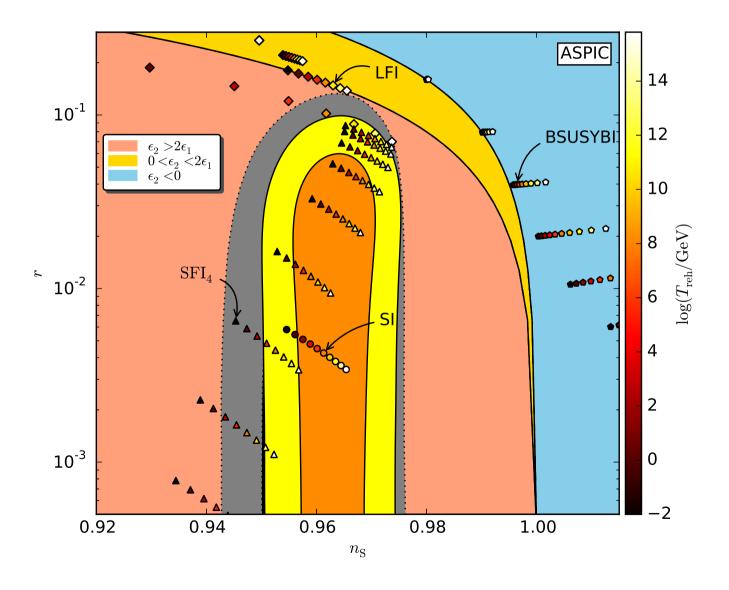




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Planck 2015 bounds on Early Universe models

Comparison between Planck 2013 and Planck 2015 + BICEP2





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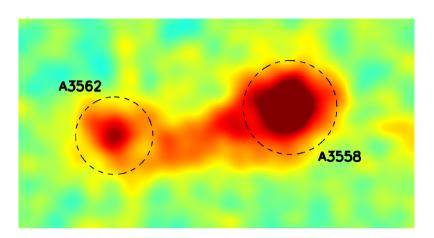
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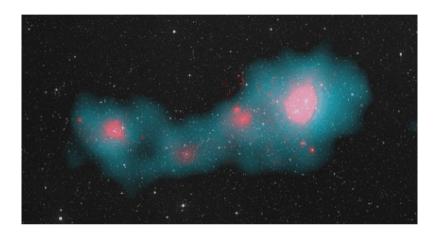
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Sunyaev-Zel'dovich effect

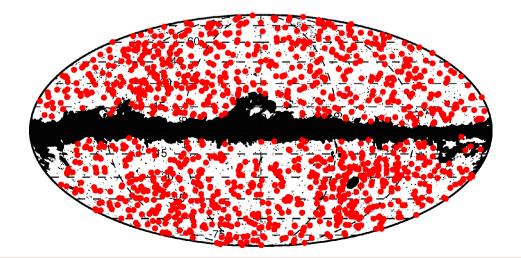
CMB



Optical, X-ray and SZ



• A thousand new clusters discovered: \Longrightarrow new measurement of Ω_{mat} , σ_{8}





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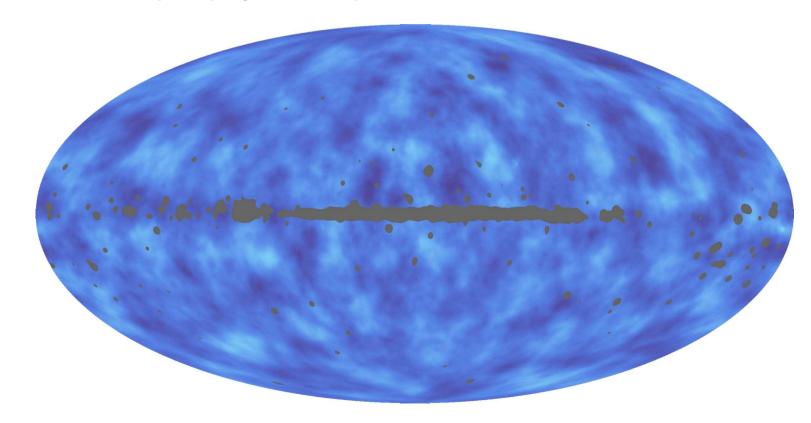
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Gravitational lensing

CMB is deflected by structures and dark matter

$$\tilde{T}(\boldsymbol{n}) = T[\boldsymbol{n} + \nabla \phi(\boldsymbol{n})] \simeq T(\boldsymbol{n}) + \nabla^i \phi(\boldsymbol{n}) \nabla_i T(\boldsymbol{n})$$

ullet Deflexion map \simeq projected map of dark matter distribution





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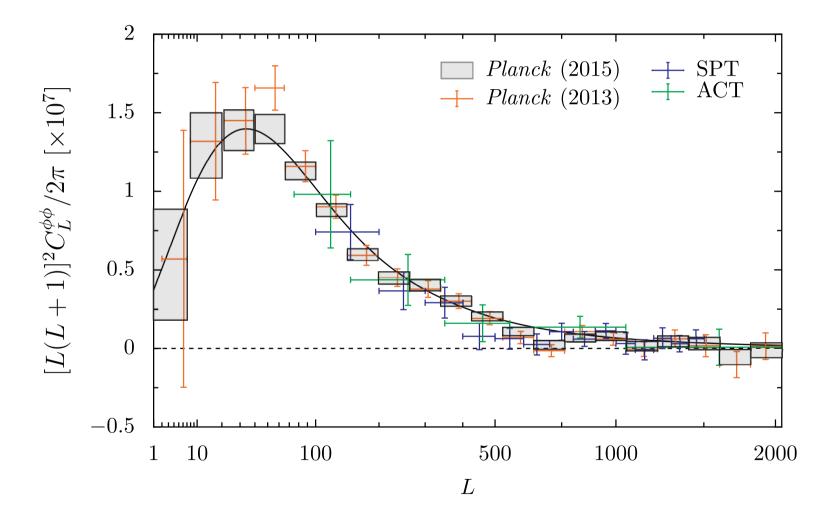
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Lensing power spectrum

 \bullet Compared to the DM + GR predictions





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Some other results

• Neutrino masses (TT+lowP+lensing+ H_0 +BAO+JLA)

$$\sum_{\nu} m_{\nu} < 0.23 \,\text{eV} \quad (95\%)$$

• Extra relativistic degrees of freedom at $T \ll \mathrm{MeV}$

$$\rho \equiv \frac{7}{8} \left(\frac{4}{11}\right)^{4/3} N_{\text{eff}} \rho_{\gamma} \qquad N_{\text{eff}} = 3.15 \pm 0.23 \qquad (\text{TT+lowP+BAO})$$



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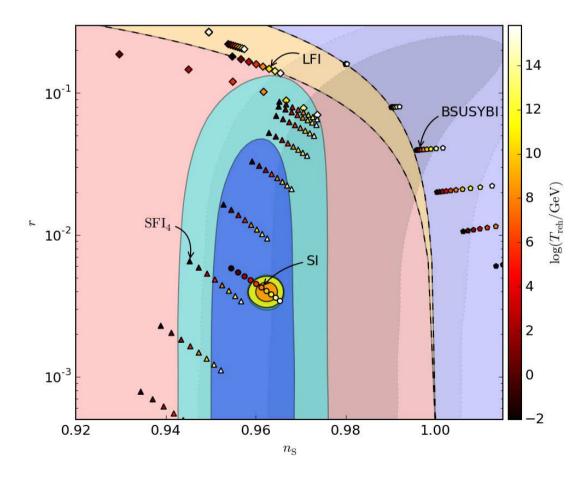
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After Planck?

- CMB: intensity measurements are over, but not polarization
- New detectors (as BICEP2) currently deployed in telescopes + balloons:
 CLASS, SPIDER, ABS, KECK, BICEP3 (2560 detectors)...
- Next step is spatial missions: Core+, LiteBird (Japan)



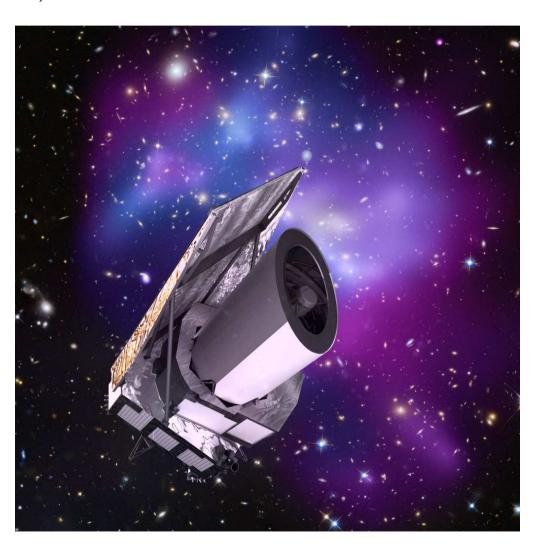


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Cosmology after the CMB?

 Galaxy surveys: SDSS IV (2014-2020), LSST (2019), EUCLID satellite (first light 2020)



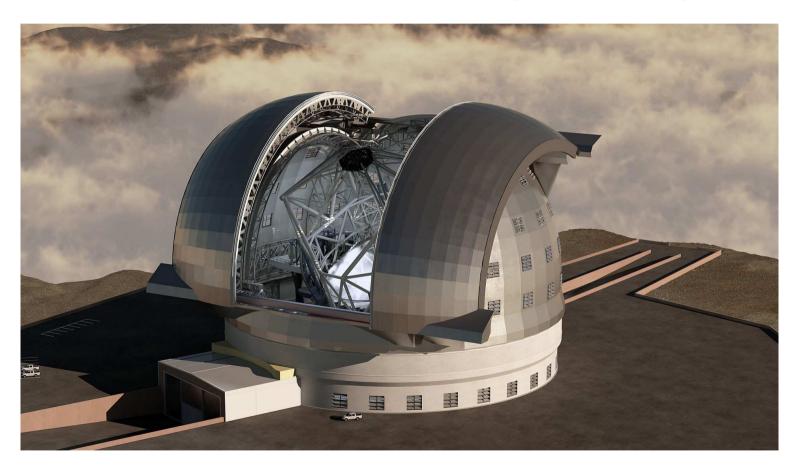


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Cosmology after the CMB?

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- 3 Extremely Large Telescopes in construction (first light 2024)





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- 21cm interferometric arrays: SKA (construction 2018)





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- 3 Extremely Large Telescopes in construction (first light 2024)
- 21cm interferometric arrays: SKA (construction 2018)
- Direct detection for GW with VIRGO/LIGO (2016-2017?)
 - ♦ eLISA has been accepted (2034!) while LPF will be launched soon 27/11

