

Before I start

Link to Tuomas' talk:

DM candidates in Technicolor/composite models
aka
TIMPs (Technicolor Interacting Massive Particle)

- Ryttov, Sannino '08 [0809.0713]  DM from UMT
- Foadi, Frandsen, Sannino '08 [0812.3406]  First? interfering DM
- Frandsen, Sannino '09 [0911.1570]  isotriplet TIMP
- EDN, Sannino '11 [1102.3116]  One zillion effective terms

Now I can go to start

Eugenio Del Nobile

CP³ - Origins



Particle Physics & Origin of Mass

CP³ UCL Louvain-la-Neuve, 30/11/2011

Dark Matter Interference

Eugenio Del Nobile

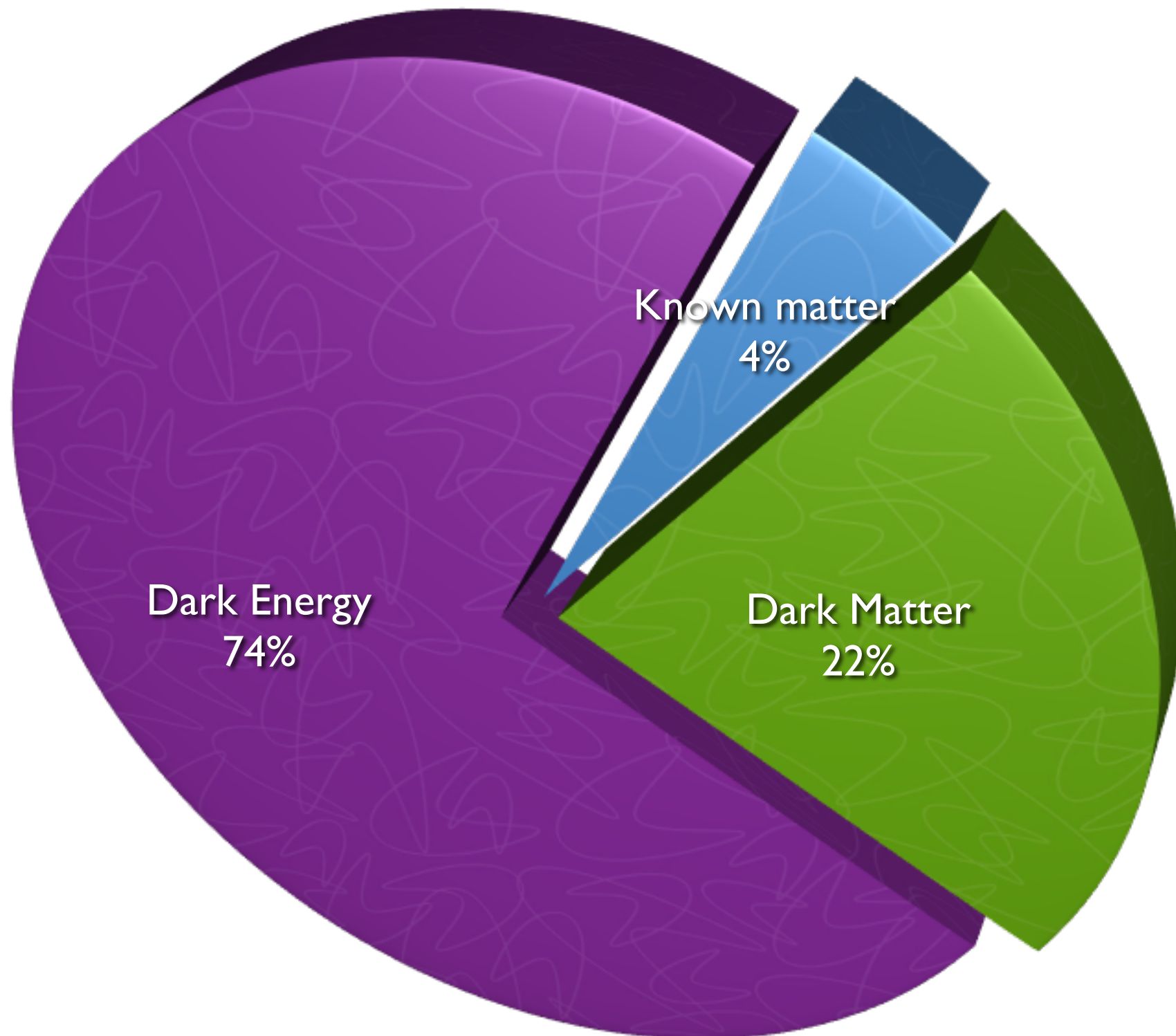
CP³ - Origins



Particle Physics & Origin of Mass

CP³ UCL Louvain-la-Neuve, 30/11/2011

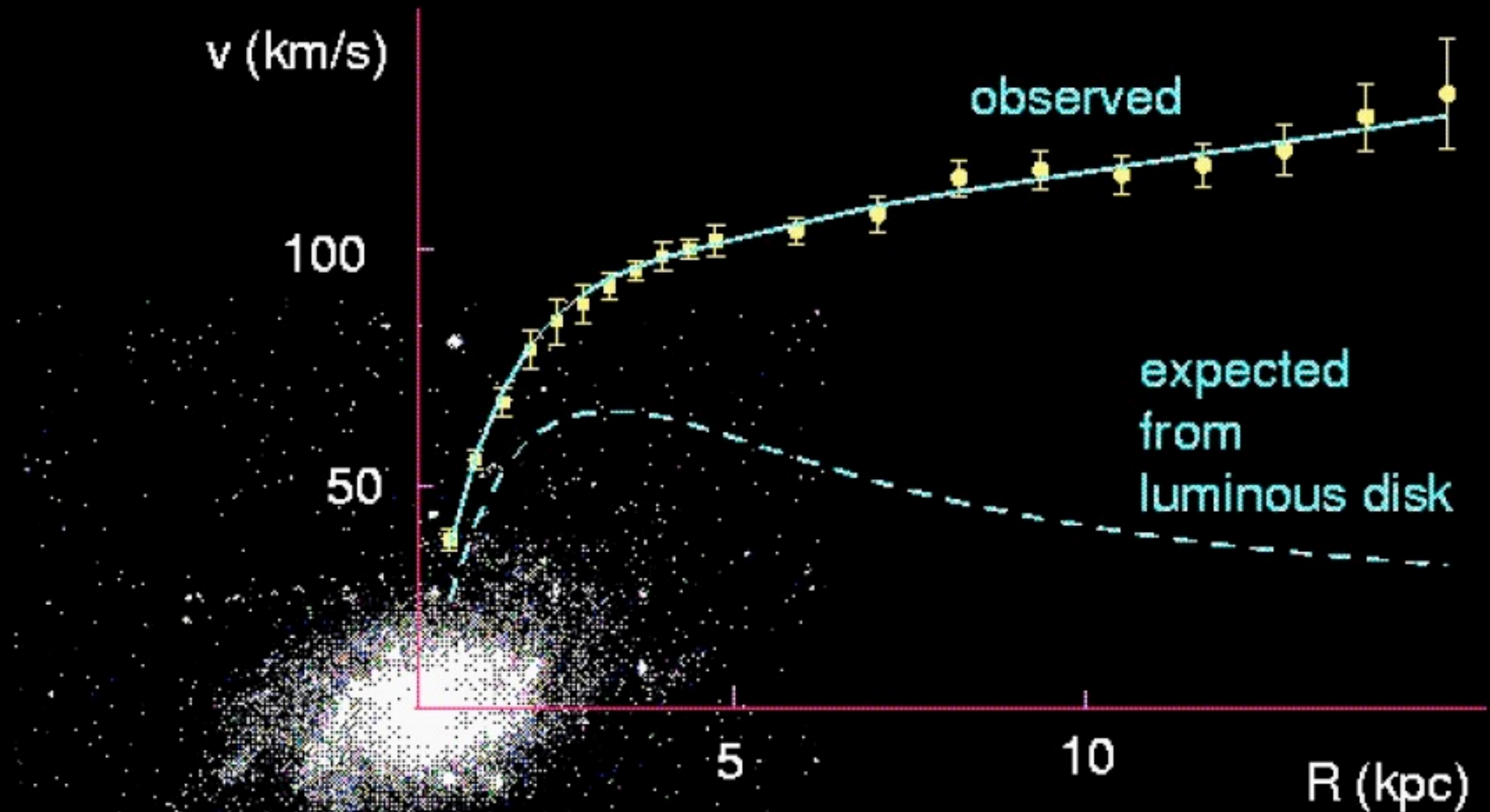
Total energy density



Astrophysical evidences for DM

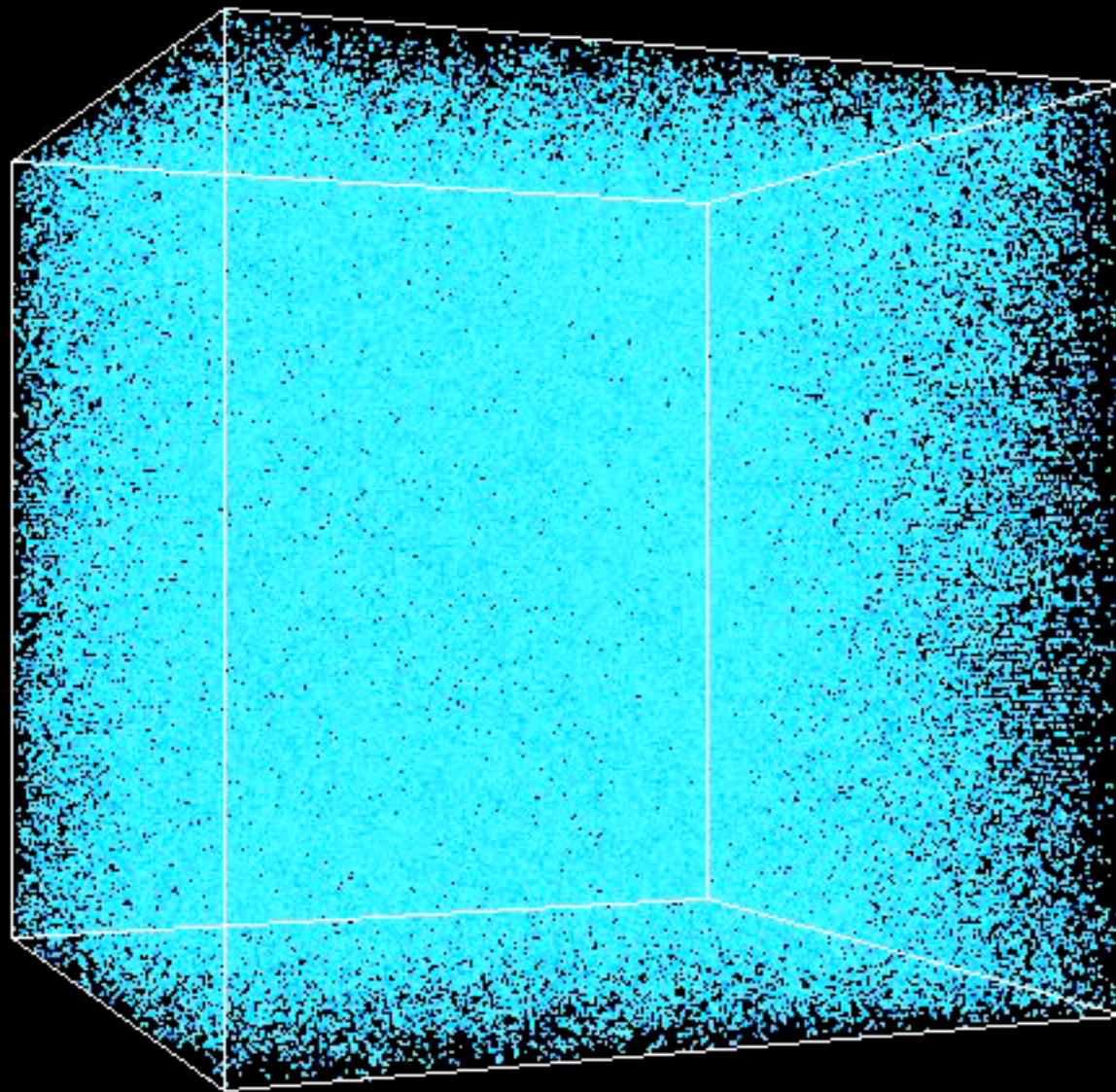
- galactic rotation curves
- gravitational lensing
- structure formation
- CMB anisotropies
- Bullet Cluster
- ...

Something missing

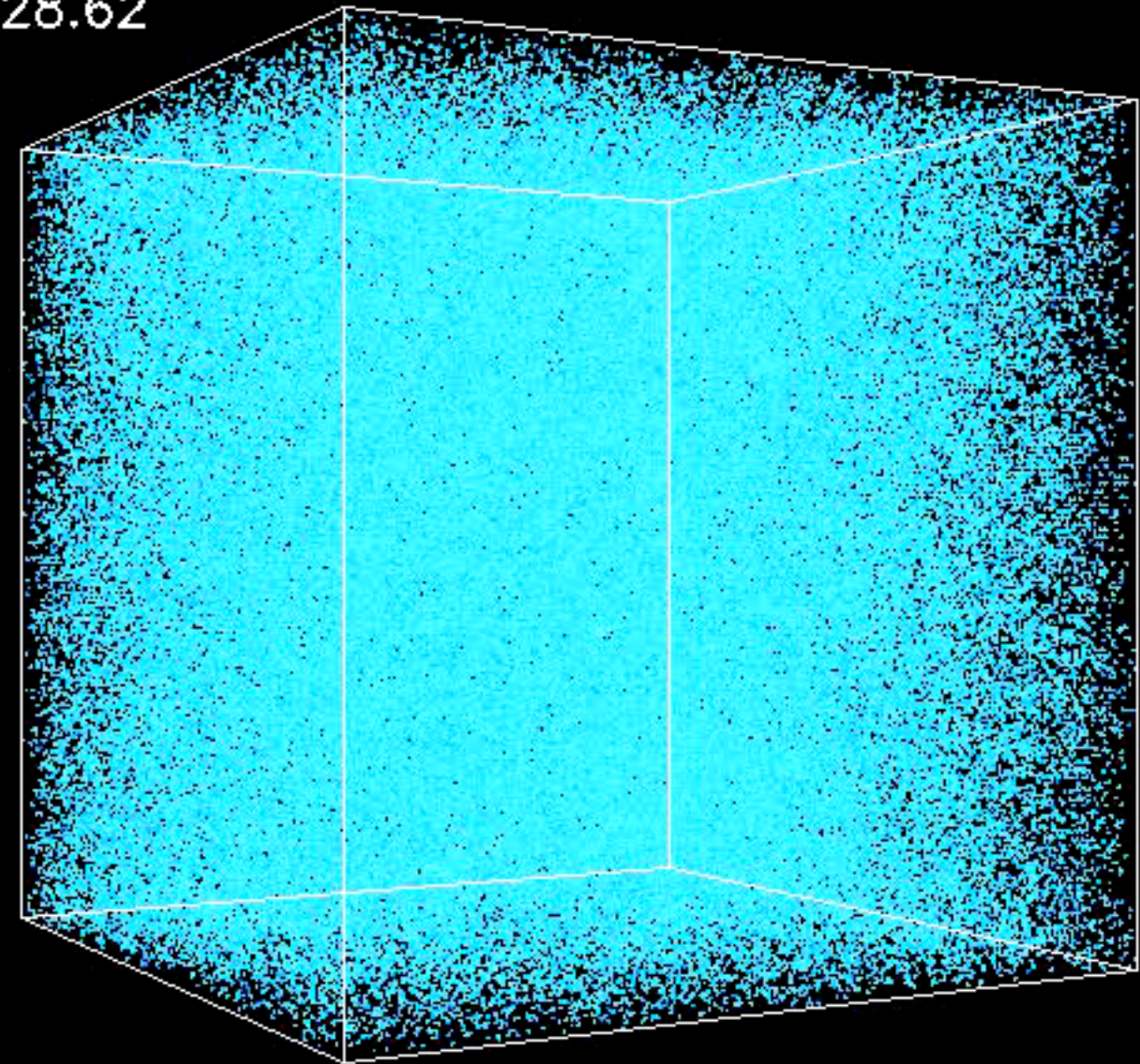


M33 rotation curve

Large structure formation



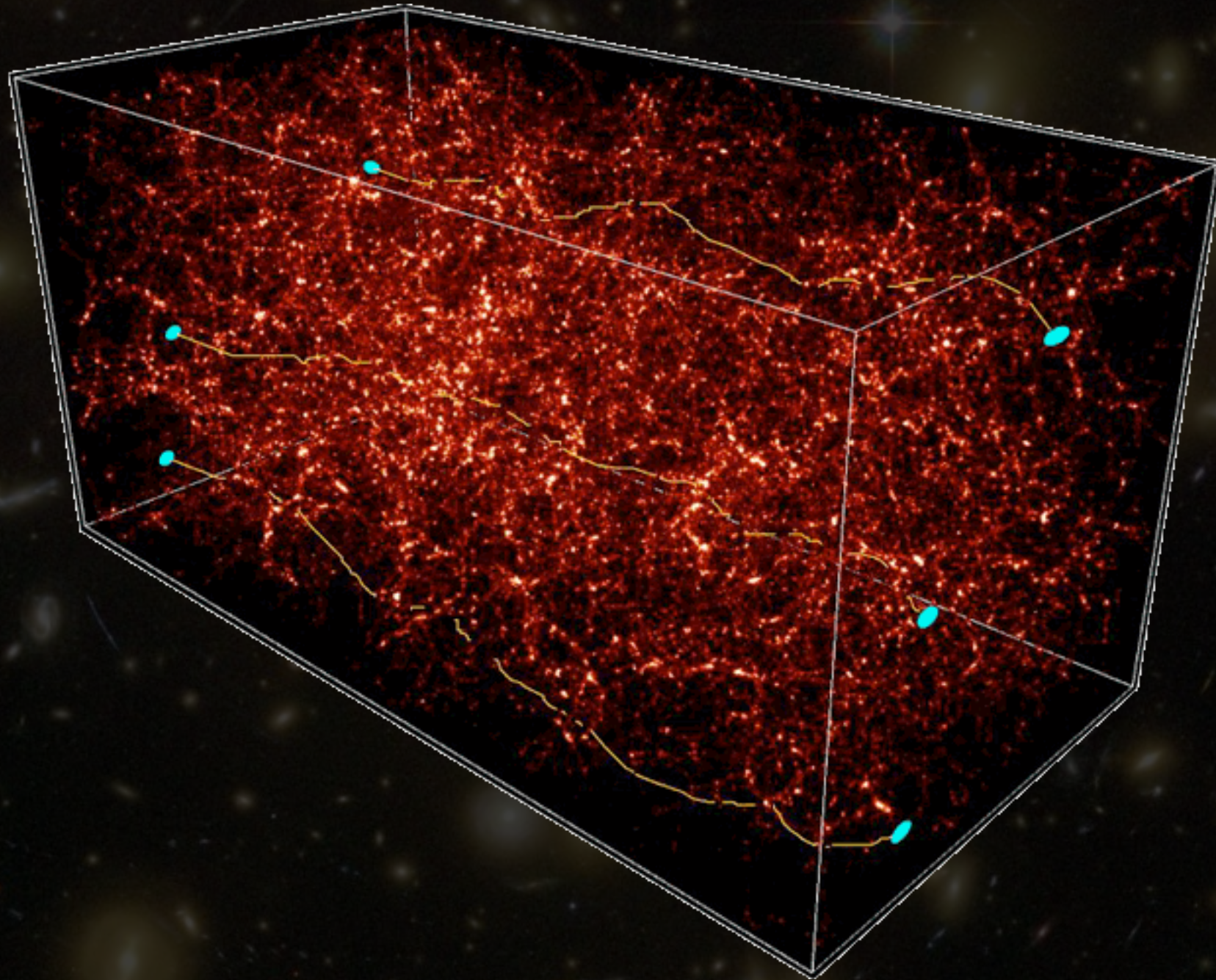
$Z=28.62$



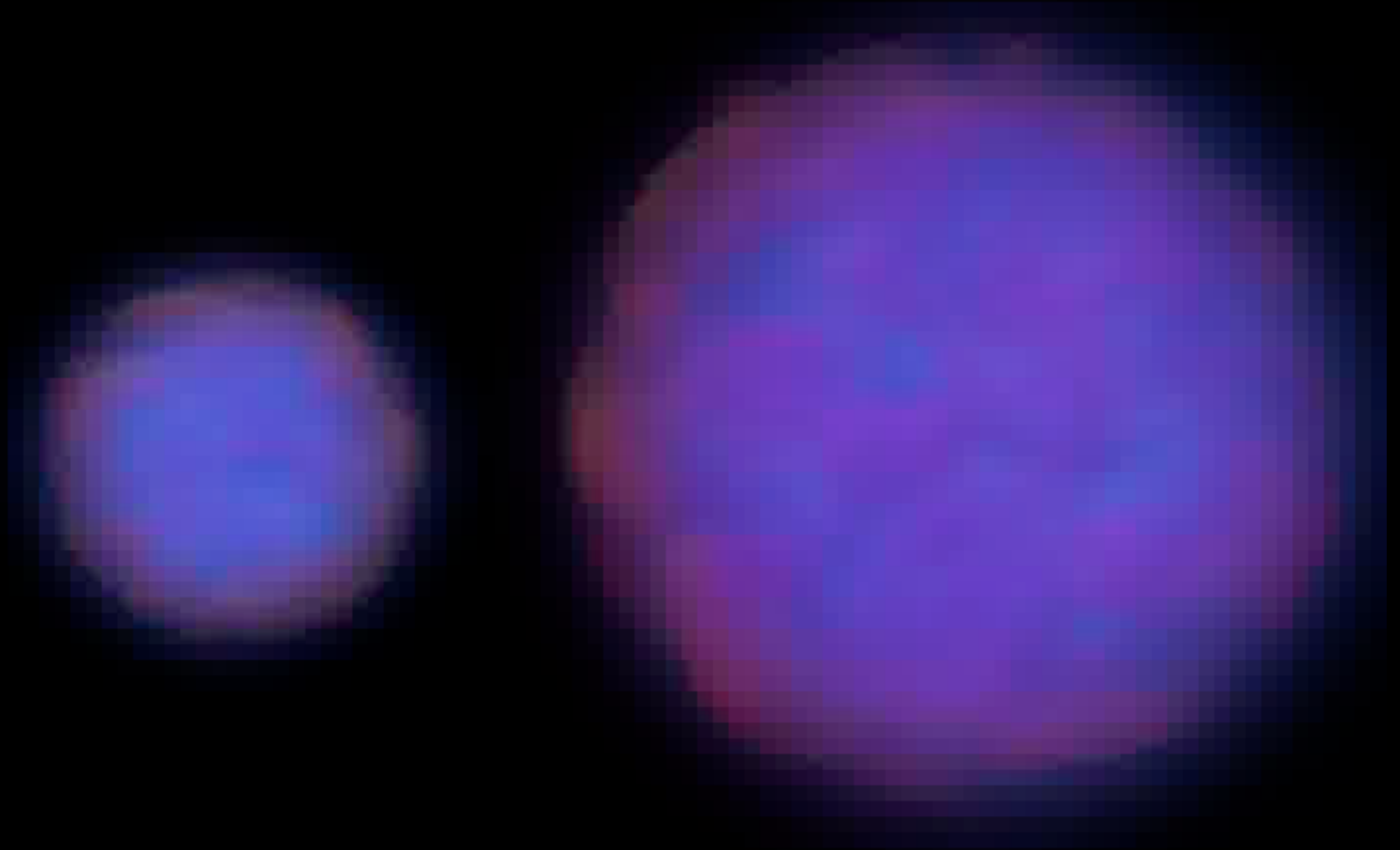
Evidences from lensing



Evidences from lensing



Bullet Cluster



Bullet Cluster



DM is not (only)

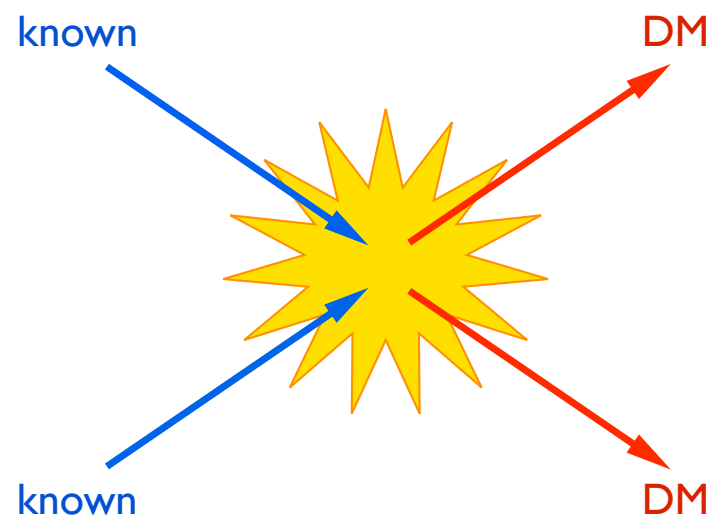
Black holes, neutron stars, brown dwarfs,
big planets that we don't see
(MAssive Compact Halo Objects: **MACHOs**)

A modification of gravity at large distances
(MOdified Newtonian Dynamics: **MOND**)

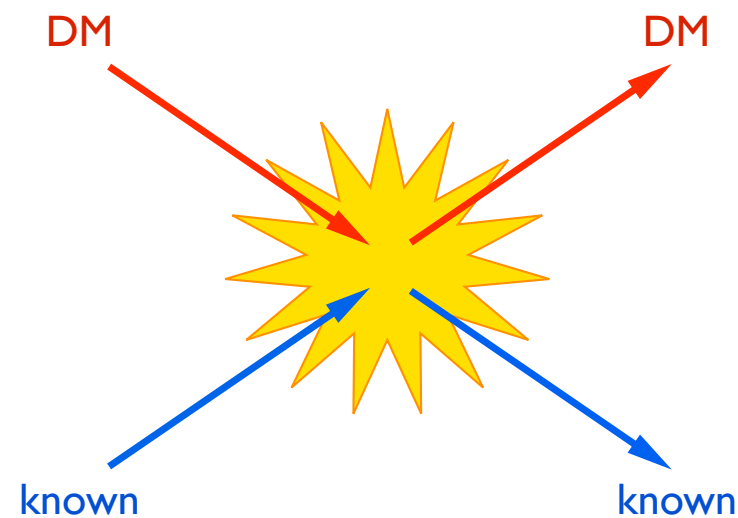
Neutrinos (**ν**)

Detection strategies

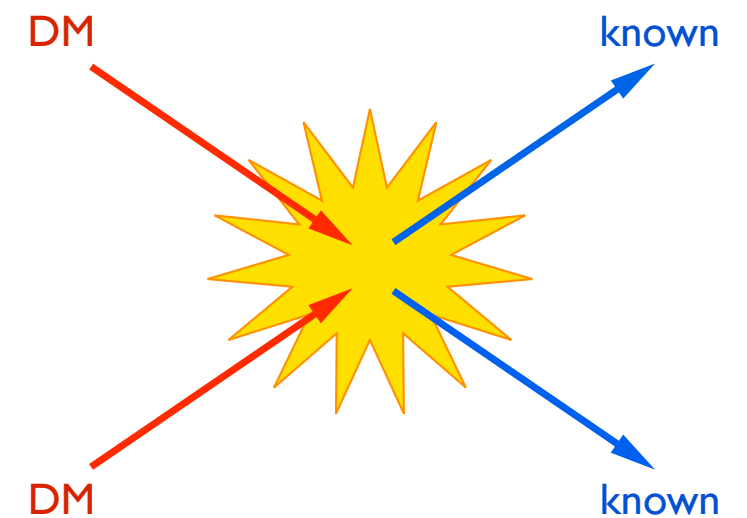
Collider



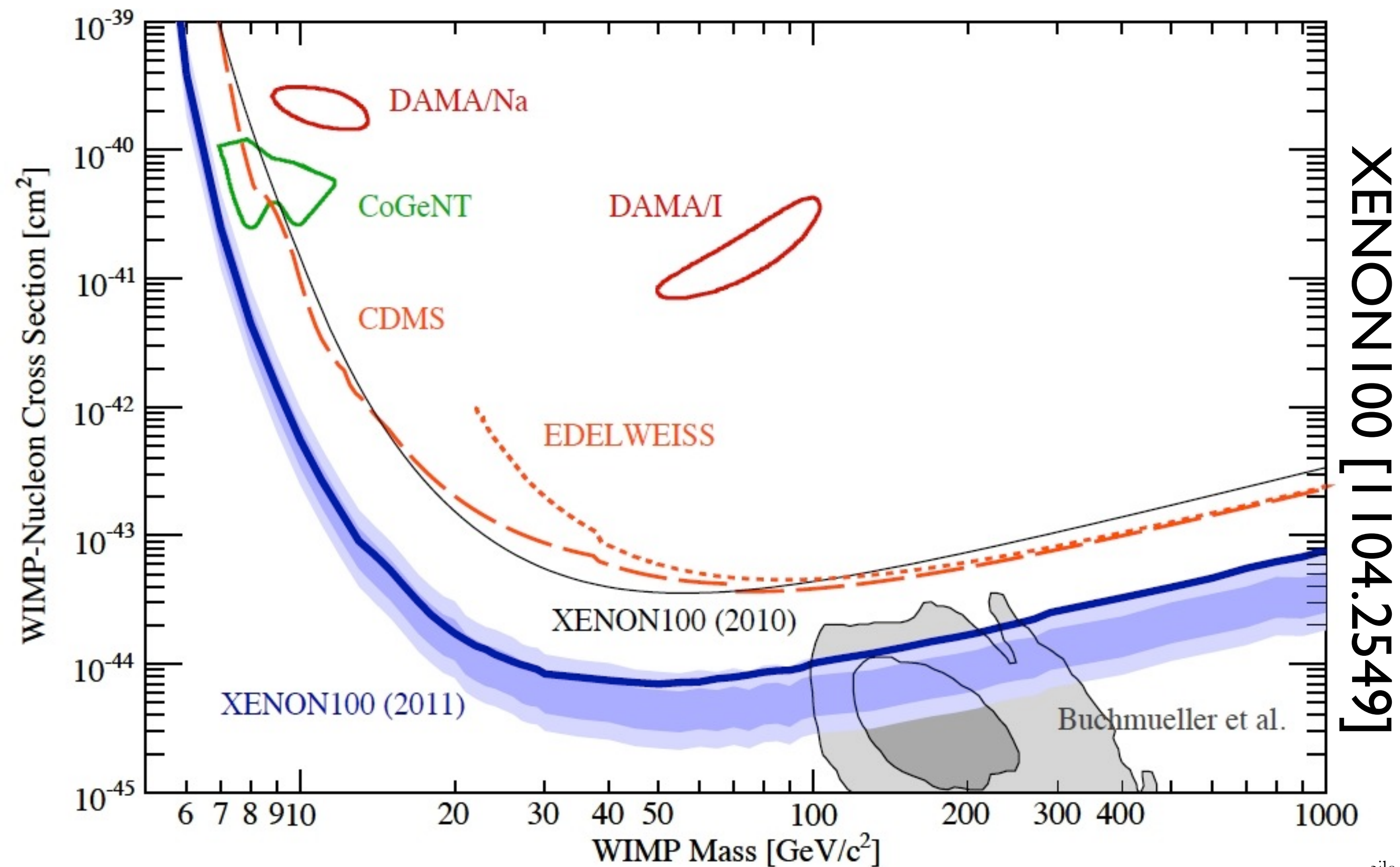
Direct



Indirect



Direct detection plot



Direct detection rate

$$R \equiv \frac{dN}{dt dm_T} = \frac{dN_T}{dm_T} n v \sigma(v)$$

Direct detection rate

$$R \equiv \frac{dN}{dt dm_T} = \frac{dN_T}{dm_T} n v \sigma(v)$$

DM density distribution $dn = n_0 f(\vec{v}, \vec{v}_E) d^3v$

Elastic kinematics: $E_R = E r (1 - \cos \theta) / 2$ $r \equiv 4 \frac{\mu_A^2}{M_{\text{DM}} m_A}$

$$R = \frac{dN_T}{dm_T} n_0 \int_0^\infty dE_R \int_{v_{\min}(E_R)}^{v_{\text{esc}}} d^3v v f(\vec{v}, \vec{v}_E) \frac{d\sigma}{dE_R}$$

$$v_{\min} = \sqrt{m_A E_R / 2 \mu_A^2}$$

Cross section

$$\frac{d\sigma}{dE_R}(E_R) \longrightarrow F^2(E_R) \frac{d\sigma_0}{dE_R} = F^2(E_R) \sigma_0 \frac{m_A}{2\mu_A^2 v^2}$$

$$\sigma_0 \equiv \int_0^{Er} \frac{d\sigma_0}{dE_R} dE_R = Er \frac{d\sigma_0}{dE_R}$$

$$\mathcal{M} = M_{\text{DM}} m_A [Z f_p + (A - Z) f_n]$$

$$\sigma_0 = \frac{\mu_A^2}{16\pi} |Z f_p + (A - Z) f_n|^2 F^2(E_R)$$

$$= \frac{\mu_A^2}{\mu_p^2} \sigma_p |Z + (A - Z) f_n / f_p|^2 F^2(E_R)$$

$$\sigma_p = \frac{\mu_p^2}{16\pi} |f_p|^2$$

Finally

Define
$$I_A \equiv \frac{dN_T}{dm_T} n_0 \int dE_R F^2(E_R) \int_{v_{\min}}^{v_{\text{esc}}} d^3v \frac{m_A}{2\mu_A^2 v} f(\vec{v}, \vec{v}_E)$$

Then
$$R = \sum_{\{A\}} \eta_A I_A \frac{\sigma_0}{F^2(E_R)} = \sum_{\{A\}} \eta_A I_A \frac{\mu_A^2}{\mu_p^2} \sigma_p |Z + (A - Z) f_n / f_p|^2$$

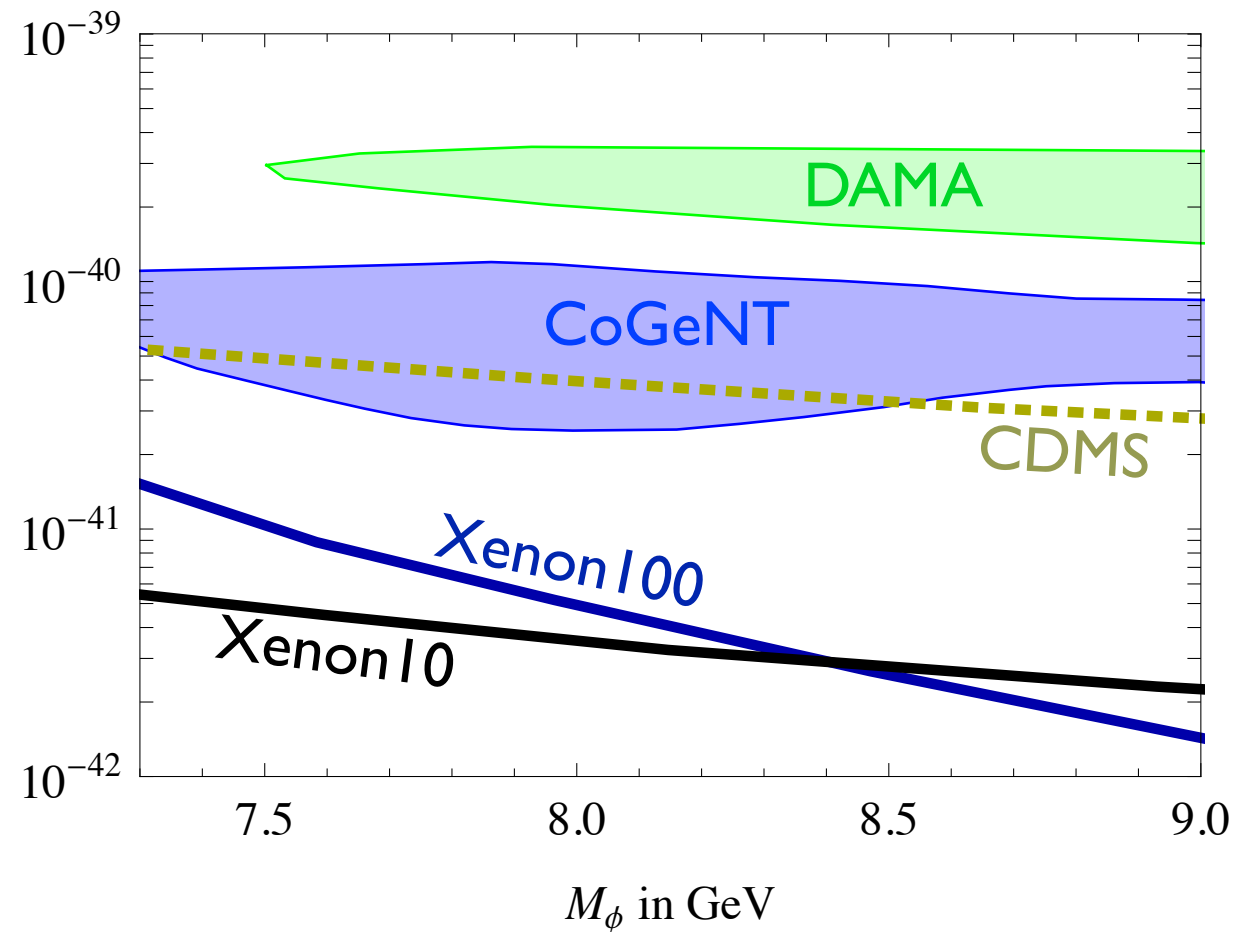
$$= \sum_{\{A\}} \eta_A I_A \frac{\mu_A^2}{\mu_p^2} \sigma_p^{\text{exp}} A^2$$

Therefore
$$\sigma_p = \sigma_p^{\text{exp}} \frac{\sum_{\{A\}} \eta_A \cancel{I_A} \frac{\mu_A^2}{\mu_p^2} A^2}{\sum_{\{A\}} \eta_A \cancel{I_A} \frac{\mu_A^2}{\mu_p^2} |Z + (A - Z) f_n / f_p|^2}$$

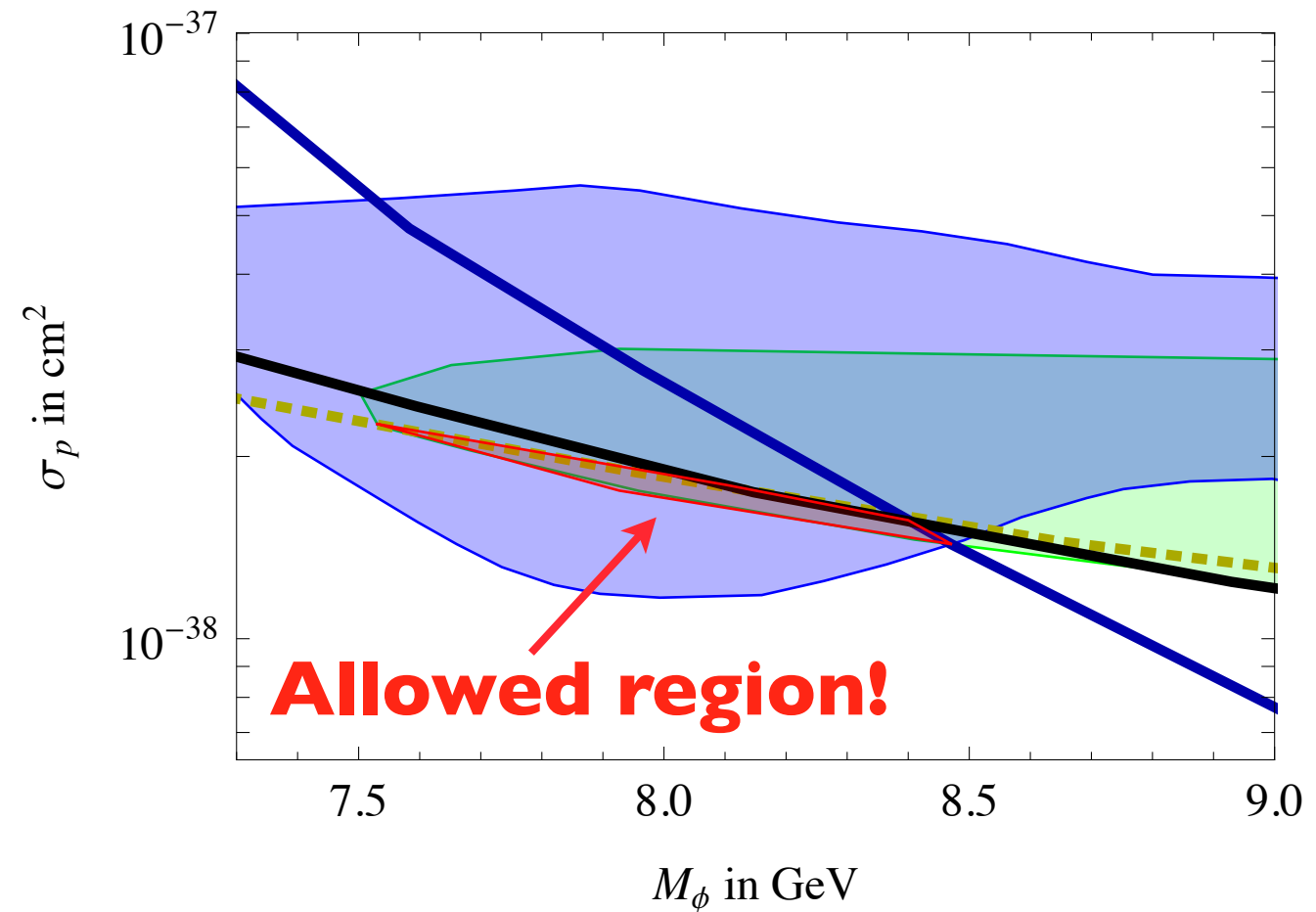
Then plot again

Feng, Kumar, Marfatia, Sanford '11 [1102.4331]

$$f_n = f_p$$



$$f_n / f_p \sim -0.7$$



$$M_{\text{DM}} \simeq 8 \text{ GeV}$$

$$\sigma_p \simeq 2 \times 10^{-38} \text{ cm}^2$$

Isospin violating DM

Z boson exchange gives $f_n / f_p \sim -10$

Higgs exchange gives $f_n / f_p \sim 1$

photon exchange gives $f_n / f_p \sim 0$

Therefore consider multiple mediators exchange

DM Interference

~ Models I

Foadi, Frandsen, Sannino '08 [0812.3406]
EDN, Kouvaris, Sannino '11 [1105.5431]

ϕ a SM-singlet scalar composite DM field

ϕ can be naturally light being a PNGB, and is protected by a new global U(1) symmetry, e.g. in UMT (Ryttov, Sannino '08 [0809.0713])

$$\mathcal{L}_{\text{int}} = -d_H H^\dagger H \phi^* \phi + i \frac{d_B}{\Lambda^2} e \phi^* \overleftrightarrow{\partial}_\mu \phi \partial_\nu F^{\mu\nu}$$

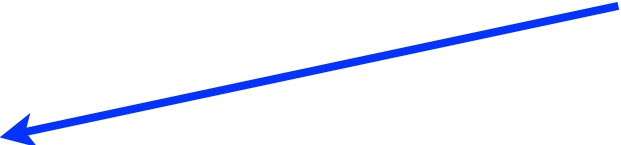
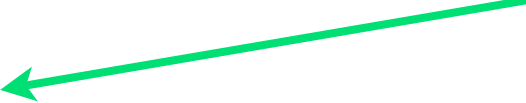
$$\frac{d_B}{\Lambda^2} \sim 3 \times 10^{-4} \text{ GeV}^{-2}$$

$$\frac{d_H}{m_H^2} \sim 5 \times 10^{-4} \text{ GeV}^{-2}$$

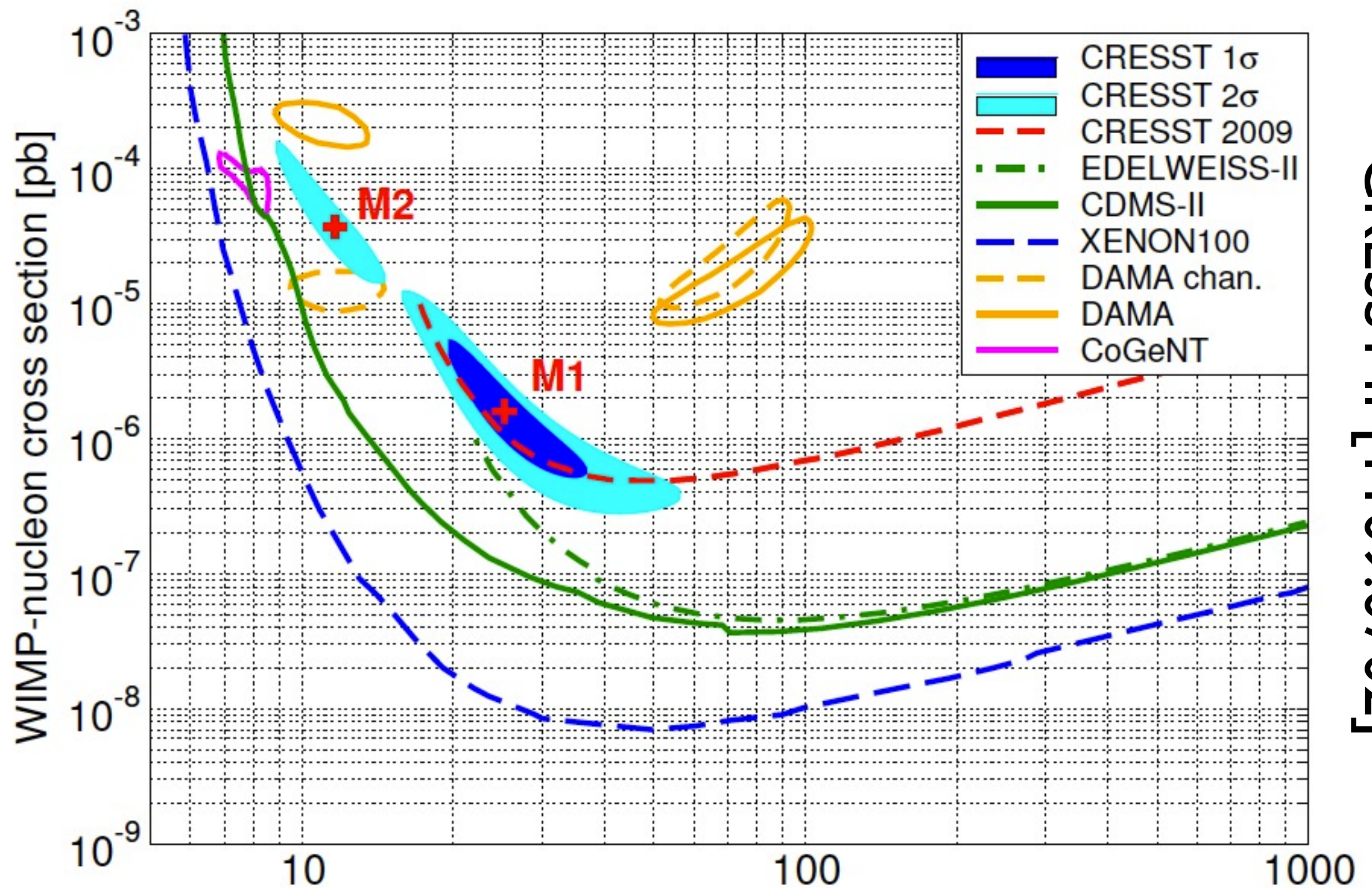
Therefore if $\Lambda, m_H \sim \mathcal{O}(100 \text{ GeV})$, d_B and d_H are natural

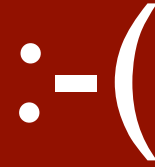
~ Models II

Other possibilities are, with fermionic DM:
(EDN, Kouvaris, Sannino, Virkajärvi '11 [1111.1902])

- Z - Higgs interference  Excluded by LEP
- Z' - Higgs interference  Works for light Higgs (~ 50 GeV)
- Z - Z' interference  Basically only Z'
- two Higgs doublets interference  It doesn't spoil Baryogenesis

Also CRESST!





CoGeNT

DAMA

Xenon10

CDMS

CRESST

$$f_n / f_p \sim -0.7$$

Xenon100

Conclusions

The experimental assumption of equal DM interaction with protons and neutrons is an assumption

DM Interference is a mechanism to get the right amount of isospin violation

Isospin violating DM is nice, but can't accommodate all the experimental results so far

Maybe what experiments see is not DM, or maybe we just need to discover some other mechanisms to find/fit it.