

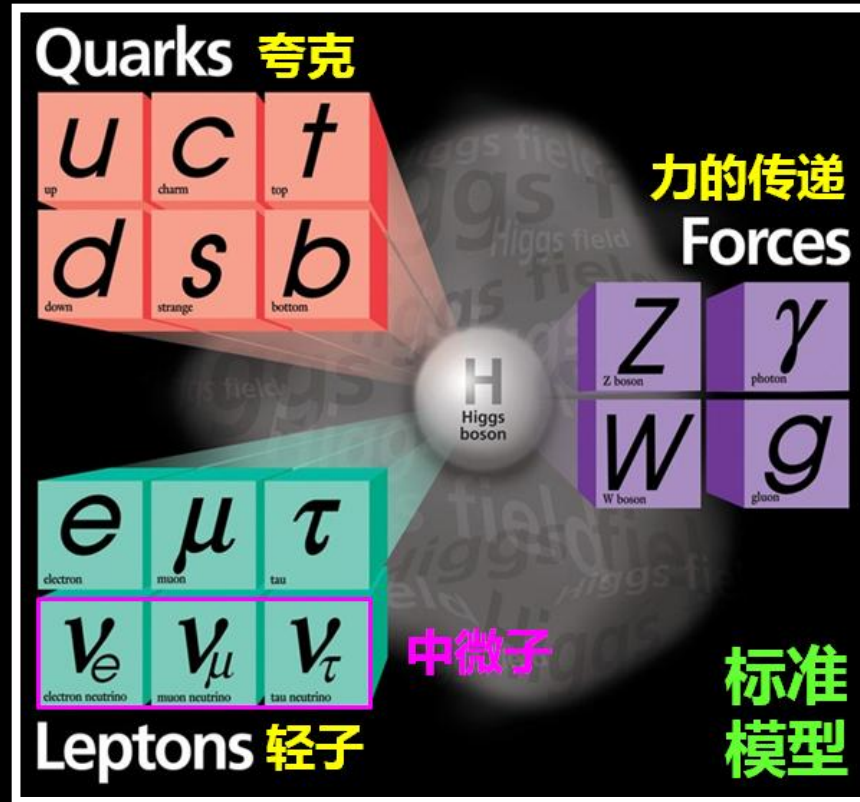
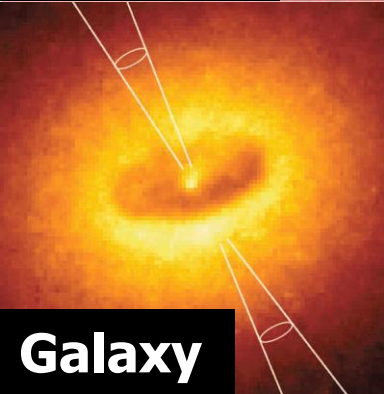
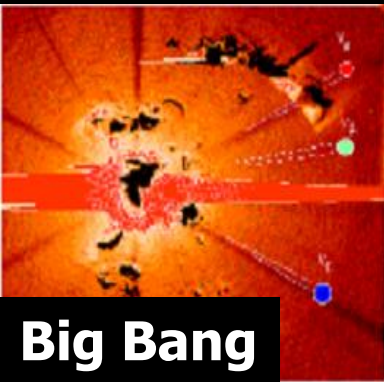
Physics of Massive Neutrinos

Zhi-zhong Xing
(IHEP, Beijing)

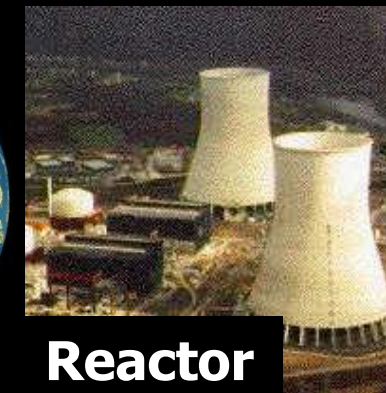
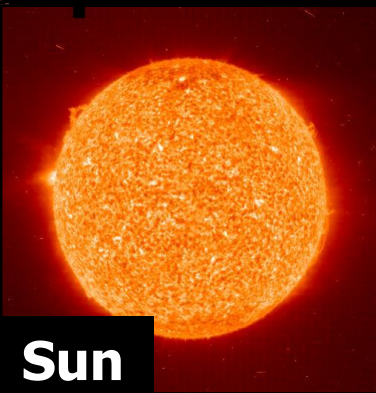
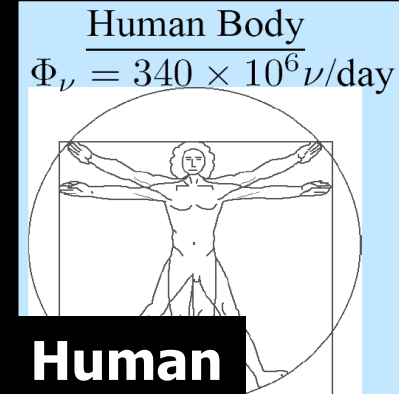
- A:** Neutrino history and lepton family
- B:** Neutrino masses and flavor mixing
- C:** Description of neutrino oscillations
- D:** Implication of neutrino oscillations
- E:** θ_{13} hint and sterile neutrino puzzle
- F:** On TeV neutrino physics at the LHC
- G:** Extreme corners in the neutrino sky
- H:** Cosmic baryon number asymmetry

@Lectures at CP3, UCL, Louvain-la-Neuve, Belgium, Jan 2012

Neutrinos: how elusive they are?



charge = 0
spin = $\frac{1}{2}$
mass = 0
speed = c



Supernova

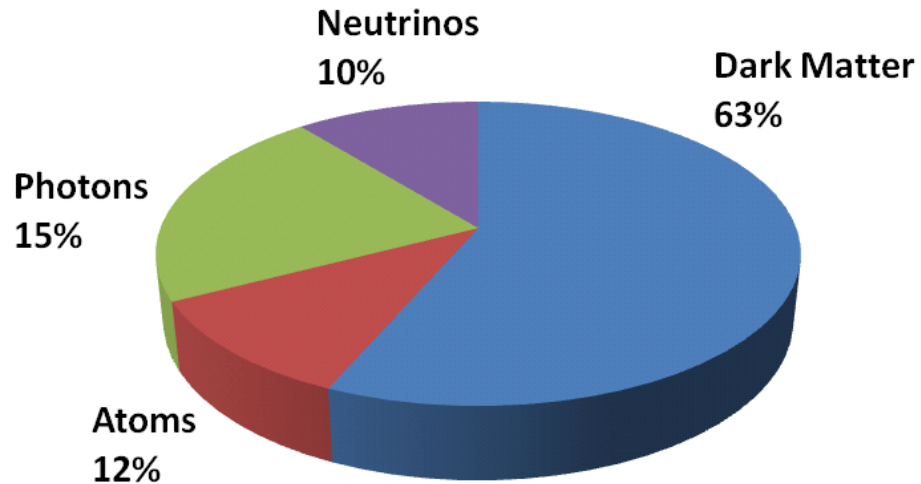
Sun

Earth

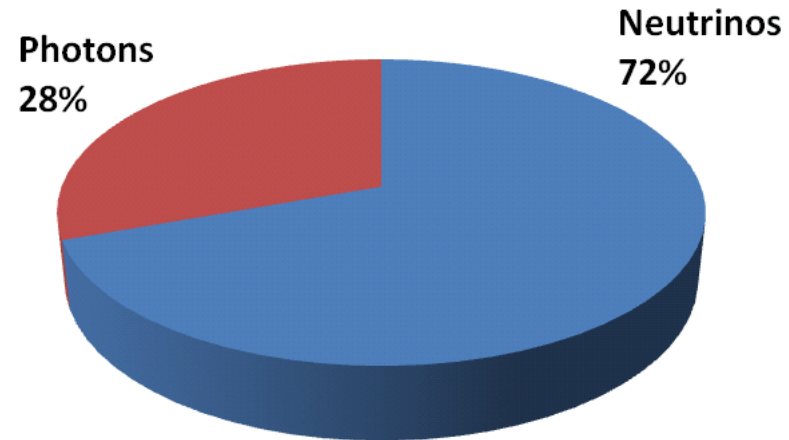
Reactor

Accelerator

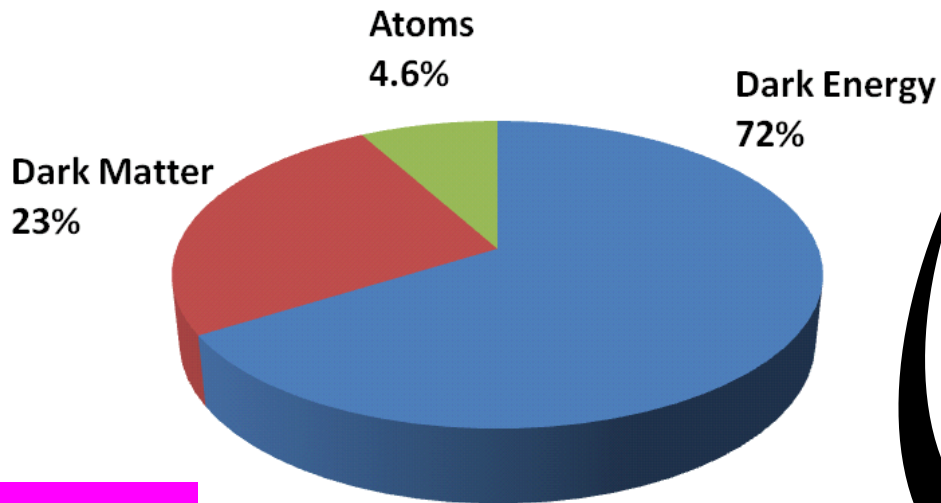
NEUTRINOS



photon decoupling
 $t = 380\,000$ years



neutrino decoupling
 $t = 1$ second

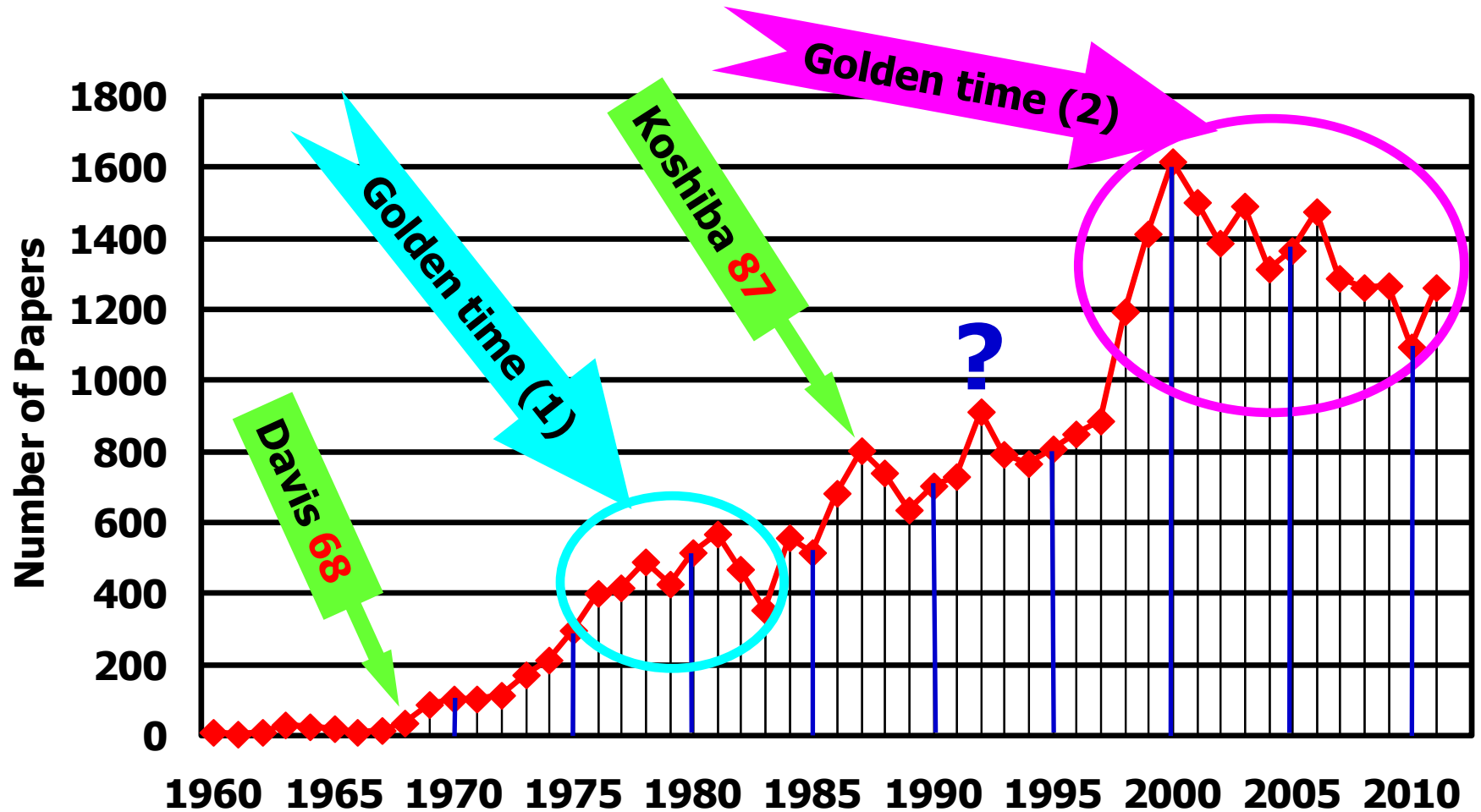


Today
 $t = 13.7$ billion years

A very important role
in the evolution of
the Universe



Annual papers on neutrinos >1000



Open Questions

the absolute mass scale?
the mass hierarchy problem?
the flavor desert problem?
how small is θ_{13} ?
leptonic CP violation?
the Majorana nature?
how many species? ...
cosmic ν background?
supernova & stellar ν 's?
UHE cosmic ν 's?
warm dark matter?
matter-antimatter asymmetry

2011 (440 pages)

ADVANCED TOPICS IN SCIENCE AND TECHNOLOGY IN CHINA

Zhizhong Xing
Shun Zhou

Neutrinos in Particle Physics, Astronomy and Cosmology



ZHEJIANG UNIVERSITY PRESS
浙江大学出版社



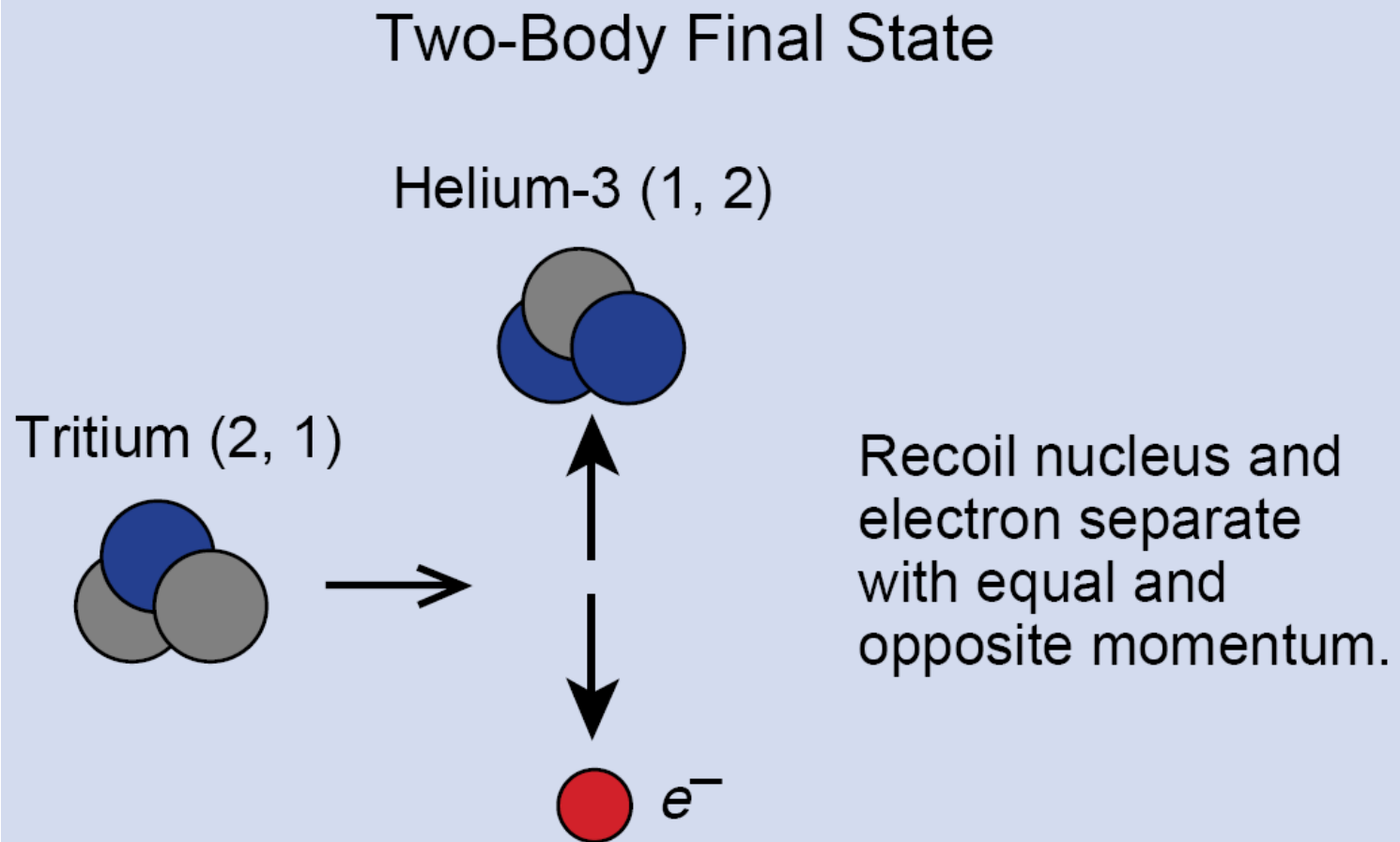
Springer

Lecture A

- ★ **Neutrino: Concept and Discovery**
- ★ **Parity, Helicity and Handedness**
- ★ **Lepton Flavors and Families**

Beta Decay in 1930

7

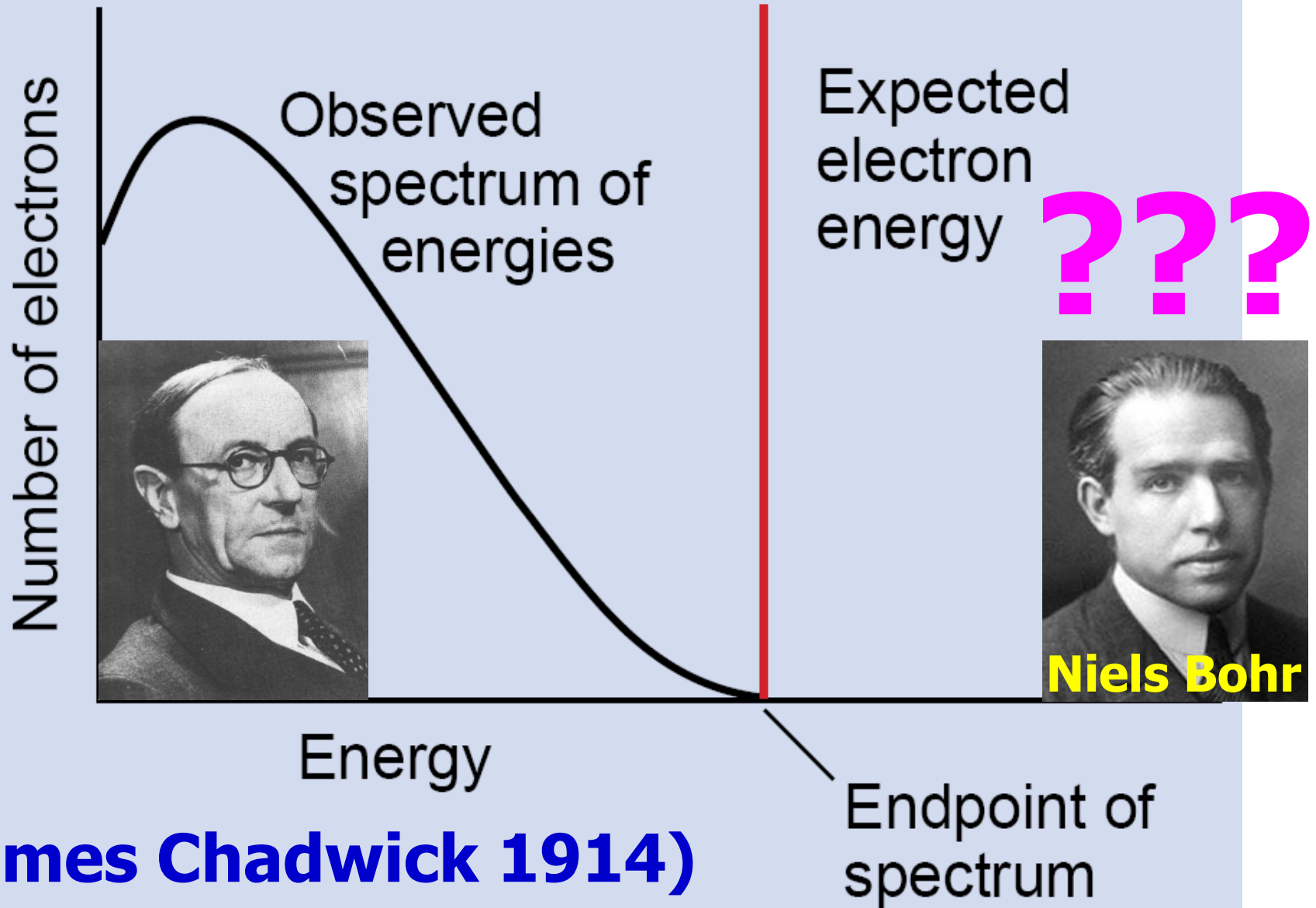


$$(N, Z) \rightarrow (N - 1, Z + 1) + e^{-},$$

where N = number of neutrons, and
 Z = number of protons.

Energy Crisis?

8



Desperate Remedy

9

Wolfgang Pauli
(1930)



The Desperate Remedy

4 December 1930
Gloriastr.
Zürich

Physical Institute of the
Federal Institute of Technology (ETH)
Zürich

Dear radioactive ladies and gentlemen,

Pauli's Letter

10

As the bearer of these lines, to whom I ask you to listen graciously, will explain more exactly, considering the 'false' statistics of N-14 and Li-6 nuclei, as well as the continuous β -spectrum, I have hit upon a desperate remedy to save the "exchange theorem"* of statistics and the energy theorem. Namely [there is] the possibility that there could exist in the nuclei electrically neutral particles that I wish to call neutrons,** which have spin $1/2$ and obey the exclusion principle, and additionally differ from light quanta in that they do not travel with the velocity of light: The mass of the neutron must be of the same order of magnitude as the electron mass and, in any case, not larger than 0.01 proton mass. The continuous β -spectrum would then become understandable by the assumption that in β decay a neutron is emitted together with the electron, in such a way that the sum of the energies of neutron and electron is constant.

But I don't feel secure enough to publish anything about this idea, so I first turn confidently to you, dear radioactives, with a question as to the situation concerning experimental proof of such a neutron, if it has something like about 10 times the penetrating capacity of a γ ray.

I admit that my remedy may appear to have a small *a priori* probability because neutrons, if they exist, would probably have long ago been seen. However, only those who wager can win, and the seriousness of the situation of the continuous β -spectrum can be made clear by the saying of my honored predecessor in office, Mr. Debye, who told me a short while ago in Brussels, "One does best not to think about that at all, like the new taxes." Thus one should earnestly discuss every way of salvation.—So, dear radioactives, put it to test and set it right.—Unfortunately, I cannot personally appear in Tübingen, since I am indispensable here on account of a ball taking place in Zürich in the night from 6 to 7 of December.—With many greetings to you, also to Mr. Back, your devoted servant,

W. Pauli

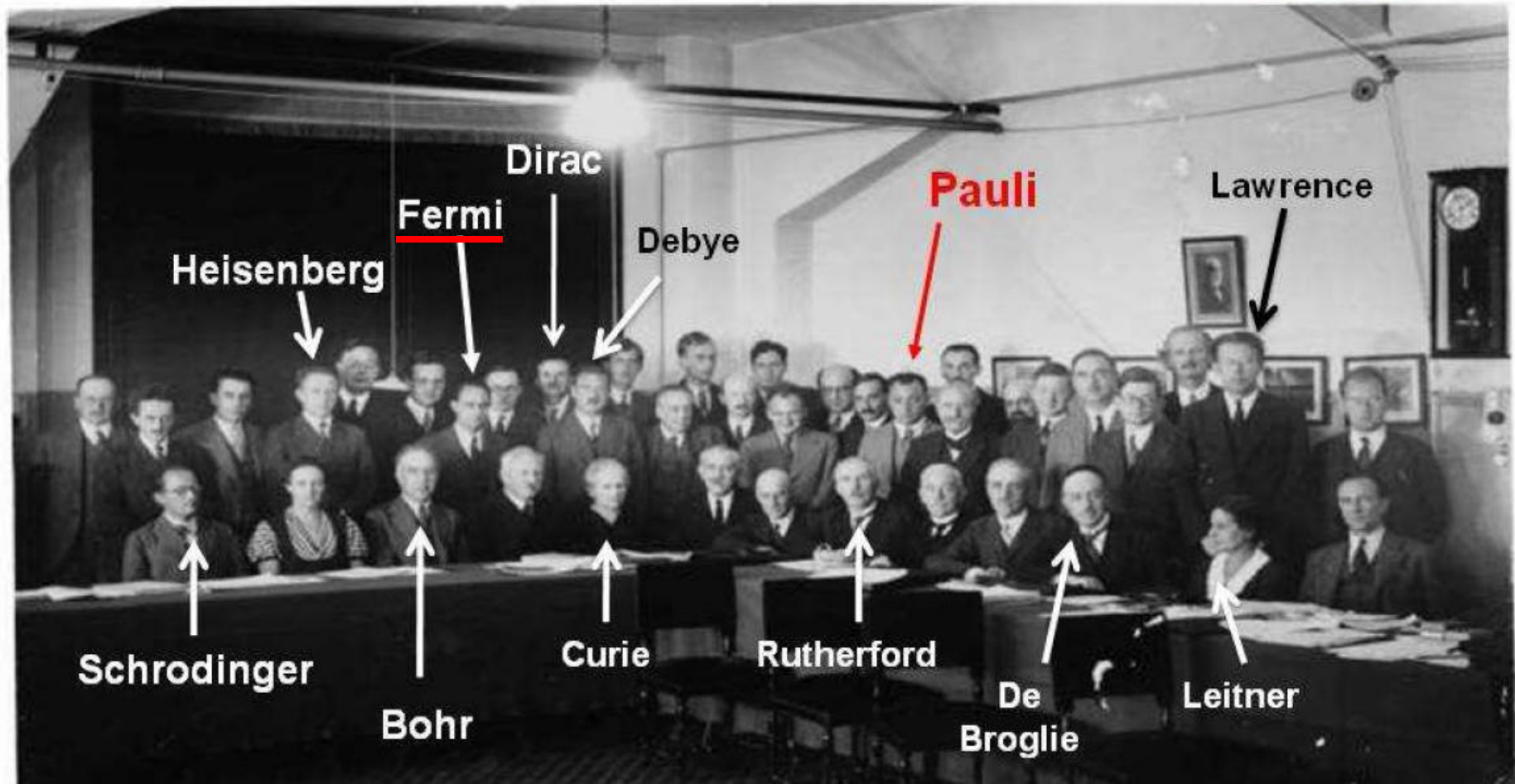
Solvay 1933

12

Pauli gave a talk on his **neutrino** proposal in this congress.

INSTITUT INTERNATIONAL DE PHYSIQUE SOLVAY
SEPTIÈME CONSEIL DE PHYSIQUE -- BRUXELLES, 22-29 OCTOBRE 1933

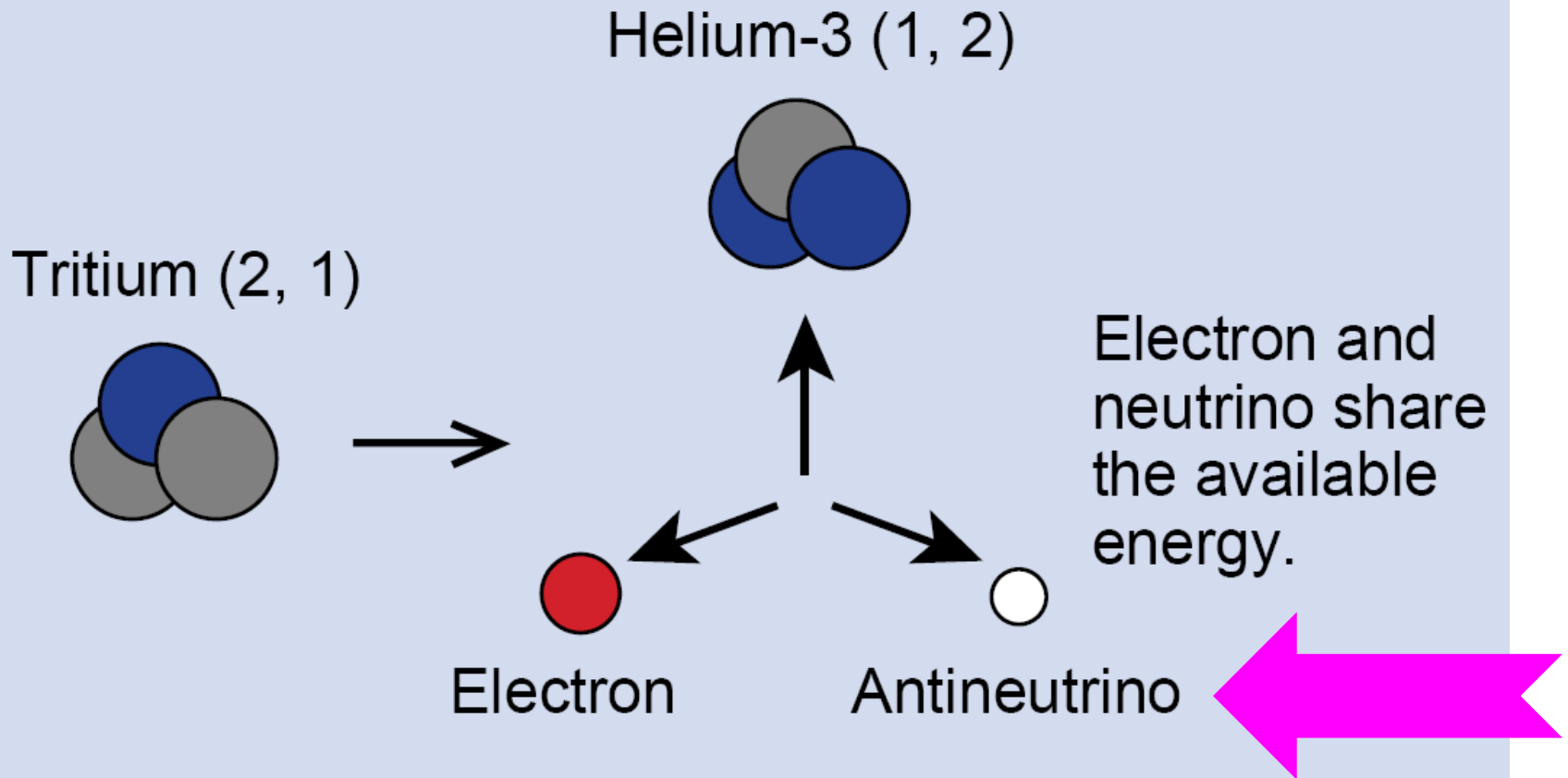
22 – 29 Octobre 1933



True Picture of β Decay

13

Three-Body Final State



$$(N, Z) \rightarrow (N - 1, Z + 1) + e^{-} + \bar{\nu} \text{ .}$$

Fermi's Paper

14

E. Fermi's publications on the Weak Interaction

REJECTED

E. Fermi, "Tentative Theory of Beta Rays"
Letter Submitted to Nature (1933)

31 Dec, 1933



Published first in this journal and latter in Z. Phys. in 1934.

Fermi's Theory

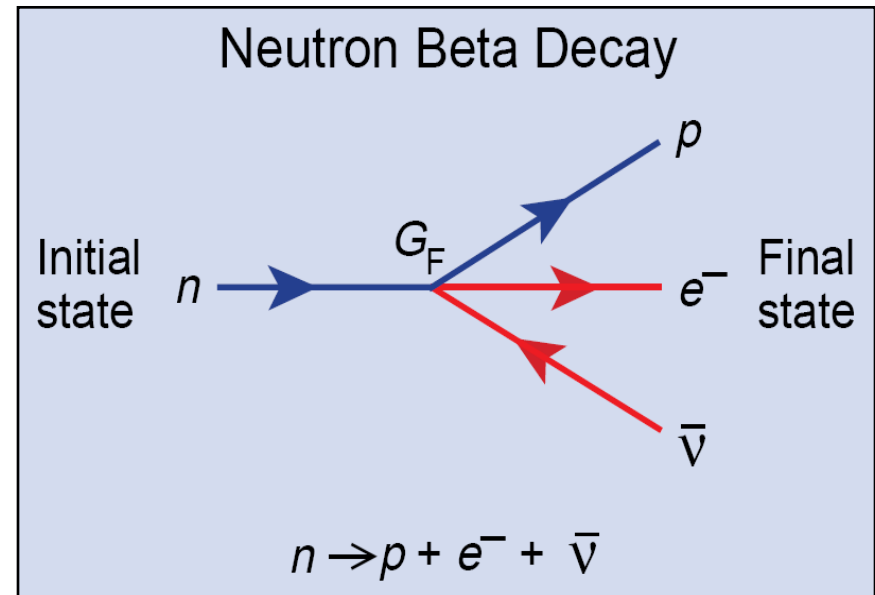
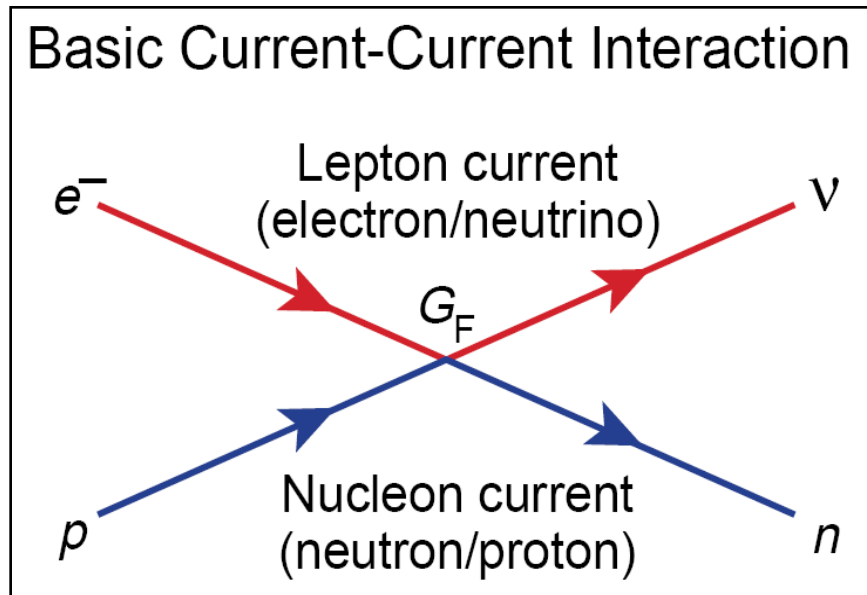
15

Enrico Fermi (1933 & 1934) assumed a new force for β decay by combining 3 brand-new concepts:

★ **Pauli's hypothesis: neutrinos**

★ **Dirac's thought: creation of particles**

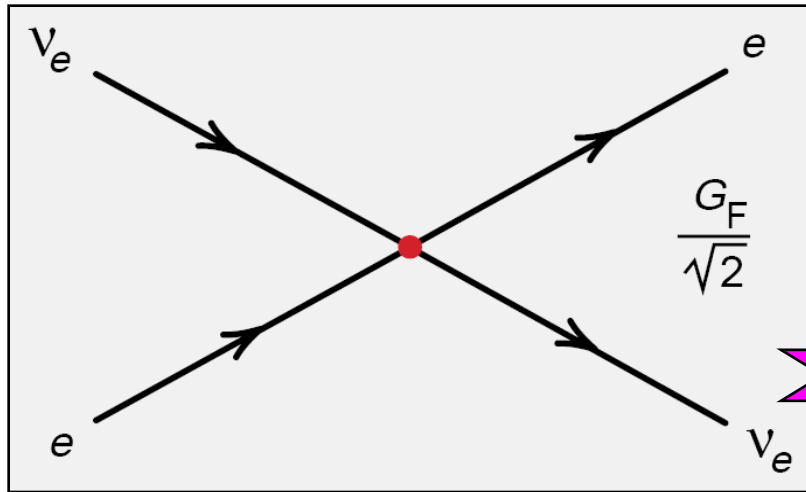
★ **Heisenberg's idea: neutron is related to proton**



Weak Interactions

16

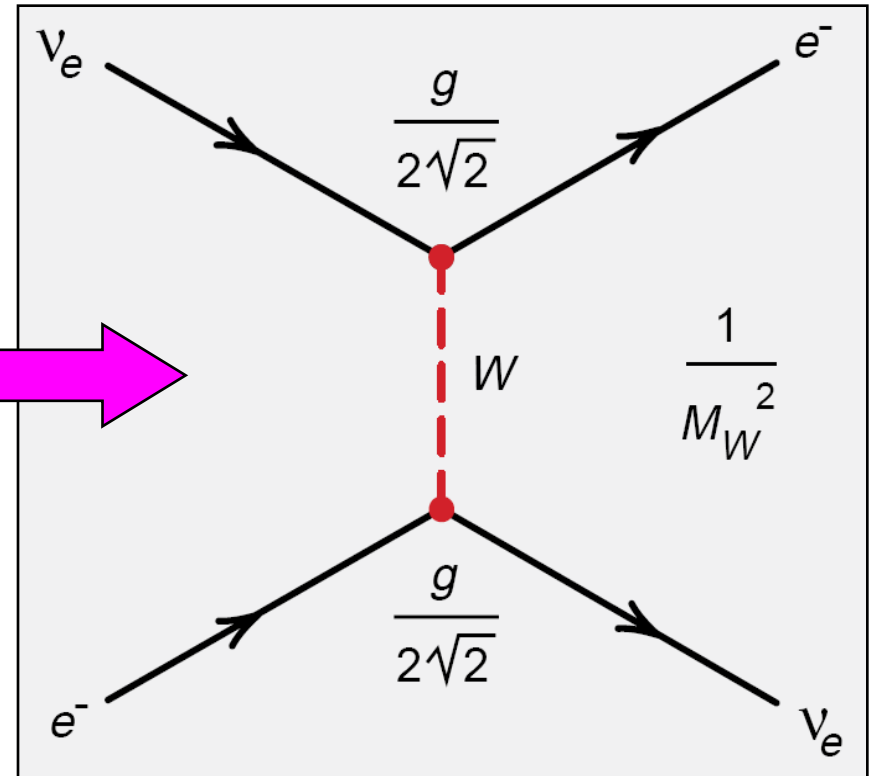
From Fermi's current-current interaction to weak charged-current gauge interactions



$$\frac{G_F}{\sqrt{2}} = \frac{g^2}{8M_W^2}$$



$$M_W = 80.4 \text{ GeV}$$

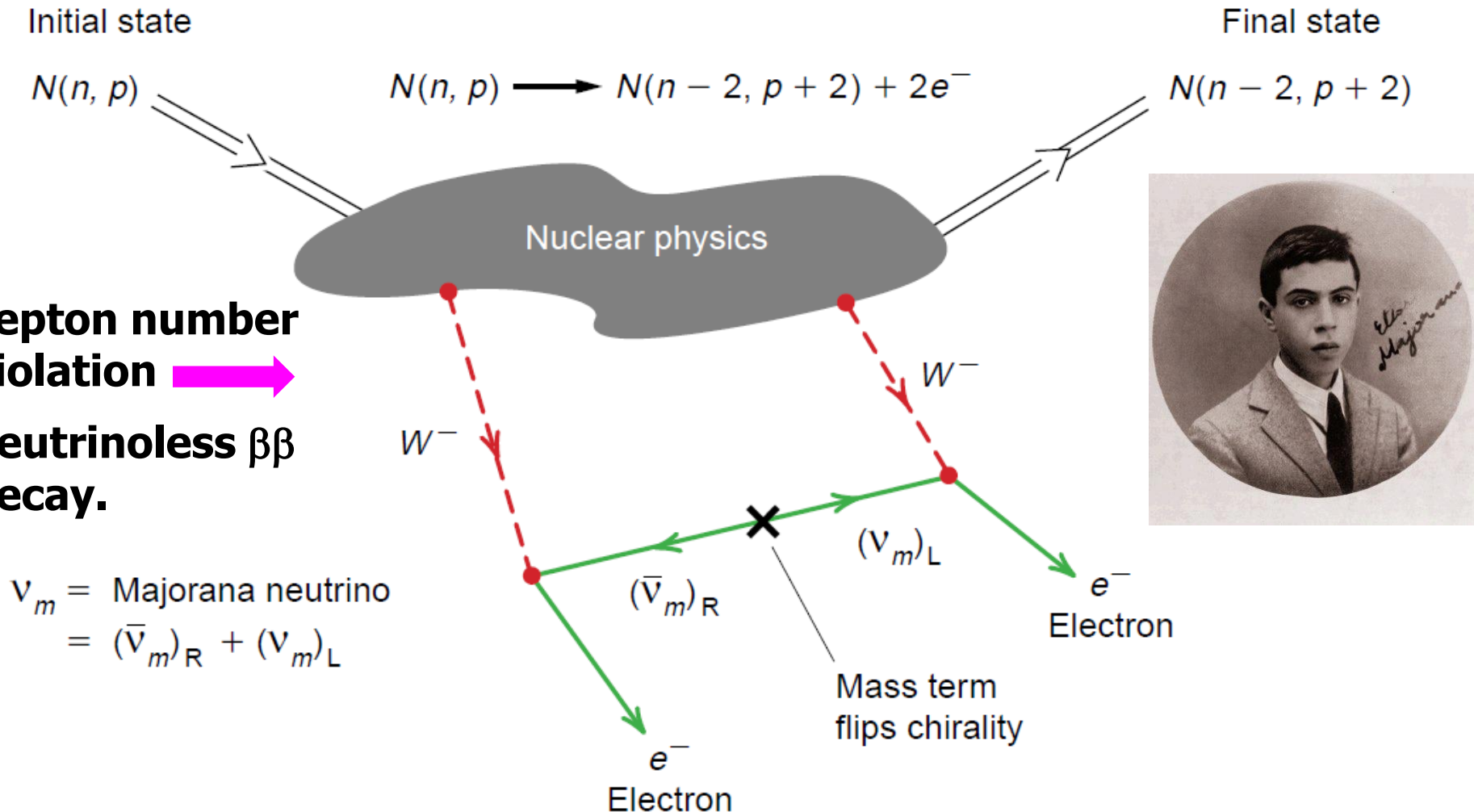


$$G_F = 1.66 \times 10^{-5} \text{ GeV}^{-2}$$

Majorana Neutrinos

17

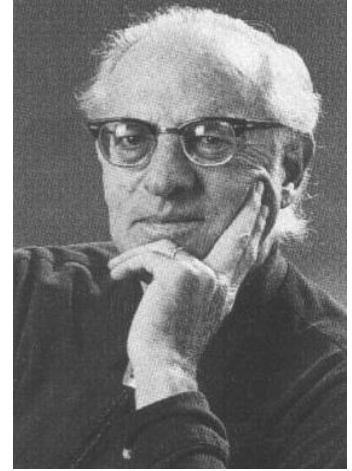
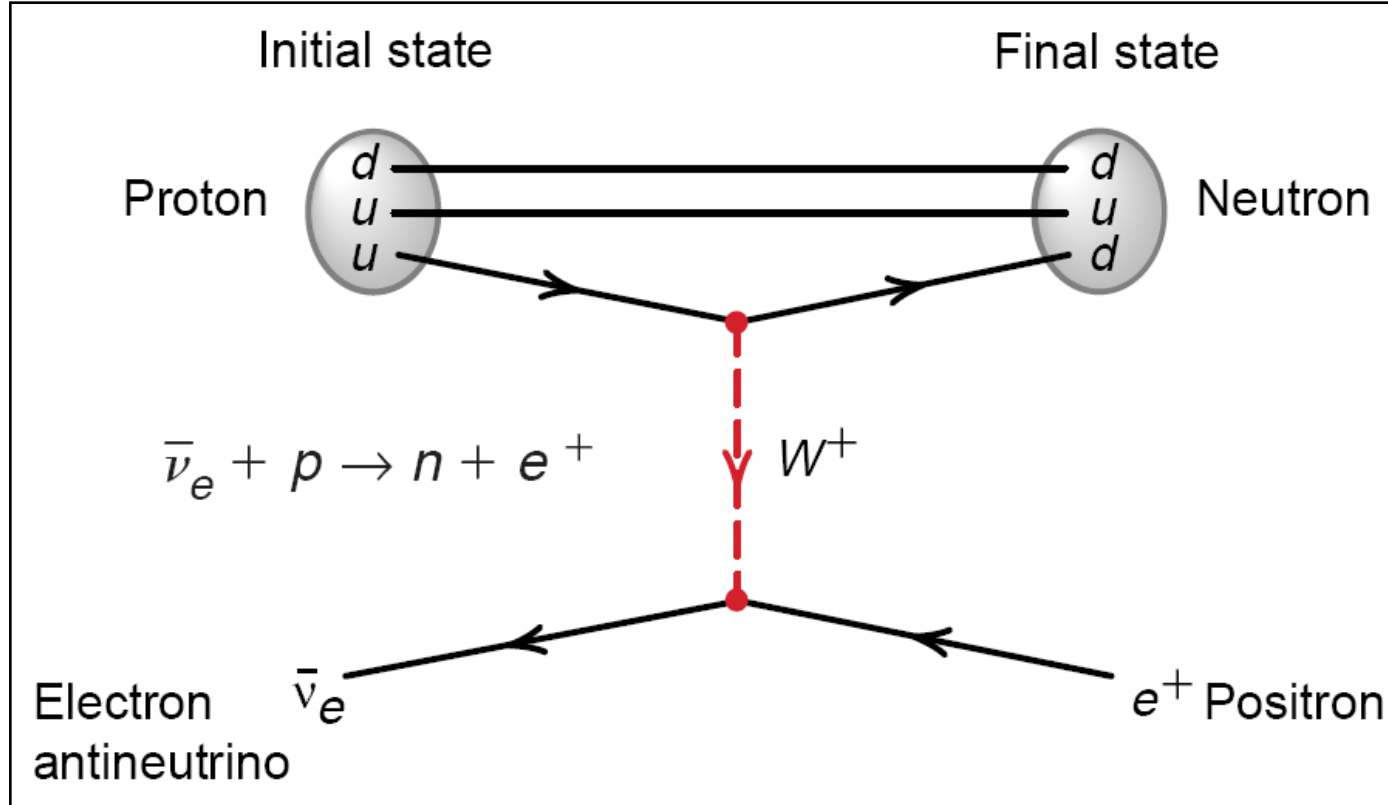
Ettore Majorana, Fermi's PhD student, proposed a new idea in 1937: neutrino is its own antiparticle.



Impossible Challenge

18

An inverse β decay to detect neutrinos (Hans Bethe 1936)



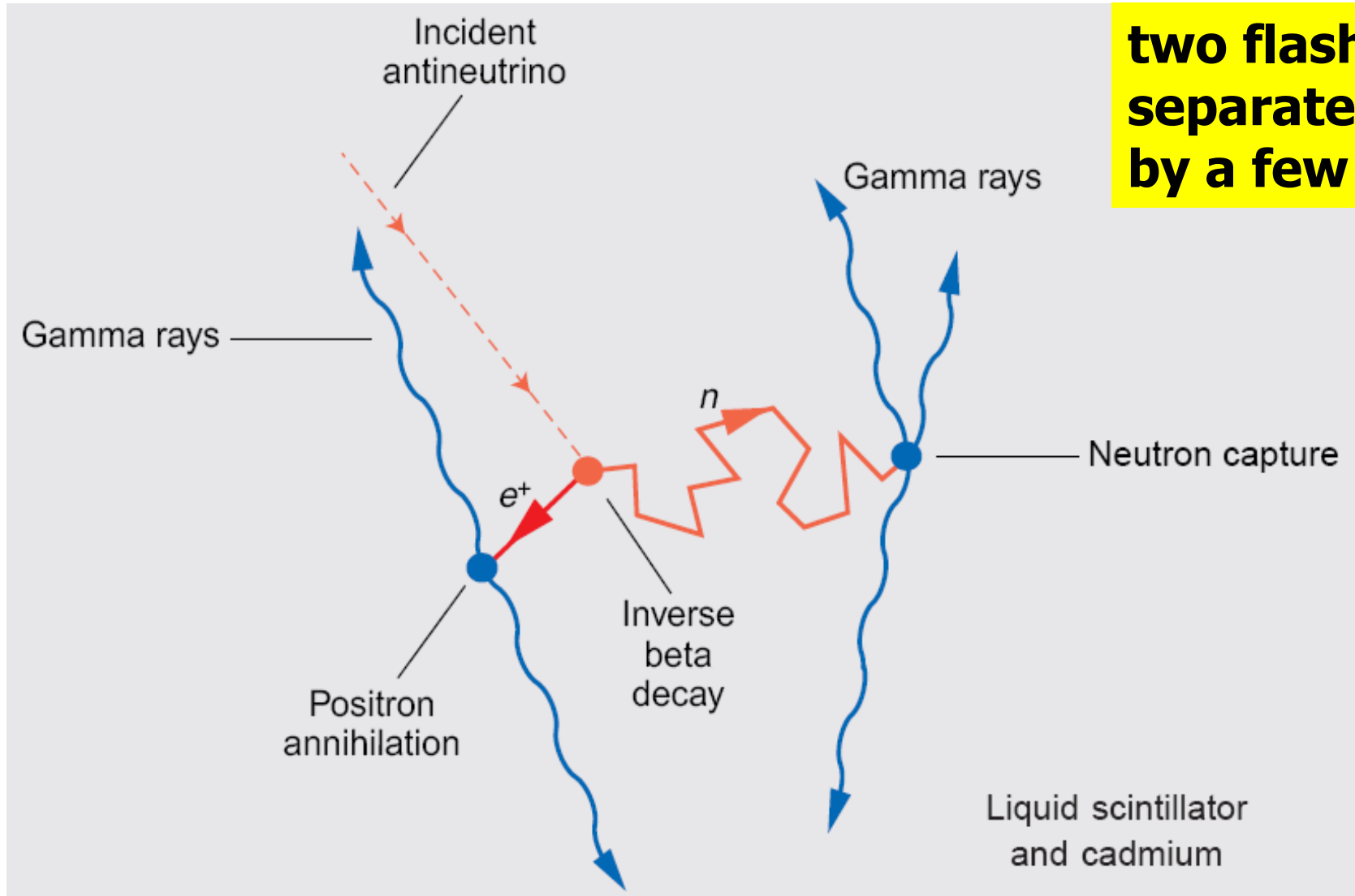
Very intense sources of neutrinos (1950's):
fission bombs and **fission reactors**.

Frederick Reines & Clyde Cowan's Project (1951).

Reactor Neutrinos

19

Decision in 1952: neutrinos from a fission reactor.



**two flashes
separated
by a few μs**

Positive Result?

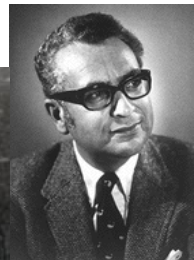
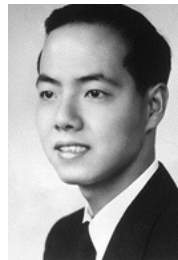
20

Reines and Cowan's telegram to Pauli on 14/06/1956:

We're happy to inform you that we've definitely detected neutrinos from fission fragments by observing inverse β decay of protons. Observed cross section agrees well with expected $6 \times 10^{-44} \text{ cm}^2$.

Such a theoretical value was based on a **parity-conserving** formulation of the β decay with **4** independent degrees of freedom for **ν 's**.

This value **doubled** after the discovery of **parity violation** in 1957, leading to the two-component **ν** theory in 1957 and the **V-A** theory in 1958.



Nobel Prize

21

A new paper on this experiment published in Phys. Rev. in 1960 reported a cross section **twice** as large as that given in 1956.

Reines (1979): **our initial analysis grossly overestimated the detection efficiency with the result that the measured cross section was at first thought to be in good agreement with [the pre-parity violation] prediction.**



Skepticism
{over 39 years}

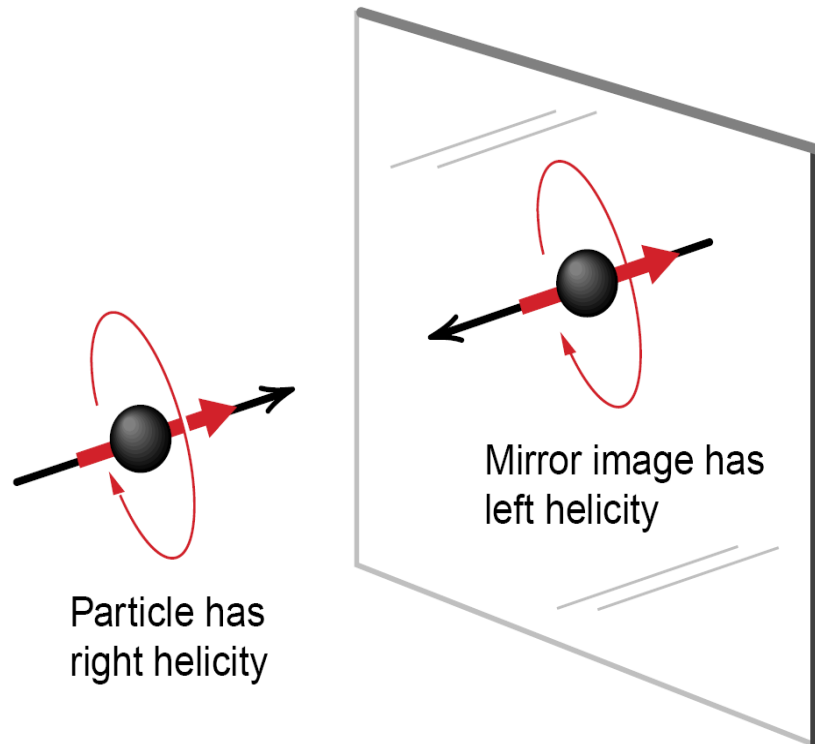
The **Nobel Prize** finally came to Frederick Reines in 1995!

Parity and Helicity

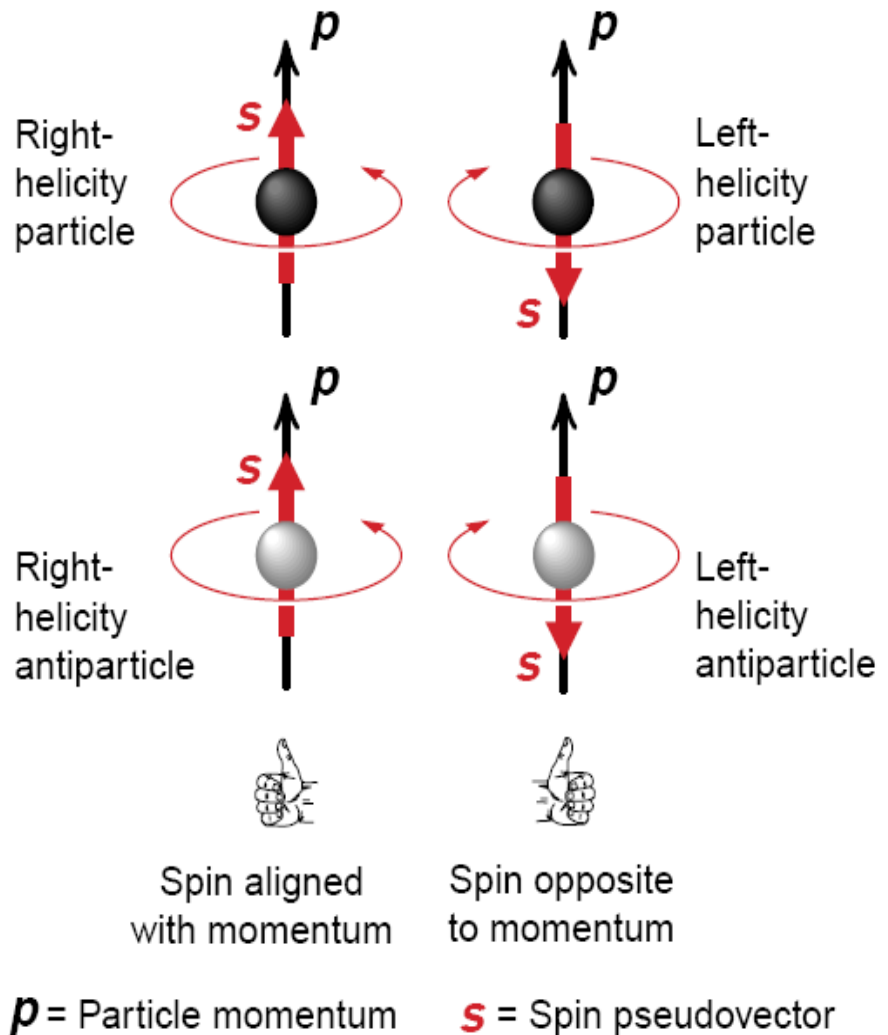
22

Parity and Helicity: if parity were conserved, a spin-1/2 particle would exist in both left- & right-helicity states.

(b) Mirror Reflection of a Right-Helicity Particle



(a) Four States of a Spin-1/2 Particle



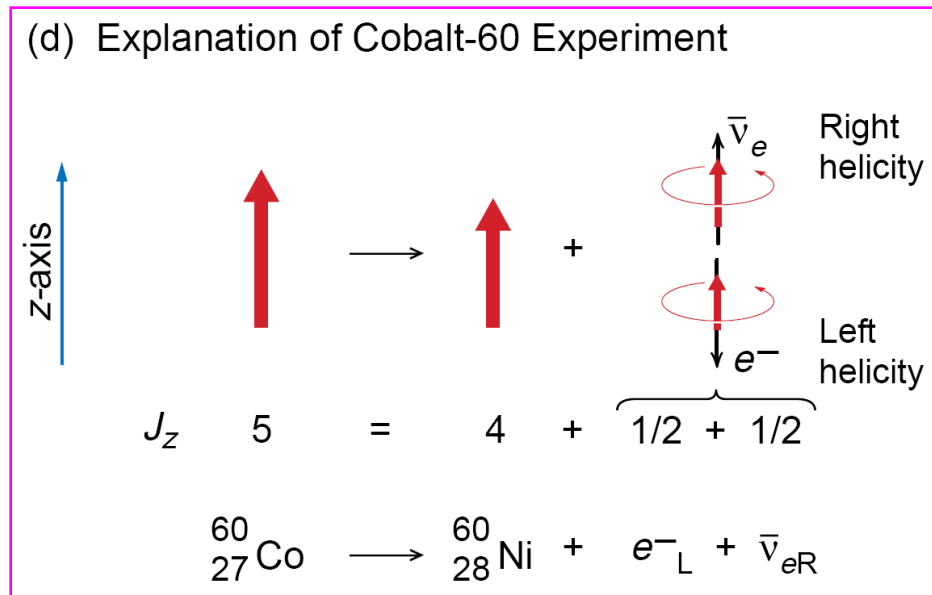
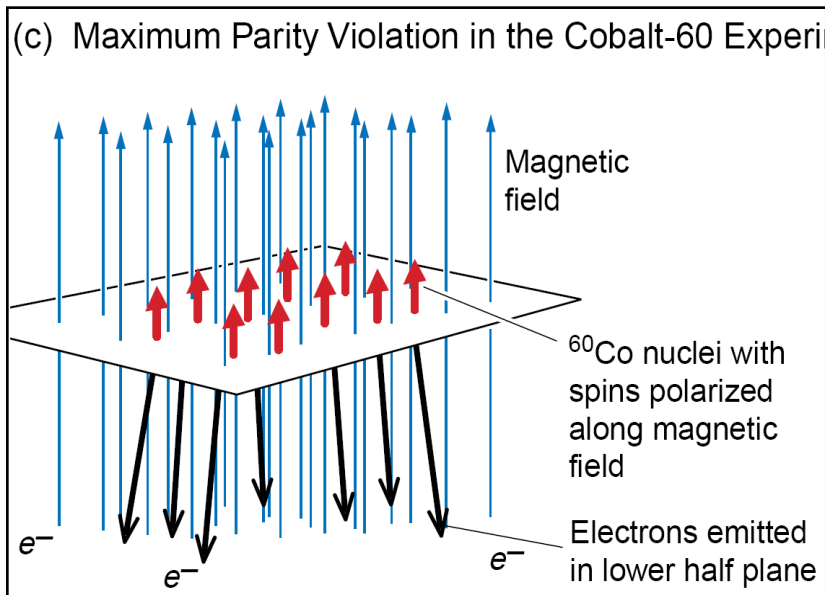
Wu's Experiment

23

C.S. Wu aligned the spins of Cobalt-60 nuclei along an external magnetic field and then measured the directions of the electrons emitted by those nuclei. She observed **maximal parity violation!**



Lee & Yang's two-component neutrino theory explained the data but implied: **neutrinos are left-handed & massless!**



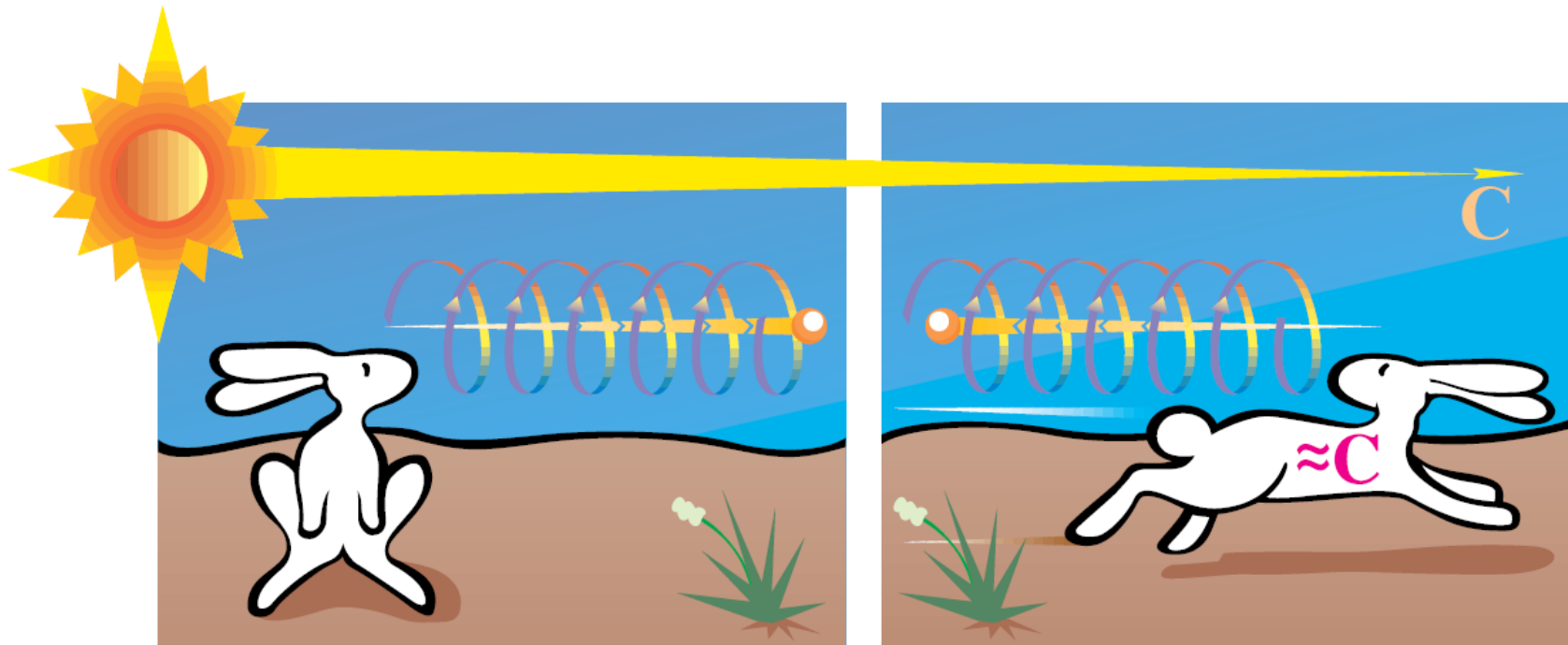
Helicity

24

The **Helicity** of a fermion relates its spin to its direction of motion

$$\lambda = \mathbf{s} \cdot \mathbf{p} / |\mathbf{p}| = \pm 1/2$$

A massive particle's **helicity** isn't relativistically invariant.



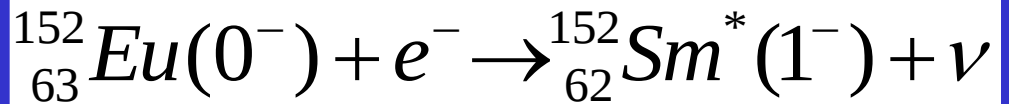
Looks like a left-handed corkscrew.

No—like a right-handed corkscrew!

The GGS Experiment

25

A proof of the **negative helicity** of a neutrino was done in 1958.



In this reaction, the Sm^* nucleus recoils from the neutrino. It decays into the ground state simply by emitting a γ particle. Then the polarization of this γ particle measures the **neutrino's helicity** in a rather non-trivial way.

$$\lambda = -1/2$$



Helicity of Neutrinos*

M. GOLDHABER, L. GRODZINS, AND A. W. SUNYAR

Brookhaven National Laboratory, Upton, New York

(Received December 11, 1957)

A COMBINED analysis of circular polarization and resonant scattering of γ rays following orbital electron capture measures the helicity of the neutrino. We have carried out such a measurement with Eu^{152m} , which decays by orbital electron capture. If we assume the most plausible spin-parity assignment for this isomer compatible with its decay scheme,¹ 0^- , we find that the neutrino is “left-handed,” i.e., $\sigma_\nu \cdot \hat{p}_\nu = -1$ (negative helicity).

Handedness

26

The **Handedness** of a spin-1/2 particle is a relativistically invariant quantity to describe this particle's spin states.

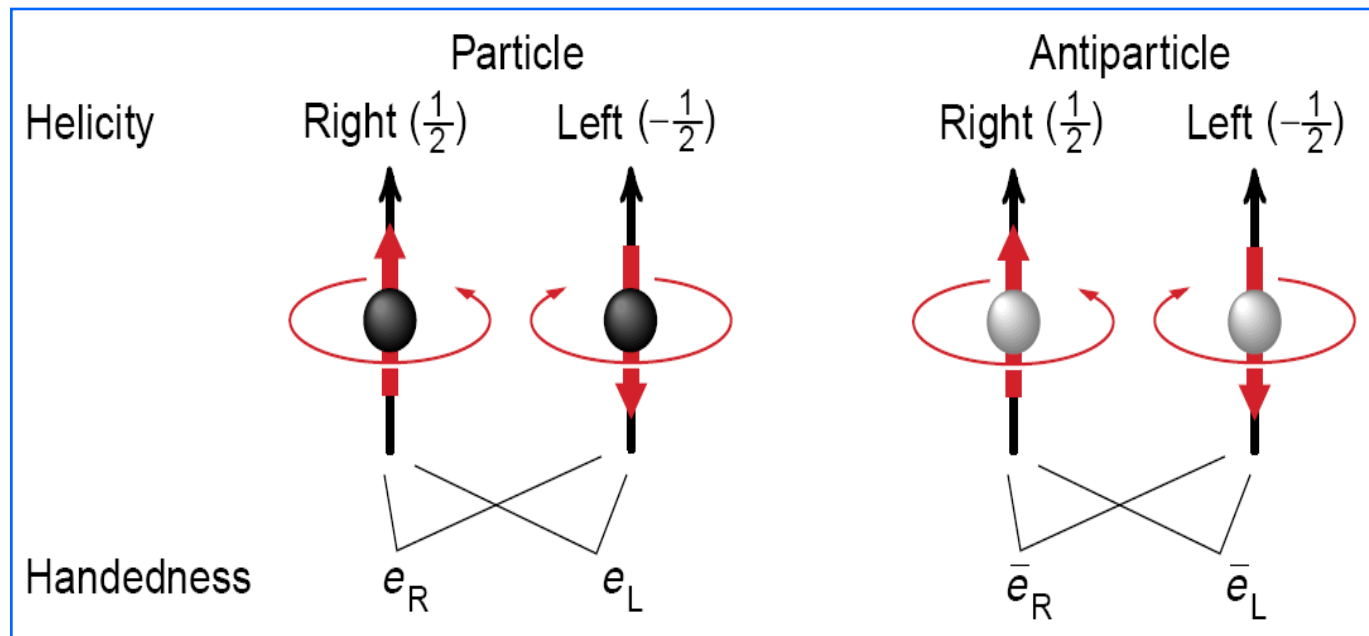
2 independent handedness or **chirality** states: **left** & **right**.

A massless or massive particle can always be decomposed into two independent components: **left-** and **right-handed**.

$$\psi = \psi_L + \psi_R$$

$$\psi_L = \frac{1 + \gamma_5}{2} \psi$$

$$\psi_R = \frac{1 - \gamma_5}{2} \psi$$



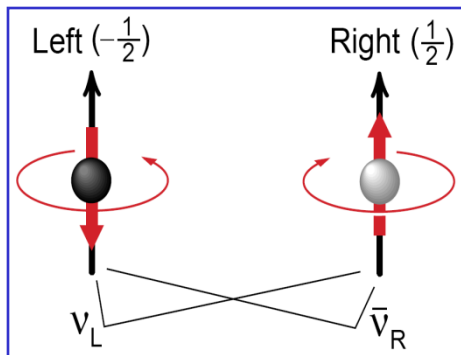
Weak States

27

The **Handedness** is an essential quantity for describing the weak force (see the table for the 1st-family in the SM).

For example:

The right-handed electron is a weak isospin singlet: it does not interact with the **W** boson, but it does couple to the photon and to the **Z** boson.

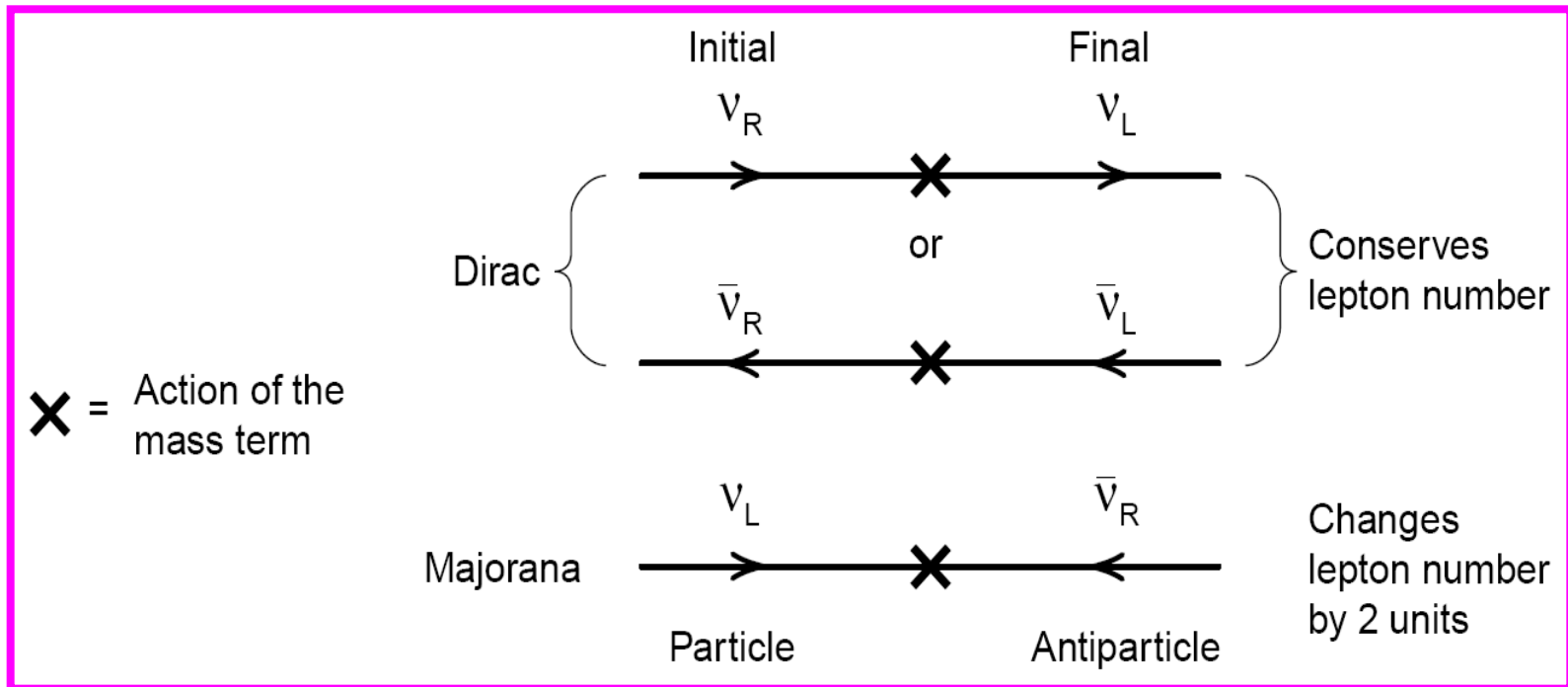


Particle Number N	Handedness N_x	Particle States	Weak Isotopic Charge I_3^w	Weak Hypercharge Y^w	Electric Charge Q
+1	L	$\begin{pmatrix} e_L \\ \nu_L \end{pmatrix}$	-1/2	-1	-1
+1	L		+1/2	-1	0
-1	R	$\begin{pmatrix} \bar{e}_R \\ \bar{\nu}_R \end{pmatrix}$	+1/2	+1	+1
-1	R		-1/2	+1	0
+1	R	e_R	0	-2	-1
-1	L	$\bar{e}_L$	0	+2	+1
+1	R	ν_R	0	0	0
-1	L	$\bar{\nu}_L$	0	0	0

Mass Terms

28

The **Handedness** is also a crucial concept for the origin of particle masses. Any interaction giving a spin-1/2 particle a nonzero mass must flip its left- and right-handed states.



$$\begin{aligned} |e_L\rangle &\propto |e_{-1/2}\rangle + m/E |e_{+1/2}\rangle \\ |e_R\rangle &\propto |e_{+1/2}\rangle - m/E |e_{-1/2}\rangle \end{aligned}$$

2 handed states in terms of **2 helicity states**:
if $m = 0$, handedness = helicity; if $m \ll E$, it holds approximately (in a relativistic limit).

Helicity + Handedness

29

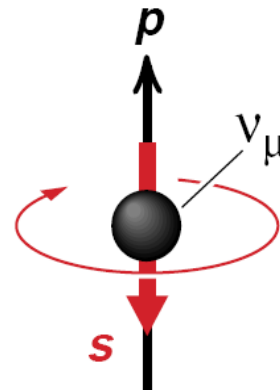
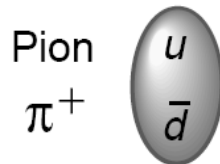
Which decay is more likely?

$$\pi^+ \rightarrow e^+ + \nu_e$$

$$\pi^+ \rightarrow \mu^+ + \nu_\mu$$

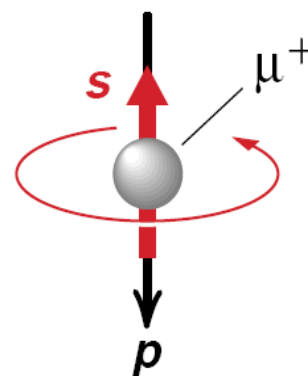
$$\begin{aligned} |\mu_L^+\rangle &\propto |\mu_{-1/2}^+\rangle + m/E |\mu_{+1/2}^+\rangle \\ |\mu_R^+\rangle &\propto |\mu_{+1/2}^+\rangle - m/E |\mu_{-1/2}^+\rangle \end{aligned}$$

z-axis



$$J_z = s_z = -\frac{1}{2}$$

The massless, left-handed neutrino is in a state of left helicity.

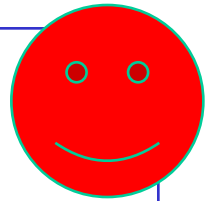


$$J_z = s_z = \frac{1}{2}$$

The right-handed antimuon must be produced in a state of left helicity.

Angular momentum

$$J_z = 0 \longrightarrow J_z = -\frac{1}{2} + \frac{1}{2} = 0$$



Electron & Its Neutrino

30

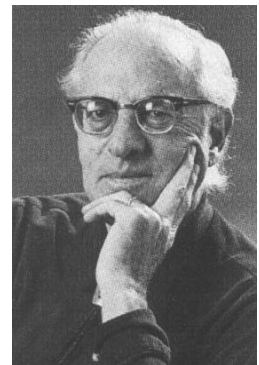
The **electron** was discovered in 1897 by **Joseph Thomson**.

The electron's anti-particle, **positron**, was discovered in 1932 by **Carl Anderson**.



The first scientist deemed to have captured **positrons** via electron-positron annihilation was **Chung-Yao Chao** (赵忠尧 1902–1998) at Caltech in 1930.

In 1956 **Clyde Cowan** and **Frederick Reines** discovered the positron's partner, **electron antineutrino**.



Muon

31

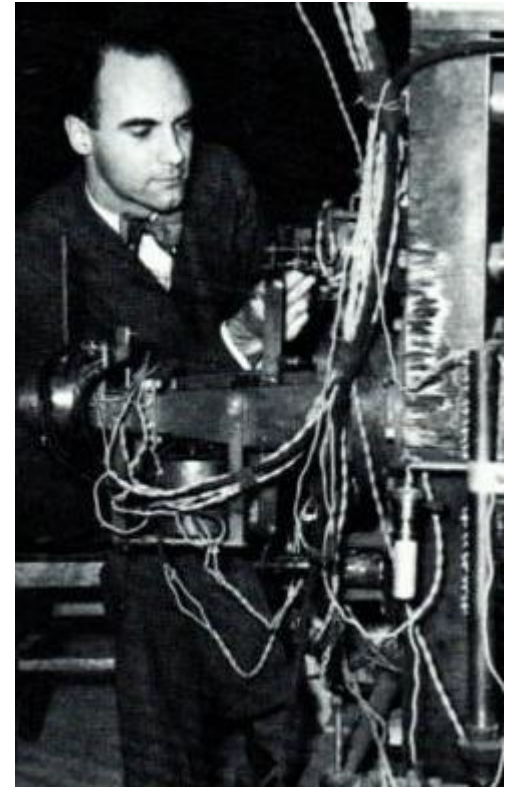
The **muon** particle, a sister of the electron, was discovered in 1936 by **Carl Anderson** and his first student **S. Neddermeyer**; and independently by **J. Street *et al.***

It was not **Hideki Yukawa's** "pion". And it was the first flavor puzzle.

Isidor Rabi famously asked:
Who ordered that?



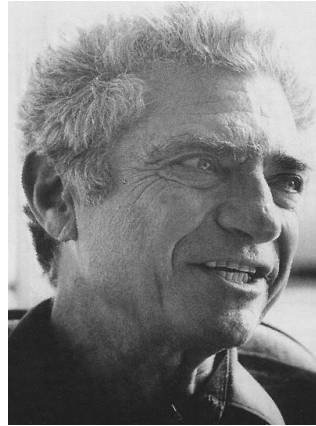
Isidor Isaac Rabi



Muon Neutrino

32

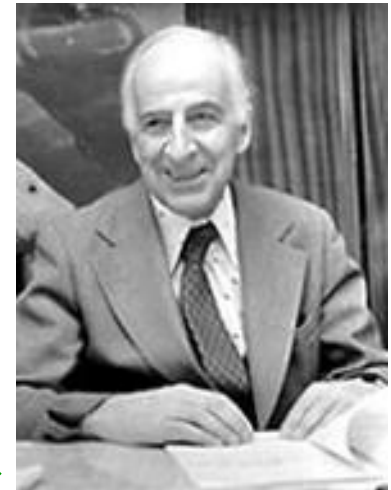
The **muon neutrino** was discovered by **Leon Lederman**, **Melvin Schwartz** and **Jack Steinberger** in 1962.



Neutrino flavor conversion was proposed by **Z. Maki**, **M. Nakagawa** and **S. Sakata** in 1962.



Neutrinos oscillate into **antineutrinos**: first proposed by **Bruno Pontecorvo** in 1957.

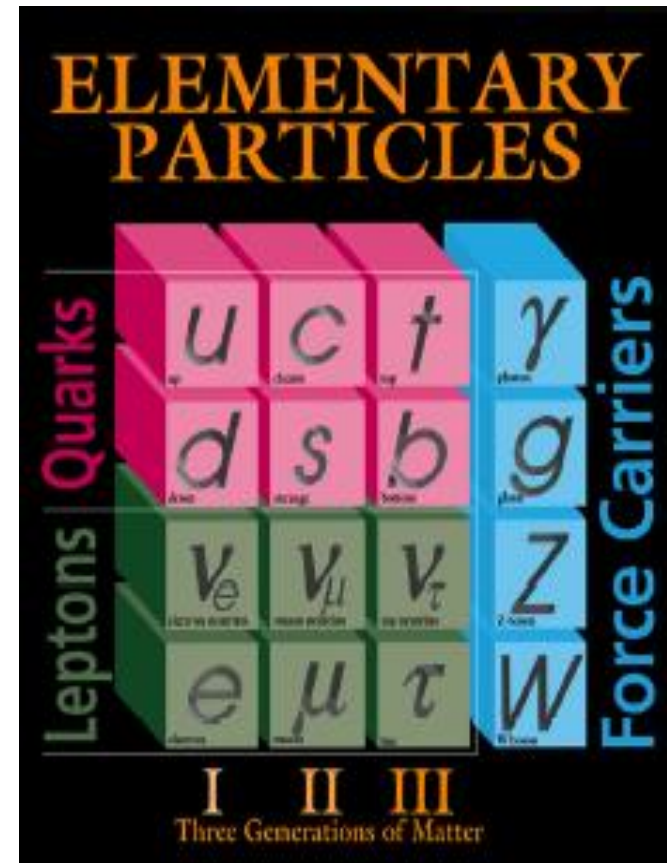
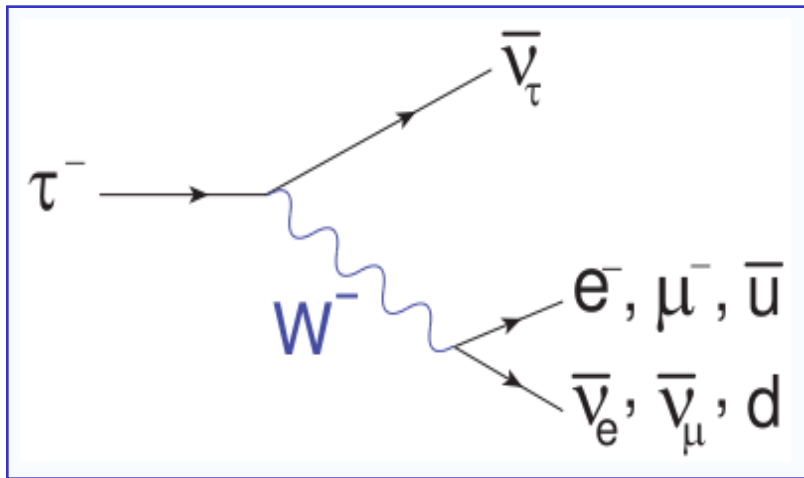


Tau & Tau Neutrino

33

The **tau** particle was discovered by **Martin Perl** in 1975 via:

$$e^+ + e^- \rightarrow e^\pm + \mu^\mp + \text{undetected particles}$$



In 2000, the **tau neutrino** was finally discovered at the Fermilab.

The lepton family is complete!

Leptons & Nobel Prizes

34

e	J.J. Thomson 1897 😊	J.J. Thomson 1906 (NP)
ν_e	C.L. Cowan et al. 1956	F.J. Reines 1995 (NP)
μ	J.C. Street et al. C.D. Anderson 1936 😊	1975 – 1936 = 1936 – 1897 = 39
ν_μ	G. Danby et al. 1962	M. Schwartz, L.M. Lederman, J. Steinberger 1988 (NP)
τ	M.L. Perl et al. 1975 😊	M.L. Perl 1995 (NP)
ν_τ	K. Kodama et al. 2000	

Antimatter: Positron.

Predicted by P.A.M. Dirac in 1928.

Discovered by C.D. Anderson in 1932; Nobel Prize in 1936.

Sarma-Xing Theorem

35

In 1995 it was an Indian theorist who first discovered the **39-year gap** of charged leptons.

2) NOBEL LEPTONS.

By **K.V.L. Sarma** (**Tata Inst.**),. TIFR-TH-95-56, Dec 1995. 13pp.

Submitted to Curr. Sci.

e-Print Archive: **hep-ph/9512420**

$$1975 + 39 = 2014$$

A summary of the discoveries made in the world of leptons is given in Table 1. We see that the third generation has started getting Nobel prizes. It is amusing that the charged-leptons crop up with a **39-year gap** and may be the 4th one would show up in the year **2114**. For the present, the available experimental information implies that there are no charged leptons which are heavier than tau and lighter than 45 GeV.

My contribution: corrected **2114** to **2014**, so the discovery would be possible **100** years earlier (two years later)!

Lecture B

- ★ **Massless Neutrinos in the SM**
- ★ **Dirac and Majorana Mass Terms**
- ★ **The Seesaw Mechanisms**

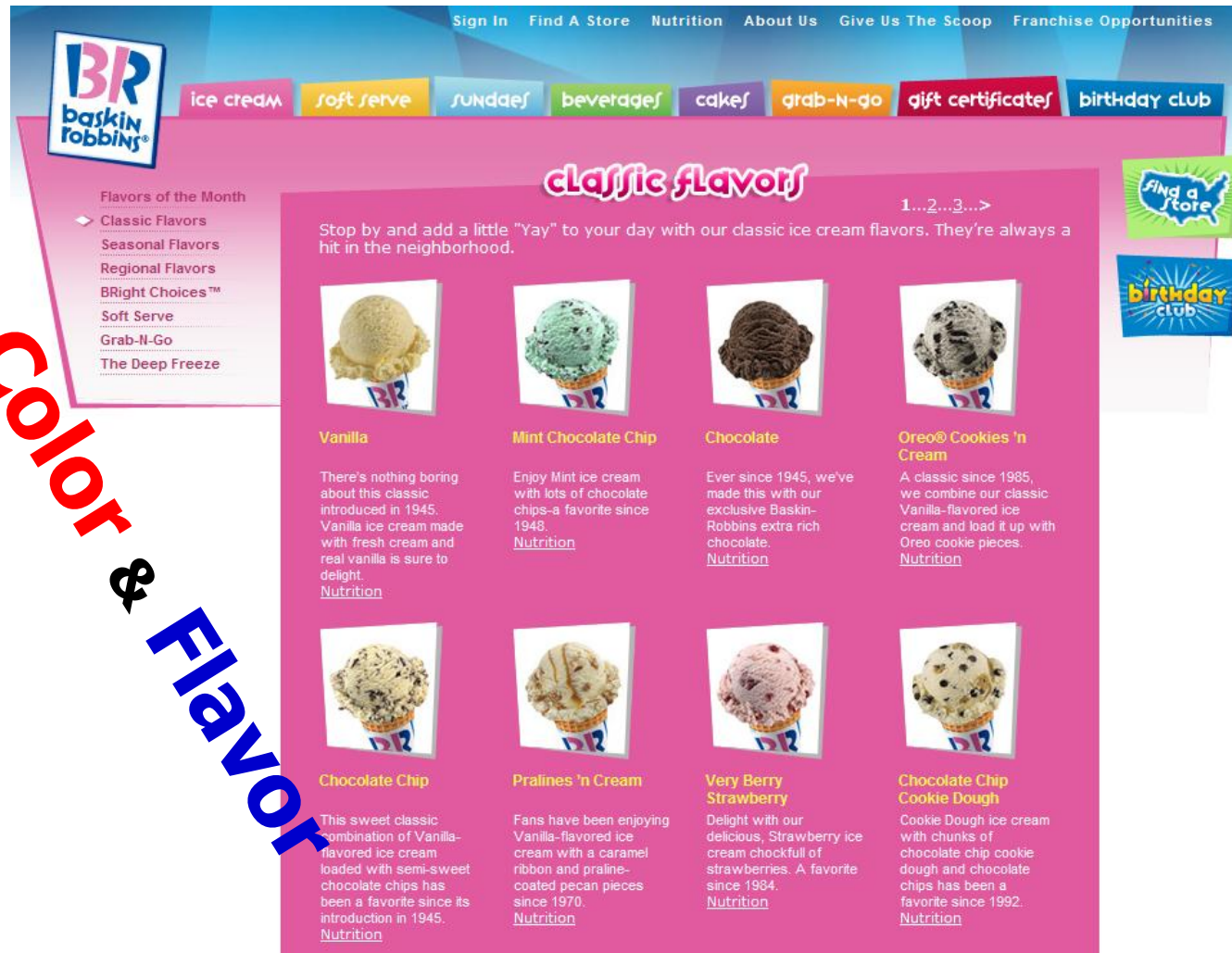
Origin of "Flavor"

37

The term **Flavor** was coined by **Harald Fritzsch** and **Murray Gell-Mann** at an ice-cream store in Pasadena in **1971**.

Flavor issues:
fermion mass
flavor mixing
CP violation

Color & Flavor



The screenshot shows the Baskin-Robbins website with a navigation bar at the top containing links: Sign In, Find A Store, Nutrition, About Us, Give Us The Scoop, and Franchise Opportunities. Below the navigation bar is a menu with categories: ice cream, soft serve, sundaes, beverages, cakes, grab-n-go, gift certificates, and birthday club. The main content area is titled "classic flavors" and features a grid of eight ice cream flavors, each with a photo of the ice cream and a description. The flavors are: Vanilla, Mint Chocolate Chip, Chocolate, Oreo® Cookies 'n Cream, Chocolate Chip, Pralines 'n Cream, Very Berry Strawberry, and Chocolate Chip Cookie Dough. Each flavor description includes a brief history and a link to "Nutrition".

Sign In Find A Store Nutrition About Us Give Us The Scoop Franchise Opportunities

ice cream soft serve sundaes beverages cakes grab-n-go gift certificates birthday club

classic flavors

1...2...3...>

Stop by and add a little "Yay" to your day with our classic ice cream flavors. They're always a hit in the neighborhood.

Vanilla

There's nothing boring about this classic introduced in 1945. Vanilla ice cream made with fresh cream and real vanilla is sure to delight.

[Nutrition](#)

Mint Chocolate Chip

Enjoy Mint ice cream with lots of chocolate chips—a favorite since 1948.

[Nutrition](#)

Chocolate

Ever since 1945, we've made this with our exclusive Baskin-Robbins extra rich chocolate.

[Nutrition](#)

Oreo® Cookies 'n Cream

A classic since 1985, we combine our classic Vanilla-flavored ice cream and load it up with Oreo cookie pieces.

[Nutrition](#)

Chocolate Chip

This sweet classic combination of Vanilla-flavored ice cream loaded with semi-sweet chocolate chips has been a favorite since its introduction in 1945.

[Nutrition](#)

Pralines 'n Cream

Fans have been enjoying Vanilla-flavored ice cream with a caramel ribbon and praline-coated pecan pieces since 1970.

[Nutrition](#)

Very Berry Strawberry

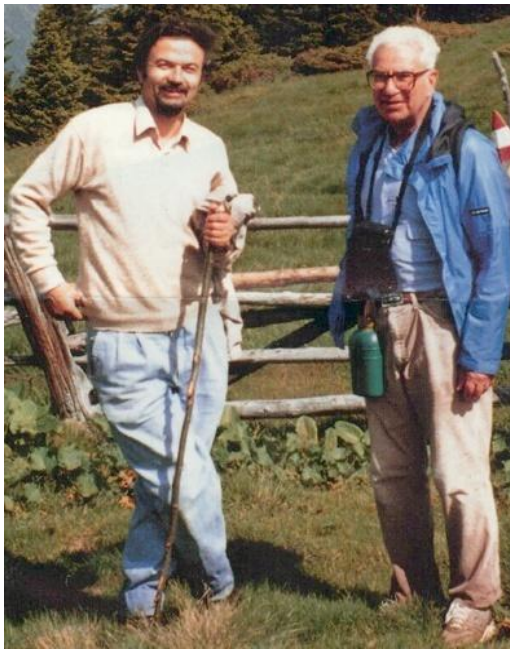
Delight with our delicious, Strawberry ice cream chockfull of strawberries. A favorite since 1984.

[Nutrition](#)

Chocolate Chip Cookie Dough

Cookie Dough ice cream with chunks of chocolate chip cookie dough and chocolate chips has been a favorite since 1992.

[Nutrition](#)



What is Mass?

38

Higgs is the **God** particle in the SM: it gives masses to other particles.

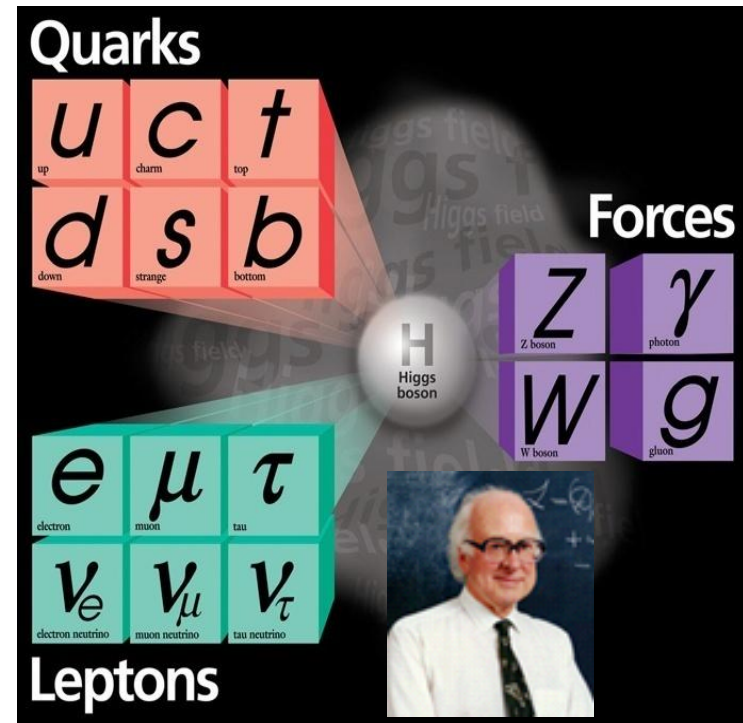
A lesson learnt from history:

All known **bosons** were discovered in **Europe** and most **fermions** were discovered in **America**.

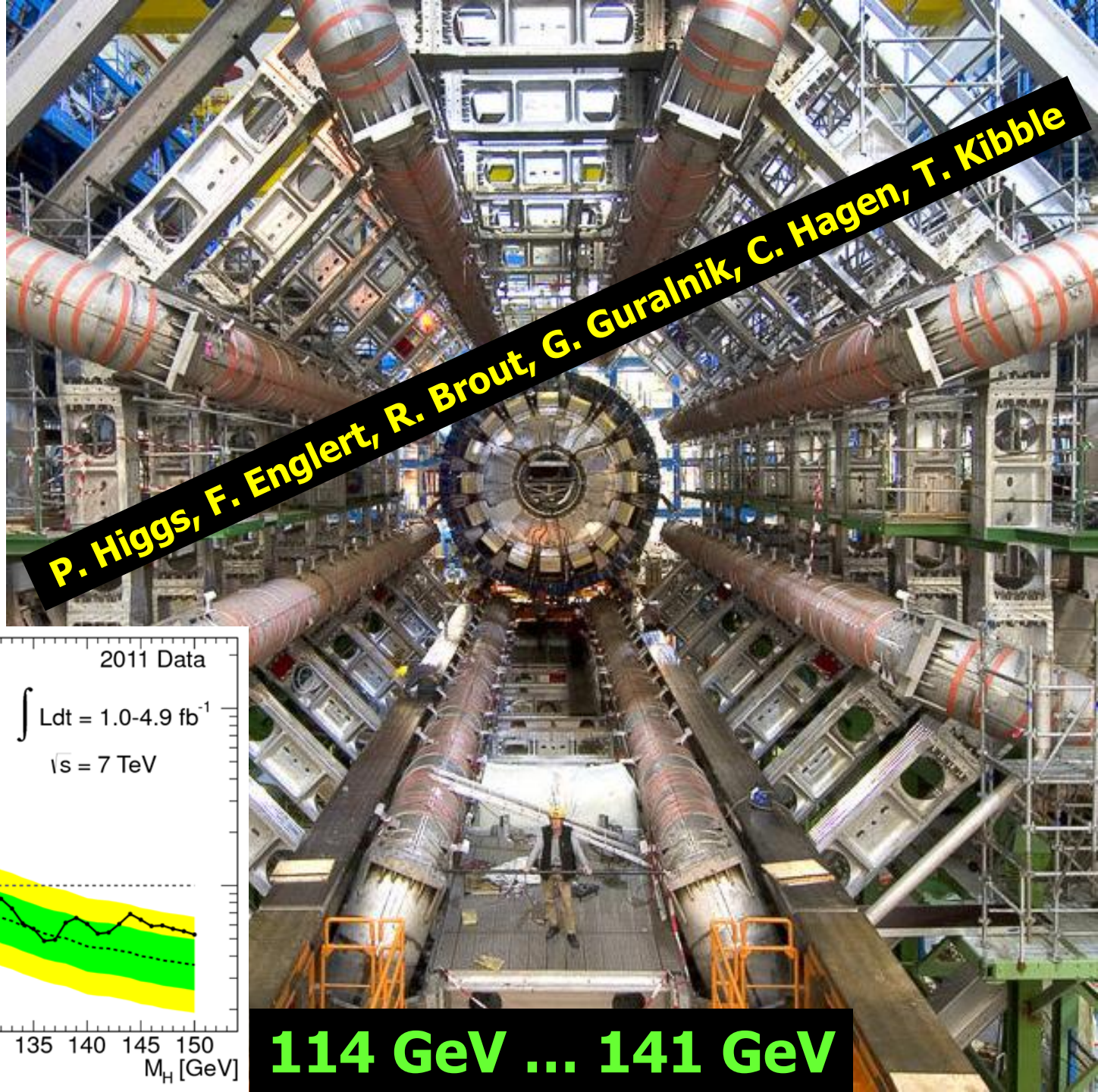
The LHC is very likely to see Higgs.

Mass is the inertial energy of a particle existing at rest.

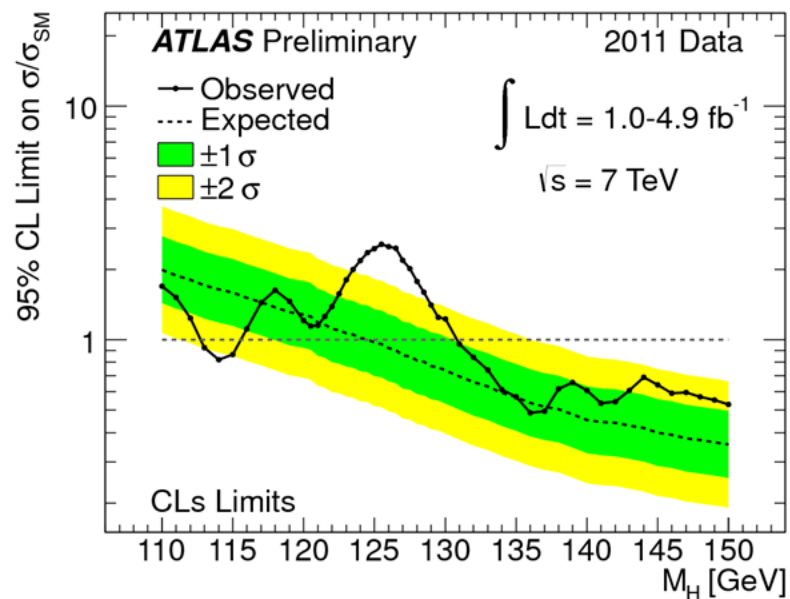
- A **massless** particle has no way to exist at rest. It must always move at the speed of light.
- A **massive** fermion (lepton or quark) must exist in both the left- and right-handed states.



2012 Higgs



P. Higgs, F. Englert, R. Brout, G. Guralnik, C. Hagen, T. Kibble



114 GeV ... 141 GeV

In the SM

40

All ν 's are **massless** due to the model's simple structure:

---- $SU(2) \times U(1)$ **gauge symmetry and Lorentz invariance**

Fundamentals of a quantum field theory

---- Economical **particle content**:

No right-handed neutrino; only a single Higgs doublet

---- Mandatory **renormalizability**:

No dimension ≥ 5 operator (**$B-L$** conserved in the SM)

Neutrinos are **massless** in the SM: Natural or not?

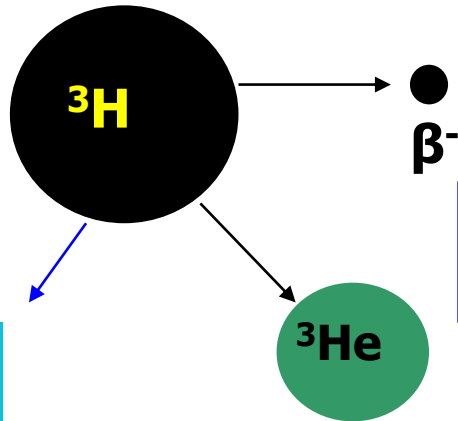
YES: It's tooooooooo light and almost left-handed;

NO: No fundamental symmetry/conservation law.

Constraints on Masses

41

Way 1: Tritium beta decay



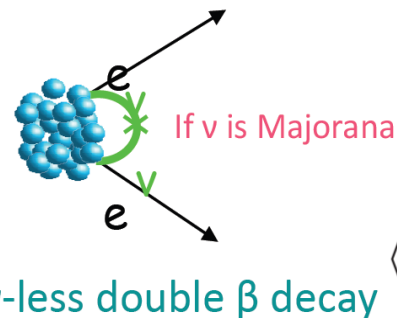
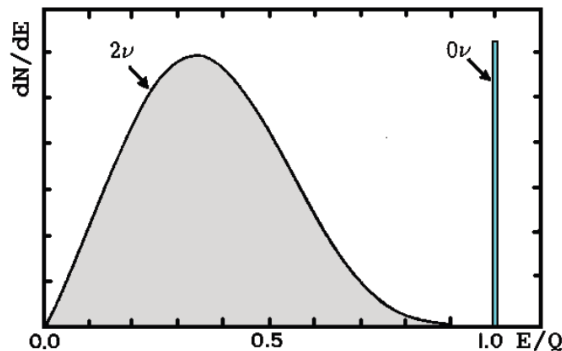
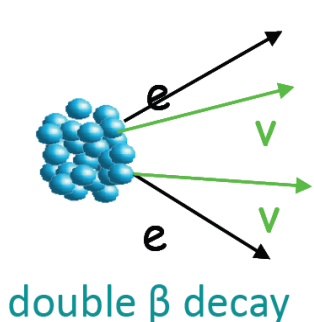
$$\langle m \rangle_e^2 \equiv \sum_i |V_{ei}|^2 m_i^2$$

$< 1 \text{ eV}$

Way 3: Cosmology



Way 2: Neutrinoless $\beta\beta$ decay



$$\langle m \rangle_{\beta\beta} \equiv \left| \sum_{i=1}^3 V_{ei}^2 m_i \right|$$

$$\langle m \rangle_{\beta\beta}^2 = m_e^2 [C_{mm} T_{1/2}^{0\nu}]^{-1}$$

Some Notations

42

Define the **left-** and **right-**handed neutrino fields:

$$\nu_L = \begin{pmatrix} \nu_{eL} \\ \nu_{\mu L} \\ \nu_{\tau L} \end{pmatrix} \quad \blacksquare \quad N_R = \begin{pmatrix} N_{1R} \\ N_{2R} \\ N_{3R} \end{pmatrix}$$

**Extend the SM's
particle content**

$$\begin{aligned} \psi_L &\equiv \frac{1 - \gamma_5}{2} \psi \\ \psi_R &\equiv \frac{1 + \gamma_5}{2} \psi \end{aligned}$$

Their **charge-conjugate counterparts** are defined below and transform as **right-** and **left-**handed fields, respectively:

$$(\nu_L)^c \equiv \mathcal{C} \overline{\nu_L}^T, \quad (N_R)^c \equiv \mathcal{C} \overline{N_R}^T$$

$$\overline{(\nu_L)^c} = (\nu_L)^T \mathcal{C}, \quad \overline{(N_R)^c} = (N_R)^T \mathcal{C}$$

$$(\nu_L)^c = (\nu^c)_R \text{ and } (N_R)^c = (N^c)_L \text{ hold} \quad \text{(can be proved easily)}$$

Properties of the charge-conjugation matrix:

$$\mathcal{C} \gamma_\mu^T \mathcal{C}^{-1} = -\gamma_\mu, \quad \mathcal{C} \gamma_5^T \mathcal{C}^{-1} = \gamma_5, \quad \mathcal{C}^{-1} = \mathcal{C}^\dagger = \mathcal{C}^T = -\mathcal{C}$$

They are from the requirement that **the charge-conjugated field** must satisfy the same **Dirac** equation ($\mathcal{C} = i\gamma^2\gamma^0$ in the **Dirac** representation)

Dirac Mass Term

43

A **Dirac** neutrino is described by a **4**-component spinor: $\nu = \nu_L + N_R$

Step 1: the gauge-invariant **Dirac** mass term and **SSB**:

$$-\mathcal{L}_{\text{Dirac}} = \bar{\ell}_L Y_\nu \tilde{H} N_R + \text{h.c.}$$



$$-\mathcal{L}'_{\text{Dirac}} = \bar{\nu}_L M_D N_R + \text{h.c.}$$

$$M_D = Y_\nu \langle H \rangle \text{ with } \langle H \rangle \simeq 174 \text{ GeV}$$

Step 2: basis transformation:

$$V^\dagger M_D U = \widehat{M}_\nu \equiv \text{Diag}\{m_1, m_2, m_3\}$$

$$-\mathcal{L}'_{\text{Dirac}} = \bar{\nu}'_L \widehat{M}_\nu N'_R + \text{h.c.}$$

Mass states link to flavor states:

$$\nu'_L = V^\dagger \nu_L \text{ and } N'_R = U^\dagger N_R$$

$$\nu' = \nu'_L + N'_R = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Step 3: physical mass term and kinetic term:

$$-\mathcal{L}'_{\text{Dirac}} = \bar{\nu}' \widehat{M}_\nu \nu' = \sum_{i=1}^3 m_i \bar{\nu}_i \nu_i$$

$$\mathcal{L}_{\text{kinetic}} = i\bar{\nu}_L \gamma_\mu \partial^\mu \nu_L + i\bar{N}_R \gamma_\mu \partial^\mu N_R = i\bar{\nu}' \gamma_\mu \partial^\mu \nu' = i \sum_{k=1}^3 \bar{\nu}_k \gamma_\mu \partial^\mu \nu_k$$

Dirac Neutrino Mixing

44

Weak charged-current interactions of leptons:

$$\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} (\overline{e \ \mu \ \tau})_L \gamma^\mu \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}_L W_\mu^- + \text{h.c.}$$

In the flavor basis

$$\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} (\overline{e \ \mu \ \tau})_L \gamma^\mu V \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}_L W_\mu^- + \text{h.c.}$$

In the mass basis

Without loss of generality, one may choose mass states=flavor states for charged leptons. So V is just the MNSP matrix of neutrino mixing.

Both the mass & CC terms are invariant with respect to a global phase transformation: lepton number (flavor) conservation (violation).

$$l(x) \rightarrow e^{i\Phi} l(x)$$

$$\nu'_L(x) \rightarrow e^{i\Phi} \nu'_L(x)$$

$$N'_R(x) \rightarrow e^{i\Phi} N'_R(x)$$

	e^-	ν_e	e^+	$\bar{\nu}_e$	μ^-	ν_μ	μ^+	$\bar{\nu}_\mu$	τ^-	ν_τ	τ^+	$\bar{\nu}_\tau$
L	+1	+1	-1	-1	+1	+1	-1	-1	+1	+1	-1	-1
L_e	+1	+1	-1	-1	0	0	0	0	0	0	0	0
L_μ	0	0	0	0	+1	+1	-1	-1	0	0	0	0
L_τ	0	0	0	0	0	0	0	0	+1	+1	-1	-1

Majorana Mass Term (1)

45

A **Majorana** mass term can be obtained by introducing a **Higgs triplet** into the SM, writing out the gauge-invariant Yukawa interactions and Higgs potentials, and then integrating out heavy degrees of freedom (**type-II seesaw mechanism**):

$$-\mathcal{L}'_{\text{Majorana}} = \frac{1}{2} \overline{\nu}_L M_L (\nu_L)^c + \text{h.c.}$$

The **Majorana** mass matrix must be a **symmetric** matrix. It can be diagonalized by a **unitary** matrix

$$\overline{\nu}_L M_L (\nu_L)^c = [\overline{\nu}_L M_L (\nu_L)^c]^T = -\overline{\nu}_L C^T M_L^T \overline{\nu}_L^T = \overline{\nu}_L M_L^T (\nu_L)^c$$

Diagonalization:

$$-\mathcal{L}'_{\text{Majorana}} = \frac{1}{2} \overline{\nu}'_L \widehat{M}_\nu (\nu'_L)^c + \text{h.c.}$$

$$V^\dagger M_L V^* = \widehat{M}_\nu \equiv \text{Diag}\{m_1, m_2, m_3\}$$

$$\nu'_L = V^\dagger \nu_L \text{ and } (\nu'_L)^c = C \overline{\nu}'_L^T$$

Physical mass term:

$$-\mathcal{L}'_{\text{Majorana}} = \frac{1}{2} \overline{\nu}' \widehat{M}_\nu \nu' = \frac{1}{2} \sum_{i=1}^3 m_i \overline{\nu}_i \nu_i$$

$$\nu' = \nu'_L + (\nu'_L)^c = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$\text{Majorana condition } (\nu')^c = \nu'$$

Majorana Mass Term (2)

46

Kinetic term (you may prove $\overline{(\psi_L)^c} \gamma_\mu \partial^\mu (\psi_L)^c = \overline{\psi_L} \gamma_\mu \partial^\mu \psi_L$)

$$\mathcal{L}_{\text{kinetic}} = i \overline{\nu_L} \gamma_\mu \partial^\mu \nu_L = i \overline{\nu'_L} \gamma_\mu \partial^\mu \nu'_L = \frac{i}{2} \overline{\nu'} \gamma_\mu \partial^\mu \nu' = \frac{i}{2} \sum_{k=1}^3 \overline{\nu_k} \gamma_\mu \partial^\mu \nu_k$$

Question: why is there a factor **1/2** in the **Majorana** mass term?

Answer: it allows us to get the correct **Dirac** equation of motion.

A proof: write out the Lagrangian of free massive **Majorana** neutrinos:

$$\begin{aligned} \mathcal{L}_\nu &= i \overline{\nu_L} \gamma_\mu \partial^\mu \nu_L - \left[\frac{1}{2} \overline{\nu_L} M_L (\nu_L)^c + \text{h.c.} \right] \\ &= i \overline{\nu'_L} \gamma_\mu \partial^\mu \nu'_L - \left[\frac{1}{2} \overline{\nu'_L} \widehat{M}_\nu (\nu'_L)^c + \text{h.c.} \right] \\ &= \frac{1}{2} \left(i \overline{\nu'} \gamma_\mu \partial^\mu \nu' - \overline{\nu'} \widehat{M}_\nu \nu' \right) = -\frac{1}{2} \left(i \partial^\mu \overline{\nu'} \gamma_\mu \nu' + \overline{\nu'} \widehat{M}_\nu \nu' \right) \end{aligned}$$



Euler-Lagrange equation:

$$\partial^\mu \frac{\partial \mathcal{L}_\nu}{\partial (\partial^\mu \overline{\nu'})} - \frac{\partial \mathcal{L}_\nu}{\partial \overline{\nu'}} = 0$$



$$i \gamma_\mu \partial^\mu \nu' - \widehat{M}_\nu \nu' = 0$$

$$i \gamma_\mu \partial^\mu \nu_k - m_k \nu_k = 0$$

Majorana Neutrino Mixing 47

Weak charged-current interactions of leptons:

$$\mathcal{L}_{\text{cc}} = \frac{g}{\sqrt{2}} (\bar{e} \ \bar{\mu} \ \bar{\tau})_{\text{L}} \gamma^{\mu} \begin{pmatrix} \nu_e \\ \nu_{\mu} \\ \nu_{\tau} \end{pmatrix}_{\text{L}} W_{\mu}^{-} + \text{h.c.}$$

In the flavor basis

$$\mathcal{L}_{\text{cc}} = \frac{g}{\sqrt{2}} (\bar{e} \ \bar{\mu} \ \bar{\tau})_{\text{L}} \gamma^{\mu} V \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}_{\text{L}} W_{\mu}^{-} + \text{h.c.}$$

In the mass basis

The **MNSP** matrix **V** contains 2 extra CP-violating phases.

Mass and CC terms are not simultaneously invariant under a global phase transformation --- **Lepton number violation**

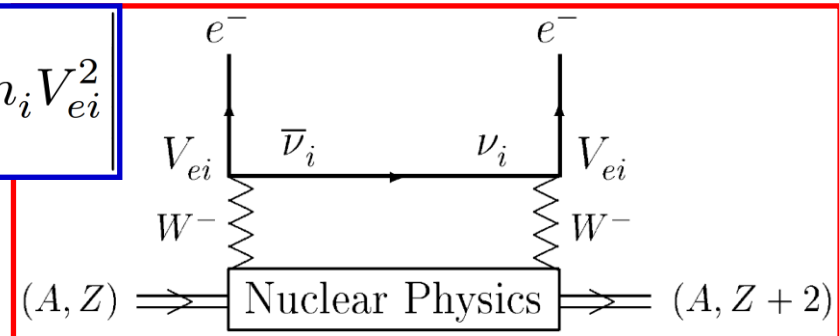
Neutrinoless double-beta decay

$$l(x) \rightarrow e^{i\Phi} l(x)$$

$$\nu'_{\text{L}}(x) \rightarrow e^{i\Phi} \nu'_{\text{L}}(x)$$

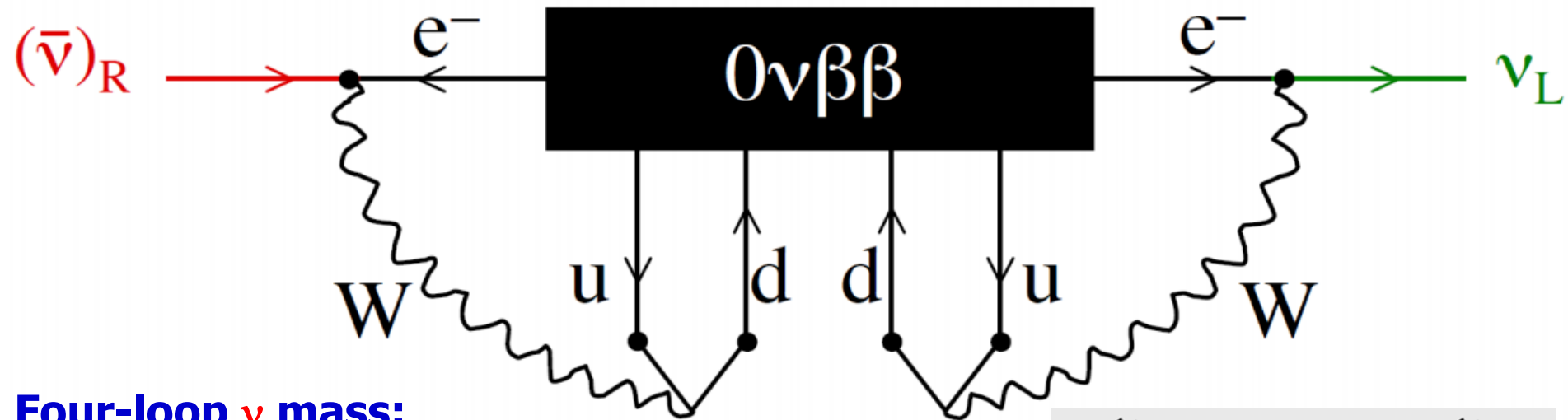
$$\overline{\nu'_{\text{L}}} \rightarrow e^{-i\Phi} \overline{\nu'_{\text{L}}} \text{ and } (\nu'_{\text{L}})^c \rightarrow e^{-i\Phi} (\nu'_{\text{L}})^c$$

$$\langle m \rangle_{ee} \equiv \left| \sum_i m_i V_{ei}^2 \right|$$



Schechter-Valle Theorem 48

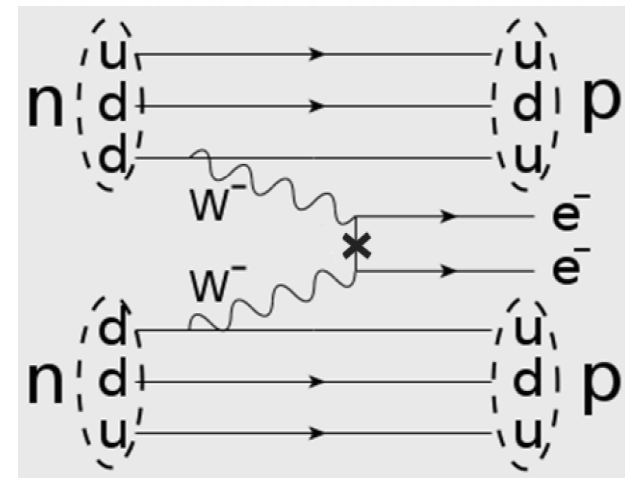
THEOREM (1982): if a $0\nu\beta\beta$ decay happens, there must be an effective **Majorana** mass term.



Four-loop ν mass:

$$\delta m_\nu = \mathcal{O}(10^{-24} \text{ eV}) \quad (\text{Duerr, Lindner, Merle, 2011})$$

Note: The **black box** can in principle have many different processes (new physics). Only in the simplest case, it is likely to constrain neutrino masses.



Hybrid Mass Term (1)

49

A **hybrid** mass term can be written out in terms of the left- and right-handed neutrino fields and their charge-conjugate counterparts:

$$\begin{aligned}
 -\mathcal{L}'_{\text{hybrid}} &= \overline{\nu_L} M_D N_R + \frac{1}{2} \overline{\nu_L} M_L (\nu_L)^c + \frac{1}{2} \overline{(N_R)^c} M_R N_R + \text{h.c.} \\
 &= \frac{1}{2} \begin{bmatrix} \overline{\nu_L} & \overline{(N_R)^c} \end{bmatrix} \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{bmatrix} (\nu_L)^c \\ N_R \end{bmatrix} + \text{h.c.} ,
 \end{aligned}$$

type-(I+II) seesaw

Here we have used

Diagonalization by means of a **6×6 unitary** matrix:

$$\overline{(N_R)^c} M_D^T (\nu_L)^c = [(N_R)^T \mathcal{C} M_D^T \mathcal{C} \overline{\nu_L}^T]^T = \overline{\nu_L} M_D N_R$$

$$\begin{pmatrix} V & R \\ S & U \end{pmatrix}^\dagger \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{pmatrix} V & R \\ S & U \end{pmatrix}^* = \begin{pmatrix} \widehat{M}_\nu & \mathbf{0} \\ \mathbf{0} & \widehat{M}_N \end{pmatrix}$$

$$\widehat{M}_\nu \equiv \text{Diag}\{m_1, m_2, m_3\}, \quad \widehat{M}_N \equiv \text{Diag}\{M_1, M_2, M_3\}$$

It is actually a Majorana mass term!

$$-\mathcal{L}'_{\text{hybrid}} = \frac{1}{2} \begin{bmatrix} \overline{\nu'_L} & \overline{(N'_R)^c} \end{bmatrix} \begin{pmatrix} \widehat{M}_\nu & \mathbf{0} \\ \mathbf{0} & \widehat{M}_N \end{pmatrix} \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} + \text{h.c.}$$

Majorana mass states

$$\begin{aligned}
 \nu' &= \begin{bmatrix} \nu'_L \\ (N'_R)^c \end{bmatrix} + \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ N_1 \\ N_2 \\ N_3 \end{pmatrix} \\
 \boxed{(\nu')^c &= \nu'}
 \end{aligned}$$

$$\nu'_L = V^\dagger \nu_L + S^\dagger (N_R)^c$$

$$N'_R = R^T (\nu_L)^c + U^T N_R$$

Hybrid Mass Term (2)

50

Physical mass term:

$$-\mathcal{L}'_{\text{hybrid}} = \frac{1}{2} \overline{\nu'} \begin{pmatrix} \widehat{M}_\nu & \mathbf{0} \\ \mathbf{0} & \widehat{M}_N \end{pmatrix} \nu' = \frac{1}{2} \sum_{i=1}^3 (m_i \overline{\nu}_i \nu_i + M_i \overline{N}_i N_i)$$

Kinetic term:

$$\begin{aligned} \mathcal{L}_{\text{kinetic}} &= i \overline{\nu}_L \gamma_\mu \partial^\mu \nu_L + i \overline{N}_R \gamma_\mu \partial^\mu N_R \\ &= \frac{i}{2} \begin{bmatrix} \overline{\nu}_L & \overline{(N_R)^c} \end{bmatrix} \gamma_\mu \partial^\mu \begin{bmatrix} \nu_L \\ (N_R)^c \end{bmatrix} + \frac{i}{2} \begin{bmatrix} (\nu_L)^c & \overline{N}_R \end{bmatrix} \gamma_\mu \partial^\mu \begin{bmatrix} (\nu_L)^c \\ N_R \end{bmatrix} \\ &= \frac{i}{2} \begin{bmatrix} \nu'_L & \overline{(N'_R)^c} \end{bmatrix} \gamma_\mu \partial^\mu \begin{pmatrix} V & R \\ S & U \end{pmatrix}^\dagger \begin{pmatrix} V & R \\ S & U \end{pmatrix} \begin{bmatrix} \nu'_L \\ (N'_R)^c \end{bmatrix} \\ &\quad + \frac{i}{2} \begin{bmatrix} (\nu'_L)^c & \overline{N'_R} \end{bmatrix} \gamma_\mu \partial^\mu \begin{pmatrix} V & R \\ S & U \end{pmatrix}^T \begin{pmatrix} V & R \\ S & U \end{pmatrix}^* \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} \\ &= \frac{i}{2} \begin{bmatrix} \nu'_L & \overline{(N'_R)^c} \end{bmatrix} \gamma_\mu \partial^\mu \begin{bmatrix} \nu'_L \\ (N'_R)^c \end{bmatrix} + \frac{i}{2} \begin{bmatrix} (\nu'_L)^c & \overline{N'_R} \end{bmatrix} \gamma_\mu \partial^\mu \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} \\ &= i \overline{\nu'_L} \gamma_\mu \partial^\mu \nu'_L + i \overline{N'_R} \gamma_\mu \partial^\mu N'_R \\ &= \frac{i}{2} \overline{\nu'} \gamma_\mu \partial^\mu \nu' = \frac{i}{2} \sum_{k=1}^3 (\overline{\nu}_k \gamma_\mu \partial^\mu \nu_k + \overline{N}_k \gamma_\mu \partial^\mu N_k) \end{aligned}$$

Non-unitary Flavor Mixing 51

Weak charged-current interactions of leptons:

In the flavor basis

$$\mathcal{L}_{\text{cc}} = \frac{g}{\sqrt{2}} \overline{(e \ \mu \ \tau)}_L \gamma^\mu \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}_L W_\mu^- + \text{h.c.}$$

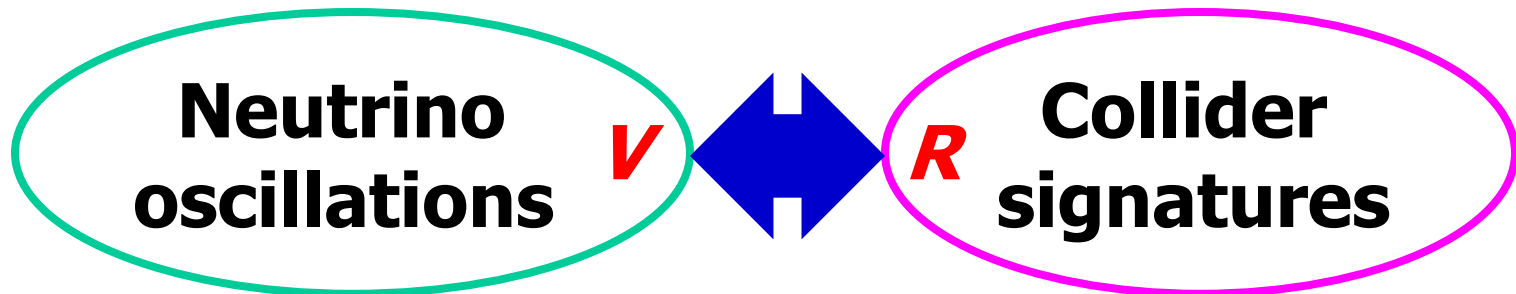
$$\nu_L = V \nu'_L + R(N'_R)^c$$

In the mass basis

$$\mathcal{L}_{\text{cc}} = \frac{g}{\sqrt{2}} \overline{(e \ \mu \ \tau)}_L \gamma^\mu \left[V \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}_L + R \begin{pmatrix} N_1 \\ N_2 \\ N_3 \end{pmatrix}_L \right] W_\mu^- + \text{h.c.}$$

V = non-unitary light neutrino mixing (**MNSP**) matrix $VV^\dagger + RR^\dagger = 1$

R = light-heavy neutrino mixing (**CC** interactions of heavy neutrinos)



TeV seesaws may bridge the gap between neutrino & collider physics.

Seesaw Mechanisms (1)

52

A **hybrid** neutrino mass Lagrangian may contain three distinct terms:

$$\begin{aligned} -\mathcal{L}'_{\text{hybrid}} &= \overline{\nu}_L M_D N_R + \frac{1}{2} \overline{\nu}_L M_L (\nu_L)^c + \frac{1}{2} \overline{(N_R)^c} M_R N_R + \text{h.c.} \\ &= \frac{1}{2} \begin{bmatrix} \overline{\nu}_L & \overline{(N_R)^c} \end{bmatrix} \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{bmatrix} (\nu_L)^c \\ N_R \end{bmatrix} + \text{h.c.} , \end{aligned}$$

- ♣ **Normal Dirac mass term**, proportional to the scale of electroweak symmetry breaking ($\sim$ **174 GeV**);
- ♣ **Light Majorana mass term**, violating the SM gauge symmetry and having a scale much lower than **174 GeV**;
- ♣ **Heavy Majorana mass term**, originating from the SU(2)_L singlet and having a scale much higher than **174 GeV**.

A strong hierarchy of **3** mass scales allows us to make approximation

$$\begin{pmatrix} V & R \\ S & U \end{pmatrix}^\dagger \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{pmatrix} V & R \\ S & U \end{pmatrix}^* = \begin{pmatrix} \widehat{M}_\nu & 0 \\ 0 & \widehat{M}_N \end{pmatrix}$$

Seesaw Mechanisms (2)

53

The above **unitary** transformation leads to the following relationships:

$$\begin{aligned} R\widehat{M}_N &= M_L R^* + M_D U^* \\ S\widehat{M}_\nu &= M_D^T V^* + M_R S^* \end{aligned}$$

$$\begin{aligned} M_R &\gg M_D \gg M_L \\ R &\sim S \sim \mathcal{O}(M_D/M_R) \end{aligned}$$

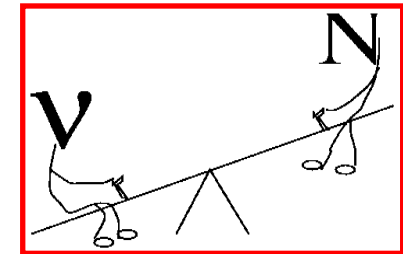
$$\begin{aligned} U\widehat{M}_N &= M_R U^* + M_D^T R^* \\ V\widehat{M}_\nu &= M_L V^* + M_D S^* \end{aligned}$$

$$\begin{aligned} U\widehat{M}_N U^T &= M_R (U U^\dagger)^T + M_D^T (R^* U^T) \approx M_R, \\ V\widehat{M}_\nu V^T &= M_L (V V^\dagger)^T + M_D (S^* V^T) \approx M_L + M_D (S^* V^T) \end{aligned}$$

$$S^* V^T = M_R^{-1} S\widehat{M}_\nu V^T - M_R^{-1} M_D^T (V V^\dagger)^T \approx -M_R^{-1} M_D^T$$

Then we arrive at the **type-(I+II)** seesaw formula:

$$M_\nu \equiv V\widehat{M}_\nu V^T \approx M_L - M_D M_R^{-1} M_D^T$$



Type-I seesaw limit: $M_\nu \approx -M_D M_R^{-1} M_D^T$ (**Fritzsch, Gell-Mann, Minkowski, 1975; Minkowski, 1977; ...**)

Type-II seesaw limit: $M_\nu = M_L$ (**Konetschny, Kummer, 1977; ...**)

History of Seesaw

54

The **seesaw** idea **originally** appeared in a paper's **footnote**.



Seesaw—A Footnote Idea:

**H. Fritzsch, M. Gell-Mann,
P. Minkowski, PLB 59 (1975) 256**

This idea was very clearly elaborated by **Minkowski** in Phys. Lett. B 67 (1977) 421 ---- but it was unjustly forgotten until **2004**.



The idea was later on embedded into the **GUT** frameworks in **1979** and **1980**:

- **T. Yanagida 1979**
- **M. Gell-Mann, P. Ramond, R. Slansky 1979**
- **S. Glashow 1979**
- **R. Mohapatra, G. Senjanovic 1980**

It was **Yanagida** who named this mechanism as "**seesaw**".

What is History?

55

History is a set of lies agreed upon

Napoleon Bonaparte

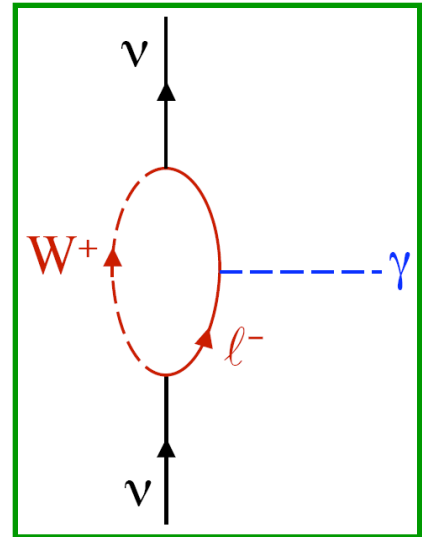


Electromagnetic Properties 56

A neutrino does not have electric charges, but it has **electromagnetic interactions** with the photon via quantum loops.

Given the SM interactions, a **massive Dirac** neutrino can only have a tiny **magnetic** dipole moment:

$$\mu_\nu \sim \frac{3eG_F}{8\sqrt{2}\pi^2} m_\nu = 3 \times 10^{-20} \frac{m_\nu}{0.1 \text{ eV}} \mu_B$$



A **massive Majorana** neutrino can **not** have **magnetic** & **electric** dipole moments, as its antiparticle is itself.

Proof: **Dirac** neutrino's electromagnetic vertex can be parametrized as

$$\Gamma_\mu(p, p') = f_Q(q^2)\gamma_\mu + f_M(q^2)i\sigma_{\mu\nu}q^\nu + f_E(q^2)\sigma_{\mu\nu}q^\nu\gamma_5 + f_A(q^2)(q^2\gamma_\mu - q_\mu q^\nu\gamma_\nu)\gamma_5$$

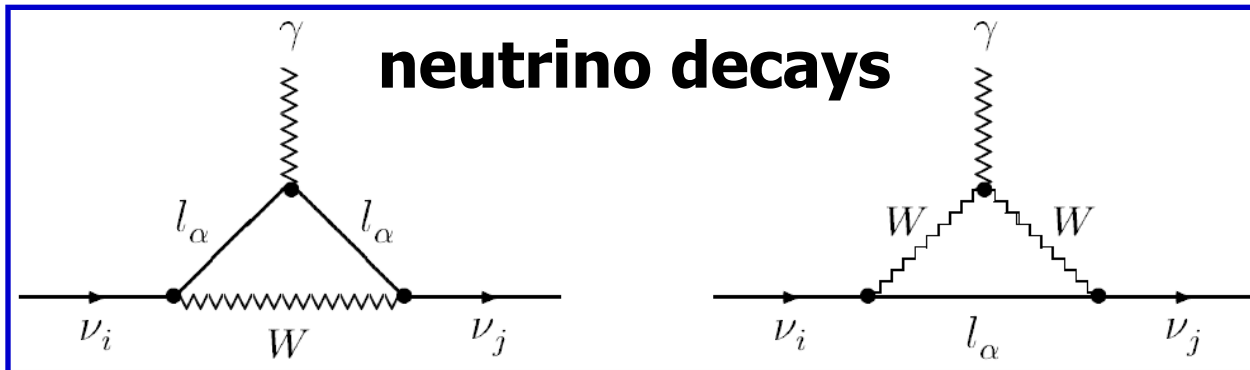
**Majorana
neutrinos**

$$\bar{\psi}\Gamma_\mu\psi = \bar{\psi}^c\Gamma_\mu\psi^c = \psi^T\mathcal{C}\Gamma_\mu\mathcal{C}\bar{\psi}^T = \left(\psi^T\mathcal{C}\Gamma_\mu\mathcal{C}\bar{\psi}^T\right)^T = -\bar{\psi}\mathcal{C}^T\Gamma_\mu^T\mathcal{C}^T\psi = \bar{\psi}\mathcal{C}\Gamma_\mu^T\mathcal{C}^{-1}\psi$$

➡ $f_Q(q^2) = f_M(q^2) = f_E(q^2) = 0$ intrinsic property of **Majorana v's**.

Transition Dipole Moments 57

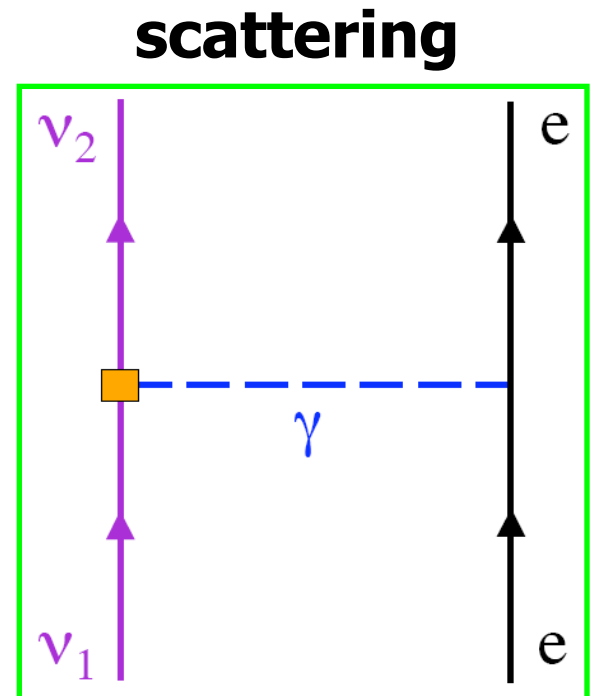
Both **Dirac** & **Majorana** neutrinos can have **transition** dipole moments (of a size comparable with μ_ν) that may give rise to neutrino decays, scattering with electrons, interactions with external magnetic field & contributions to ν masses. (**Data:** $< \text{a few} \times 10^{-11}$ Bohr magneton).



$$\mu_{\text{eff}} \equiv \sqrt{|\mu_{ij}|^2 + |\epsilon_{ij}|^2}$$

$$\Gamma_{\nu_i \rightarrow \nu_j + \gamma} = 5.3 \times \left(1 - \frac{m_j^2}{m_i^2}\right)^3 \left(\frac{m_i}{1 \text{ eV}}\right)^3 \left(\frac{\mu_{\text{eff}}}{\mu_B}\right)^2 \text{ s}^{-1}$$

$$\frac{d\sigma'_\mu}{dT} = \frac{\alpha^2 \pi}{m_e^2} \sum_{k=1}^3 \left| \sum_{j=1}^3 e^{iq_j L} V_{ej} \left(i \frac{\mu_{jk}}{\mu_B} + \frac{\epsilon_{jk}}{\mu_B} \right) \right|^2 \left(\frac{1}{T} - \frac{1}{E_\nu} \right)$$



Summary of Lecture B

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- (A) Three reasons for neutrinos to be massless in the SM.
- (B) The **Dirac** mass term and **lepton number conservation**.
- (C) The **Majorana** mass term and **lepton number violation**.
 - the **Majorana** mass matrix must be **symmetric**;
 - factor **1/2** in front of the mass term makes sense.
- (D) The **hybrid** mass term and **seesaw mechanisms**.
 - light and heavy neutrinos are **Majorana** particles;
 - the **3×3 light** flavor mixing matrix is **non-unitary**;
 - light neutrino masses: the type-(I+II) seesaw.
- (E) **Electromagnetic dipole moment** of massive neutrinos.
 - **Dirac** neutrinos have **magnetic** dipole moments;
 - **Majorana** neutrinos have no dipole moments;
 - **Dirac & Majorana** neutrinos: **transition** moments.